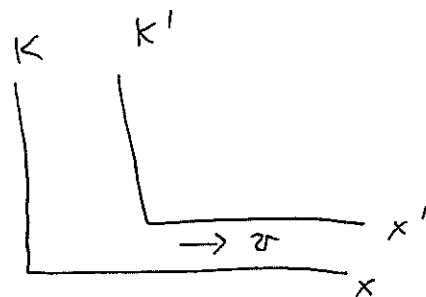


Classical Field Theory

4-vectors, notations

defining $x^0 = ct, x^1 = x, x^2 = y, x^3 = z$

write Lorentz transformation



$$\text{as } \begin{pmatrix} x'^0 \\ x'^1 \\ x'^2 \\ x'^3 \end{pmatrix} = \begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} x^0 \\ x^1 \\ x^2 \\ x^3 \end{pmatrix}$$

$$\beta = v/c$$

$$\gamma = \frac{1}{\sqrt{1-\beta^2}}$$

$$x'^M = \Lambda^M_{\nu} x^{\nu}$$

Definition a 4-vector $A^M = \begin{pmatrix} A^0 \\ A^1 \\ A^2 \\ A^3 \end{pmatrix}$ is an object

which under Lorentz transformation transforms

$$\text{as } \begin{pmatrix} A'^0 \\ A'^1 \\ A'^2 \\ A'^3 \end{pmatrix} = \underbrace{\begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}}_{\Lambda^M_{\nu}} \begin{pmatrix} A^0 \\ A^1 \\ A^2 \\ A^3 \end{pmatrix}$$

(example: x^M is a contravariant 4-vector)

$$A'^M = \Lambda^M_{\nu} A^{\nu}$$

$\Rightarrow A^M$ is a contravariant vector: $A'^M = \frac{\partial x'^M}{\partial x^{\nu}} A^{\nu}$

B_{μ} is a covariant vector: $B'_{\mu} = \frac{\partial x^{\nu}}{\partial x'^{\mu}} B_{\nu}$

$\Rightarrow \frac{\partial \varphi}{\partial x^{\mu}} \equiv \partial_{\mu} \varphi$ with φ , scalar field is a covariant vector

$$\text{as } \frac{\partial \varphi}{\partial x'^{\mu}} = \frac{\partial x^{\nu}}{\partial x'^{\mu}} \frac{\partial \varphi}{\partial x^{\nu}}$$

Tensors: $A^{\mu} B^{\nu}$ ~ contravariant, $A_{\mu} B_{\nu}$ ~ covariant
(rank 2), can have higher ranks.

Def. Scalar (inner) product of 2 vectors is $A_{\mu} B^{\mu}$.
(assume summation).

It is Lorentz-invariant: $A'_{\mu} B'^{\mu} = \frac{\partial x^{\alpha}}{\partial x'^{\mu}} A_{\alpha} \frac{\partial x'^{\mu}}{\partial x^{\beta}} B^{\beta}$
 $= \frac{\partial x^{\alpha}}{\partial x^{\beta}} A_{\alpha} B^{\beta} = \delta^{\alpha}_{\beta} A_{\alpha} B^{\beta} = A_{\alpha} B^{\alpha}$.

Def. The interval $ds^2 (= dx_{\mu} dx^{\mu}) = (dx^0)^2 - (dx^1)^2 - (dx^2)^2 - (dx^3)^2$

It's a Lorentz-invariant too.

Def. The metric tensor $g_{\mu\nu}$ is defined by

$$ds^2 = g_{\mu\nu} dx^{\mu} dx^{\nu}$$

In our Minkowski space $g_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \equiv \eta_{\mu\nu}$
(throughout the course we'll use this notation)

$dx_{\mu} dx^{\mu}$ ~ also a Lorentz-scalar $\Rightarrow dx_{\mu} = g_{\mu\nu} dx^{\nu}$

$\Rightarrow g_{\mu\nu}$ lowers & raises indices!

Example: $x^M = (ct, \vec{x}) \Rightarrow x_\mu = g_{\mu\nu} x^\nu = (ct, -\vec{x})$ (+)
 contravariant covariant

In general $A_\mu = g_{\mu\nu} A^\nu$, $A^\mu = g^{\mu\nu} A_\nu$

where $g^{\mu\nu}$ is defined by requiring that

$g^{\mu\alpha} g_{\alpha\nu} = \delta^\mu_\nu$: if that is true \Rightarrow start

with $A_\mu = g_{\mu\nu} A^\nu \Rightarrow g^{\alpha\mu} A_\mu = g^{\alpha\mu} g_{\mu\nu} A^\nu =$
 $= \delta^\alpha_\nu A^\nu = A^\alpha \Rightarrow A^\alpha = g^{\alpha\mu} A_\mu.$

$g^{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \equiv \eta^{\mu\nu} + 00. \quad (= g_{\mu\nu})$

Def. $\partial_\mu \equiv \frac{\partial}{\partial x^\mu}$, $\partial^\mu \equiv \frac{\partial}{\partial x_\mu} \Rightarrow \partial_\mu \varphi$ is a covariant vector,

$\partial^\mu \varphi$ is a contravariant vector. (check!)

$\partial_\mu A^\mu$ is a Lorentz-invariant.

$\partial_\mu \partial^\mu = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \vec{\nabla}^2$ is also Lorentz-invariant.

Examples: other important 4-vectors are

$p^\mu = \begin{pmatrix} E/c \\ \vec{p} \end{pmatrix}$, $p_\mu = \begin{pmatrix} E/c \\ -\vec{p} \end{pmatrix} \Rightarrow p_\mu p^\mu = \left(\frac{E}{c}\right)^2 - \vec{p}^2 = m^2 c^2.$

$A^\mu = (\Phi, \vec{A})$ in E&M, Φ ~ electric potential,
 \vec{A} - vector potential.

$J^\mu = (c\rho, \vec{J})$ with ρ the charge density, \vec{J} the
current density.

Notations: from now on $c = 1$ and $\hbar = 1$

"natural units".

\Rightarrow mass, momentum, energy are measured in
the same units (eV, keV, MeV, GeV, ...)

$$1 \text{ eV} = 1.6 \times 10^{-19} \text{ J.}$$

distances, time are measured in femto-
-meters aka fermis (fm):

$$1 \text{ fm} \approx 5 \text{ GeV}^{-1}$$

$$1 \text{ GeV} = 10^9 \text{ eV}, \quad 1 \text{ femto-} = 10^{-15} \text{ m.} = 1 \text{ fm.}$$

-meter

proton's mass $m_p = 0.938 \text{ GeV} \approx 1 \text{ GeV}$

electron's mass $m_e = 0.511 \text{ MeV} = 0.5 \times 10^{-3} \text{ GeV}$

Classical Scalar Field Theory (9) (real field)

$\varphi(x^\mu) = \varphi(x^0, \vec{x}) \sim$ a function of space-time points x^μ .

(example \sim temperature field $T(t, \vec{x})$)

in Classical Mechanics one has point particles $i=1, \dots, N$ with the Lagrangian $L(q_i, \dot{q}_i, t)$

and the action $S = \int dt L(q_i, \dot{q}_i, t)$

$q_i \sim$ degrees of freedom (e.g. particle coordinates)

$\dot{q}_i = \frac{dq_i}{dt} \sim$ generalized velocities

Now, instead of discrete point particles we

have a field $\varphi(\vec{x}, t) \Rightarrow$

Classical Mechanics

Classical Field Theory

$q_i \rightarrow \varphi(x^0, \vec{x})$

$i \rightarrow \vec{x}, t$

$\dot{q}_i \rightarrow \partial_\mu \varphi, \mu=0, 1, 2, 3$

$L(q_i, \dot{q}_i, t) \rightarrow \int d^3x \mathcal{L}(\varphi, \partial_\mu \varphi)$

\mathcal{L} is Lagrangian density. (usually called the Lagrangian)

The action is $\mathcal{S} = \int dt \mathcal{L} = \underbrace{\int dt d^3x}_{d^4x} \mathcal{L}(\varphi, \partial_\mu \varphi)$ (remember $c=1$)

\mathcal{S} is a Lorentz-scalar (better be, physics is Lorentz-invariant)

What about $d^4x = dx^0 dx^1 dx^2 dx^3$? Remember that

$x'^M = \Lambda^M_\nu x^\nu$ with $\Lambda^M_\nu = \frac{\partial x'^M}{\partial x^\nu}$ a matrix of L. tr.

$\Rightarrow d^4x' = \underbrace{|\det \Lambda|}_{\text{Jacobian}} d^4x$

Now, $\det \Lambda = \det \begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} = \gamma^2(1-\beta^2) = 1$ (true in general)

$\Rightarrow d^4x' = d^4x \Rightarrow d^4x$ is a Lorentz-scalar

$\Rightarrow \mathcal{L}$ is a Lorentz-scalar!

Just like in classical mechanics, in classical field theory dynamics is given by the least action

principle; field φ is determined by requiring that \mathcal{S} is stationary with respect to small perturbation around φ : $\mathcal{S}[\varphi + \delta\varphi] = \mathcal{S}[\varphi] + o(\delta\varphi^2)$.