

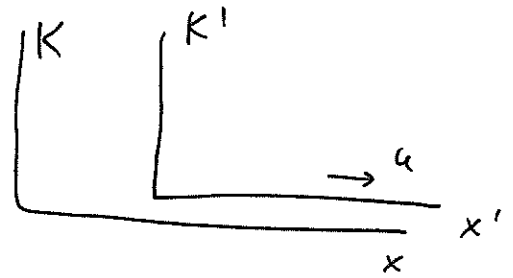
Last time

Velocity Transformations

$$V_x = \frac{V_x' + u}{1 + \frac{uV_x'}{c^2}}$$

$$V_y = \gamma \frac{V_y'}{1 + \frac{uV_x'}{c^2}}$$

$$V_z = \gamma \frac{V_z'}{1 + \frac{uV_x'}{c^2}}$$



$$\tan \theta = \frac{V' \sin \theta'}{\gamma (V' \cos \theta' + u)}$$

Relativistic Doppler Shift

plane monochromatic wave with frequency ω and wave vector \vec{k} in one frame (K) and ω' & \vec{k}' in K'

$$k^0 \equiv \omega/c, \quad k'^0 = \omega'/c, \quad \vec{\beta} = \vec{v}/c \sim \text{boost velocity}/c$$

$$\Rightarrow \begin{cases} k'^0 = \gamma (k^0 - \vec{\beta} \cdot \vec{k}) \\ k'_{\parallel} = \gamma (k_{\parallel} - \beta k^0) \\ \vec{k}'_{\perp} = \vec{k}_{\perp} \end{cases}$$

$$\Rightarrow \omega' = \gamma \omega (1 - \beta \cos \theta) \quad \text{Doppler shift}$$

$$\tan \theta' = \frac{\sin \theta}{\gamma (\cos \theta - \beta)} \quad \text{light aberration}$$

For an EM wave: $|\vec{k}| = k^0 = \frac{\omega}{c}$, $|\vec{k}'| = k'^0 = \frac{\omega'}{c}$ (114)

\Rightarrow if the angle between \vec{k} and \vec{v} is θ in K and θ' in K' $\Rightarrow \omega' = \gamma(\omega - \beta \omega \cos \theta) \Rightarrow$

$$\Rightarrow \omega' = \gamma \omega (1 - \beta \cos \theta) \quad \text{Doppler shift}$$

$$\begin{cases} \frac{\omega'}{c} \cos \theta' = \gamma \left(\frac{\omega}{c} \cos \theta - \beta \cdot \frac{\omega}{c} \right) \\ \frac{\omega'}{c} \sin \theta' = \frac{\omega}{c} \sin \theta \end{cases}$$

$$\Rightarrow \tan \theta' = \frac{\sin \theta}{\gamma (\cos \theta - \beta)}$$

(cf. with light aberration.)

Four - vectors.

We have seen one example: $x^0 = ct$, $x^1 = x$, $x^2 = y$, $x^3 = z$

$$\begin{pmatrix} x_0' \\ x_1' \\ x_2' \\ x_3' \end{pmatrix} = \begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x_0 \\ x_1 \\ x_2 \\ x_3 \end{pmatrix}$$

A general Lorentz transformation can be

- a boost along x , y , or z axis (\forall angle θ)
- a rotation around the x , y , or z axis ($\forall \beta$)

Definition A 4-vector A^M is a set of 4

quantities (A^0, A^1, A^2, A^3) , which under Lorentz

transformation transform as $A'^M = \frac{\partial x'^M}{\partial x^N} A^N$ (just like x^M)

Example boost along the x-axis:

$$\begin{pmatrix} A^{0'} \\ A^{1'} \\ A^{2'} \\ A^{3'} \end{pmatrix} = \begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} A^0 \\ A^1 \\ A^2 \\ A^3 \end{pmatrix}$$

$\mu, \nu = 0, 1, 2, 3$
summation over ν is implied.

$\Rightarrow A^M, M=0,1,2,3$ is a contravariant

vector if it transforms according to:

$$A'^M = \frac{\partial x'^M}{\partial x^N} A^N \quad (\text{equivalent to above})$$

$\Rightarrow B_M, M=0, \dots, 3$ is a covariant vector

if
$$B'_M = \frac{\partial x^N}{\partial x'^M} B_N$$

Example: $\frac{\partial \varphi}{\partial x^M}$ is a covariant vector as

$$\frac{\partial \varphi}{\partial x'^M} = \frac{\partial x^N}{\partial x'^M} \frac{\partial \varphi}{\partial x^N}$$

One can define tensors by

$$A'^M B'^N = \frac{\partial x'^M}{\partial x^\alpha} \frac{\partial x'^N}{\partial x^\beta} A^\alpha B^\beta \Rightarrow \text{rank two contravariant}$$

tensor would be $C'^{MN} = \frac{\partial x'^M}{\partial x^\alpha} \frac{\partial x'^N}{\partial x^\beta} C^{\alpha\beta}$, etc.

In general $T_{\mu_1 \dots \mu_n}^{\nu_1 \dots \nu_m}$ transforms as $A_{\mu_1} \dots A_{\mu_n} B^{\nu_1} \dots B^{\nu_m}$.

Definition Scalar (inner) product of 2 vectors (1/6)

is defined by $A_\mu \cdot B^\mu$ (summation assumed)

Let's prove that it's Lorentz-invariant:

$$A'_\mu \cdot B'^\mu = \frac{\partial x^\alpha}{\partial x'^\mu} A_\alpha \frac{\partial x'^\mu}{\partial x^\beta} B^\beta = \frac{\partial x^\alpha}{\partial x^\beta} A_\alpha B^\beta = \delta^\alpha_\beta \cdot$$

$$A_\alpha B^\beta = A_\alpha B^\alpha \quad \text{Q.E.D.}$$

The interval is a scalar: (it's Lorentz-invariant)

$$ds^2 = (dx^0)^2 - (dx^1)^2 - (dx^2)^2 - (dx^3)^2$$

Define the metric tensor by

$$ds^2 \equiv g_{\mu\nu} dx^\mu dx^\nu$$

$$g_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$$

(Minkowski
space
Cartesian
coordinates)

Note that $dx_\mu dx^\mu$ is also a Lorentz-scalar.

Identifying $dx_\mu = g_{\mu\nu} dx^\nu$ we see that

$g_{\mu\nu}$ lowers indices of 4-vectors, tensors, etc.

Example $x^\mu = (ct, \vec{x})$, $x_\mu = g_{\mu\nu} x^\nu = (ct, -\vec{x})$.

(Def.) $g^{\mu\nu}$ is an inverse matrix for $g_{\mu\nu}$: $g^{\mu\nu} g_{\nu\sigma} = \delta^{\mu}_{\sigma}$ (117)

$A_{\mu} = g_{\mu\nu} A^{\nu} \Rightarrow A^{\mu} = g^{\mu\nu} A_{\nu}$ where $S^{\mu}_{\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$

(e.g. $x_{\mu} = g_{\mu\nu} x^{\nu}, \dots$)
 $\Rightarrow g^{\mu\nu}$ raises indices of 4-vectors, tensors, etc.
 $S^{\mu}_{\nu} = g^{\mu\alpha} \cdot g_{\alpha\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \Rightarrow g^{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} = g_{\mu\nu}$ (only for Minkowski space!)

indeed, as $A_{\mu} B^{\mu} = S^{\mu}_{\nu} \cdot A_{\mu} B^{\nu} = g^{\mu\nu} A_{\mu} B^{\nu}$

Define an abbreviated notation: $\partial_{\mu} \equiv \frac{\partial}{\partial x^{\mu}}$

$\partial^{\mu} \equiv \frac{\partial}{\partial x_{\mu}}$

$\Rightarrow \partial_{\mu} \varphi$ is a covariant vector

$\partial^{\mu} \varphi$ is a contravariant vector (check!)

$\partial_{\mu} A^{\mu}$ is Lorentz-invariant $\left(\frac{\partial \varphi}{\partial x^{\mu}} = \frac{\partial x^{\nu}}{\partial x^{\mu}} \frac{\partial \varphi}{\partial x^{\nu}} \right)$

4.1 Laplace operator $\frac{\partial^2}{c^2 dt^2} - \vec{\nabla}^2 = \partial_{\mu} \partial^{\mu}$ is (d'Alembertian)

also Lorentz-invariant. (can you prove this?)

4-velocity

Let's define a 4-vector for velocity:

$dx^{\mu} = (dx^0, dx^1, dx^2, dx^3) \Rightarrow v^{\mu} \stackrel{?}{=} \frac{dx^{\mu}}{dt}$

But: time is not a scalar!

$\frac{dx^\mu}{dt} \sim \frac{dx^\mu}{dx^0} \sim$ not a Lorentz-vector.

\Rightarrow try proper time $d\tau = \frac{ds}{c} \Rightarrow u^\mu \equiv \frac{dx^\mu}{d\tau}$ 4-velocity

as $d\tau = \frac{dt}{\gamma} \Rightarrow u^0 = \frac{cdt}{dt/\gamma} = c\gamma$

$\vec{u} = \frac{d\vec{x}}{dt} \cdot \gamma = \gamma \cdot \vec{v} \Rightarrow u^\mu = \gamma(c, \vec{v})$

Note $u_\mu u^\mu = c^2$.

Boost in terms of rapidity.

~~$$\begin{pmatrix} A^0 \\ A^1 \\ A^2 \\ A^3 \end{pmatrix} = \begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \cdot \begin{pmatrix} A^0 \\ A^1 \\ A^2 \\ A^3 \end{pmatrix}$$~~

~~\Rightarrow Define rapidity η~~

~~by $\beta \equiv \tanh \eta = \frac{e^\eta - e^{-\eta}}{e^\eta + e^{-\eta}}$~~

~~Define light-cone coordinates~~

$A^+ = \frac{A^0 + A^1}{\sqrt{2}}$

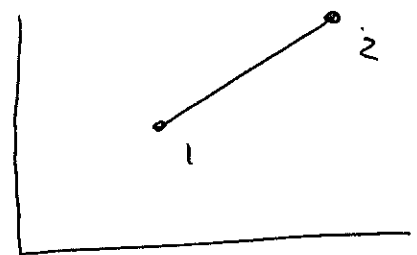
$A^- = \frac{A^0 - A^1}{\sqrt{2}}$

~~\Rightarrow then~~

~~$A_\mu A^\mu = 2A^+A^- - (A^2)^2 - (A^3)^2$~~

Relativistic Mechanics.

Consider a free particle (moving along a straight line). We need to construct a Lorentz-invariant action for such particle. It's characterized



by a 4-vector $X^M \Rightarrow$ the only Lorentz-invariant is the interval \Rightarrow

$\Rightarrow \int_1^2 ds$ (can't have $\int_1^2 (ds)^2 \sim$ still infinitesimal)

\Rightarrow write action as $S = -A \cdot \int_1^2 ds$

As $ds^2 = c^2 dt^2 - (dx^i)^2 = dt^2 (1 - \beta^2(t)) \Rightarrow$

$\Rightarrow ds = c dt \sqrt{1 - \beta^2(t)} \Rightarrow S = -Ac \int_{t_1}^{t_2} dt \sqrt{1 - \beta^2(t)}$

\Rightarrow as $S = \int_{t_1}^{t_2} dt \cdot L$, where L is the

Lagrangian.