

Last time

# Canonical Quantization of Free

Vector Field (cont'd)

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu}, \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

Lorenz gauge  $\partial_\mu A^\mu = 0$ .

if  $A_\mu$  are 4 fields  $\Rightarrow \pi^\mu = F^{\mu 0} \Rightarrow \pi^0 = 0$

$\Rightarrow$  problem

$$\text{Use } \mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{\lambda}{2} (\partial_\mu A^\mu)^2$$

$$\pi^\mu = \frac{\delta \mathcal{L}}{\delta (\partial_0 A_\mu)} = F^{\mu 0} - \lambda g^{\mu 0} (\partial_\rho A^\rho)$$

$$\Rightarrow \pi^0 = -\lambda \partial_\rho A^\rho \neq 0$$

$$[A_\mu(\vec{x}, t), \pi_\nu(\vec{y}, t)] = i g_{\mu\nu} \delta^3(\vec{x} - \vec{y})$$

$$[A_\mu(\vec{x}, t), A_\nu(\vec{y}, t)] = 0 = [\pi_\mu(\vec{x}, t), \pi_\nu(\vec{y}, t)]$$

$$A_\mu(x) = \int \frac{d^3k}{(2\pi)^3 2E_k} \sum_{\lambda=0}^3 \epsilon_\mu^\lambda(k) \left[ \hat{a}_{\vec{k}, \lambda}^- e^{-ik \cdot x} + \hat{a}_{\vec{k}, \lambda}^{+\dagger} e^{ik \cdot x} \right]$$

real polarization vector  $\epsilon^\lambda, \epsilon^{\lambda'} = g^{\lambda\lambda'}$

$$[\hat{a}_{\vec{k}, \lambda}, \hat{a}_{\vec{k}', \lambda'}^\dagger] = -g_{\lambda\lambda'} (2\pi)^3 2E_{\vec{k}} \delta^3(\vec{k} - \vec{k}')$$

$$\partial_\nu \left[ \frac{\delta \mathcal{L}}{\delta (\partial_\nu A_\mu)} \right] = 0 \Rightarrow \partial_\nu F^{\mu\nu} - \lambda \partial^\mu \partial_\rho A^\rho = 0 \quad (72)$$

$$\Rightarrow \partial_\nu \partial^\mu A^\nu - \square A^\mu - \lambda \partial^\mu \partial_\rho A^\rho = 0$$

$$\square A^\mu - (1-\lambda) \partial^\mu \partial_\nu A^\nu = 0$$

$\square \partial_\mu A^\mu = 0 \Rightarrow$  NOT SO in QM  
 $\Rightarrow$  classically can set boundary conditions to have  $\partial_\mu A^\mu = 0$  and  $\square A^\mu = 0$ .

$\lambda = 1 \Rightarrow$  "Feynman gauge"  $\Rightarrow$  get  $\square A^\mu = 0$  back.

$\Rightarrow$  the same classical physics with a different

Lagrangian:  $\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{2} (\partial_\mu A^\mu)^2$ .

Now,  $\pi^0 = \frac{\delta \mathcal{L}}{\delta \dot{A}_0} = -\partial_\mu A^\mu$ . No longer need to

impose  $\partial_\mu A^\mu = 0$  explicitly  $\Rightarrow \pi^0$  does not have to be 0.

Quantize the system:

$$[A_\mu(\vec{x}, t), \pi_\nu(\vec{y}, t)] = i g_{\mu\nu} \delta(\vec{x} - \vec{y})$$

$$[A_\mu(\vec{x}, t), A_\nu(\vec{y}, t)] = [\pi_\mu(\vec{x}, t), \pi_\nu(\vec{y}, t)] = 0$$

$$A_\mu(\vec{x}, t) = \int \frac{d^3 k}{(2\pi)^3 2\epsilon_k} \sum_{\lambda=0}^3 \epsilon_\mu^\lambda(\vec{k}) \left[ \hat{a}_{\vec{k}, \lambda} e^{-ik \cdot x} + \hat{a}_{\vec{k}, \lambda}^\dagger e^{ik \cdot x} \right]$$

choose the basis of polarization vectors in such a way that  $\epsilon^\lambda \cdot \epsilon^{\lambda'} = g^{\lambda\lambda'}$

One then gets:

(73)

$$\left[ \hat{a}_{\vec{k}, \lambda}^{\dagger}, \hat{a}_{\vec{k}', \lambda'}^{\dagger} \right] = -g_{\lambda \lambda'} (2\pi)^3 2\varepsilon_k \delta(\vec{k} - \vec{k}')$$

all other commutators are zero.

if  $k^M = (k, 0, 0, k) \Rightarrow$

$$\varepsilon^0 = \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix}, \quad \varepsilon^1 = \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix}, \quad \varepsilon^2 = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}, \quad \varepsilon^3 = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix}$$

$\varepsilon^{1,2}, k = 0$  (transverse)

$\varepsilon^3 \sim$  longitudinal,  $\varepsilon^0 \sim$  time-like.

Fock space:  $\hat{a}_{\vec{k}, \lambda}^{\dagger} |0\rangle, \dots$

$$|1, \lambda\rangle = \int \frac{d^3k}{(2\pi)^3 2\varepsilon_k} f(\vec{k}) \hat{a}_{\vec{k}, \lambda}^{\dagger} |0\rangle \quad \sim \text{one-photon state}$$

Problem:  $\langle 1, \lambda=0 | 1, \lambda=0 \rangle = \int \frac{d^3k}{(2\pi)^3 2\varepsilon_k} \frac{d^3k'}{(2\pi)^3 2\varepsilon_{k'}} f(\vec{k}) f^*(\vec{k}')$

$$\langle 0 | \hat{a}_{\vec{k}', \lambda=0}^{\dagger} \hat{a}_{\vec{k}, 0}^{\dagger} |0\rangle = - \int \frac{d^3k}{(2\pi)^3 2\varepsilon_k} |f(\vec{k})|^2 \langle 0|0\rangle$$

$$\left[ \hat{a}_{\vec{k}', 0}^{\dagger}, \hat{a}_{\vec{k}, 0}^{\dagger} \right] = - (2\pi)^3 2\varepsilon_k \delta(\vec{k} - \vec{k}')$$

$\Rightarrow$  negative norm states!

$$H = \int \frac{d^3k}{(2\pi)^3 2\varepsilon_k} \varepsilon_k \left[ \sum_{\lambda=1}^3 \hat{a}_{\vec{k}, \lambda}^{\dagger} \hat{a}_{\vec{k}, \lambda} - \hat{a}_{\vec{k}, 0}^{\dagger} \hat{a}_{\vec{k}, 0} \right] \sim \text{negative energies problem ...}$$

=> way out: demand that while  $\partial_\mu A^\mu \neq 0$  at the operator level, for physical states one

has  $\partial_\mu A^{(+)\mu} |\psi\rangle = 0$

where (+) denotes positive-energy part of  $A_\mu$  (the term with  $\hat{a}_{\vec{k},\lambda}^+$ ).

$\langle \psi | \partial_\mu A^\mu | \psi \rangle = \langle \psi | \partial_\mu A^{(+)\mu} + \partial_\mu A^{(-)\mu} | \psi \rangle = 0$ .

(Gupta & Bleuler method).  $\langle \psi | \partial_\mu A^{(-)\mu} = [\partial_\mu A^{(+)\mu} | \psi \rangle]^+ \Rightarrow$

$\Rightarrow \partial_\mu A^{(+)\mu} | \psi \rangle = 0 \Rightarrow \sum_{\lambda=0}^3 k^\mu \epsilon_\mu^\lambda(\vec{k}) \hat{a}_{\vec{k},\lambda}^+ | \psi \rangle = 0$

$\Rightarrow (k^\mu \epsilon_\mu^0 \hat{a}_{\vec{k},0}^+ + k^\mu \epsilon_\mu^3 \hat{a}_{\vec{k},3}^+) | \psi \rangle = 0$

as  $\epsilon^{1,2} \cdot k = 0$ . Now  $k \cdot \epsilon^0 = k^0 = -k \cdot \epsilon^3$

$\Rightarrow (\hat{a}_{\vec{k},0}^+ - \hat{a}_{\vec{k},3}^+) | \psi \rangle = 0$

=> physical states are mixtures of longitudinal & time-like photons (along with the transverse photons)

$\Rightarrow \langle \psi | \hat{a}_{\vec{k},0}^+ \hat{a}_{\vec{k},0}^+ | \psi \rangle = \langle \psi | \hat{a}_{\vec{k},3}^+ \hat{a}_{\vec{k},3}^+ | \psi \rangle$

=> only transverse photons contribute in H:

$\langle \psi | H | \psi \rangle = \int \frac{d^3k}{(2\pi)^3 2\epsilon_k} \epsilon_k \cdot \langle \psi | \sum_{\lambda=1}^2 \hat{a}_{\vec{k},\lambda}^+ \hat{a}_{\vec{k},\lambda}^+ | \psi \rangle \geq 0$

$\Rightarrow$  no negative energy problem any more.

(75)

## Massive Vector Field

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{m^2}{2} A_\mu A^\mu$$

$m$  is the mass

$$\partial_\mu F^{\mu\nu} + m^2 A^\nu = 0$$

Proca  
equ's.

$\partial_\mu A^\mu = 0$  always  $\Rightarrow$  only  $A^i$  are independent d.o.f.  $\Rightarrow$

$$A_\mu(x) = \int \frac{d^3k}{(2\pi)^3 2\epsilon_k} \sum_{\lambda=1}^3 \epsilon_\mu^{(\lambda)}(\vec{k}) \left[ \hat{a}_{\vec{k},\lambda}^- e^{-ik \cdot x} + \hat{a}_{\vec{k},\lambda}^{1+} e^{ik \cdot x} \right]$$

with  $[\hat{a}_{\vec{k},\lambda}^-, \hat{a}_{\vec{k}',\lambda'}^{1+}] = \delta_{\lambda\lambda'} (2\pi)^3 2\epsilon_k \delta(\vec{k}-\vec{k}')$

$k \cdot \epsilon^{(\lambda)} = 0$  for  $\lambda = 1, 2, 3 \Rightarrow$  in rest frame

have  $k^\mu = (m, 0, 0, 0) \Rightarrow$

$$\epsilon_\mu^{(1)} = (0, 1, 0, 0), \quad \epsilon_\mu^{(2)} = (0, 0, 1, 0), \quad \epsilon_\mu^{(3)} = (0, 0, 0, 1).$$

3 degrees of freedom (all physical)

$$H = \int \frac{d^3k}{(2\pi)^3 2\epsilon_k} \epsilon_k \sum_{\lambda=1}^3 \hat{a}_{\vec{k},\lambda}^{1+} \hat{a}_{\vec{k},\lambda}^-$$

Time evolution:  $-i \partial_t A_\mu = [H, A_\mu]$  for both massive & massless vector fields.

(Note that  $A^0$  is not a d.o.f. as it is eliminated using  $\partial_\mu A^\mu = 0$ . No worries about  $\pi^0 = 0$  again.)