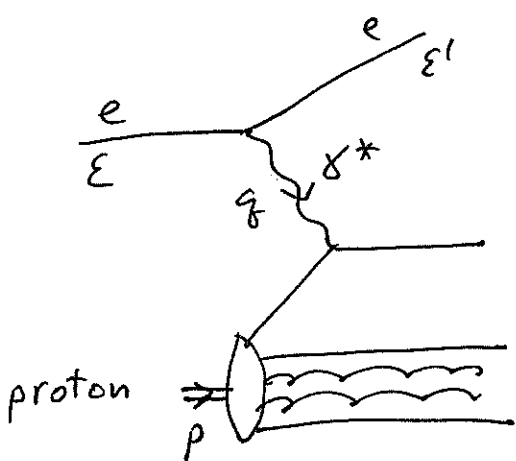


Last time | Parton Model and DIS

DIS (cont'd)



$$e + p \rightarrow e + X$$

$$Q^2 = -q^2$$

photon virtuality

$$x = \frac{Q^2}{2p \cdot q}$$

Bjorken x

$$\frac{d\sigma}{d^3k'} = \frac{dE_M^2}{Q^4 \epsilon \epsilon'} l_{\mu\nu} W^{\mu\nu}$$

← rest frame of the proton

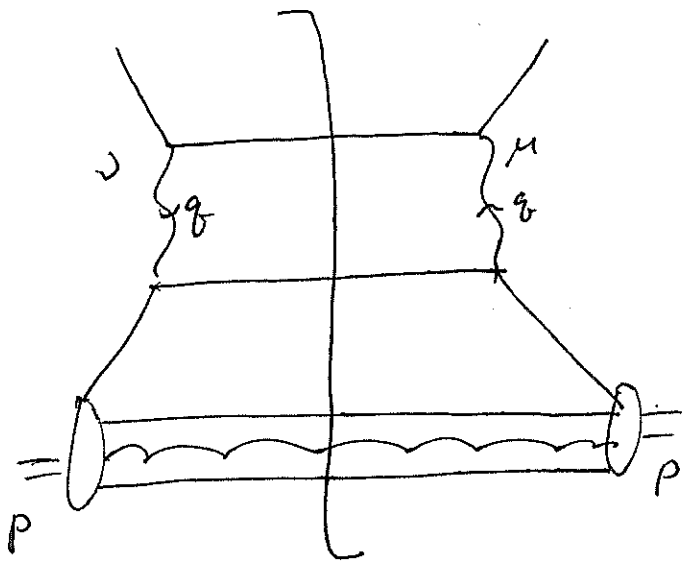
$$l_{\mu\nu} = 2(k_\mu k'_\nu + k'_\mu k_\nu - g_{\mu\nu} k \cdot k') \quad \text{Leptonic tensor}$$

(put $m_e = 0$)

$$W_{\mu\nu} = \frac{1}{4\pi m_p} \int d^4x e^{iq \cdot x} \langle p | j_\mu(x) j_\nu(0) | p \rangle$$

hadronic tensor

$$j_\mu(x) = \sum_f e_f \bar{q}_f(x) \gamma_\mu q_f(x) \sim \text{EM current}$$

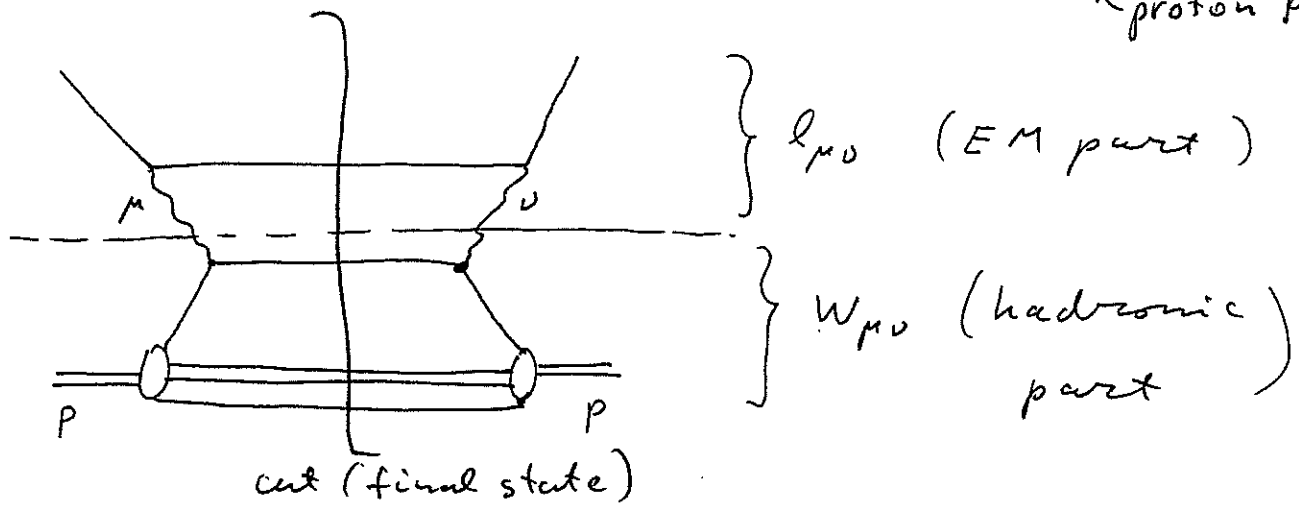


} $e^{\mu\nu}$

} $w_{\mu\nu}$

$$W_{\mu\nu} = \frac{1}{4\pi m_p} \int d^4x e^{iq \cdot x} \langle p | j_\mu(x) j_\nu(0) | p \rangle$$

(averaged over σ)
 \nwarrow proton polarization



$$W_{\mu\nu}(p, q) = A(x, Q^2) p_\mu p_\nu + B(x, Q^2) g_\mu g_\nu + C(x, Q^2) g_{\mu\nu} + D(x, Q^2) (p_\mu g_\nu + p_\nu g_\mu) + E(x, Q^2) (p_\mu g_\nu - p_\nu g_\mu) + F(x, Q^2) \epsilon_{\mu\nu\alpha\beta} p^\alpha q^\beta$$

$F = 0$ in $\gamma^* p, \gamma^* A$ (F comes from γ_5 's, appears in ν DIS).

(i) $q_\mu W^{\mu\nu} = 0$ (current conservation)
 $q_\mu \rightarrow \partial_\mu \Rightarrow \partial_\mu j^\mu(x) = 0$

$$A p_\nu (p \cdot q) + B g_\nu \cdot q^2 + C g_\nu + D (p \cdot q g_\nu + q^2 p_\nu) + E (p \cdot q g_\nu - q^2 p_\nu) = 0$$

(ii) $q_\nu W^{\mu\nu} = 0 = A p_\mu (p \cdot q) + B q^2 g_\mu + D (p \cdot q g_\mu + q^2 p_\mu) + E (p_\mu q^2 - p \cdot q g_\mu) = 0$



$$(1) - (2) = 0 \Rightarrow (E = 0.)$$

(84)

p_μ and q_μ are independent \Rightarrow

$$0 = A p \cdot q + D q^2$$

$$0 = B q^2 + C + D p \cdot q$$

$$D = -A \frac{p \cdot q}{q^2}$$

$$B = -\frac{1}{q^2} C + A \left(\frac{p \cdot q}{q^2} \right)^2$$

$$W_{\mu\nu} = A \left[p_\mu p_\nu - \frac{p \cdot q}{q^2} (p_\mu q_\nu + p_\nu q_\mu) + \left(\frac{p \cdot q}{q^2} \right)^2 q_\mu q_\nu \right] + C \left[g_{\mu\nu} - \frac{q_\mu q_\nu}{q^2} \right]$$

Usually one writes

$$W_{\mu\nu} = -W_1(x, Q^2) \left[g_{\mu\nu} - \frac{q_\mu q_\nu}{q^2} \right] + \frac{W_2(x, Q^2)}{m_p^2} \cdot \left[p_\mu p_\nu - \frac{p \cdot q}{q^2} (p_\mu q_\nu + p_\nu q_\mu) + \left(\frac{p \cdot q}{q^2} \right)^2 q_\mu q_\nu \right]$$

W_1 & W_2 are structure functions (Def.)

Using $q_\mu l^{\mu\nu} = q_\nu l^{\mu\nu} = 0$ yields

$$l_{\mu\nu} W^{\mu\nu} = -W_1 (-4 k \cdot k') + \frac{2W_2}{m_p^2} [2 p_0 k \cdot p_0 k' - m^2 k \cdot k']$$

$$2 \epsilon \epsilon' \sin^2 \frac{\theta}{2} \qquad 2m^2 \epsilon \epsilon' - 2m^2 \epsilon \epsilon' \sin^2 \frac{\theta}{2} =$$

$$= 2m^2 \epsilon \epsilon' \cos^2 \frac{\theta}{2}$$

$$g_{\mu\nu} l^{\mu\nu} = (k-k')_{\mu} 2(k^{\mu} k'^{\nu} + k^{\nu} k'^{\mu} - g^{\mu\nu} k \cdot k') =$$

$$= 2(k^2 k'^{\nu} + k^{\nu} k'^2 - k^{\nu} k \cdot k' - k \cdot k k'^{\nu} - k^{\nu} k'^2 + k^{\nu} k k')$$

$$= 2(k^2 k'^{\nu} - k'^2 k^{\nu}) \approx 0 \text{ as } k^2 \approx k'^2 \approx 0$$

(neglect electron's mass), $g_{\nu\lambda} l^{\mu\nu} = 0$ (similar)

$$\Rightarrow l_{\mu\nu} W^{\mu\nu} = l^{\mu\nu} \left[-W_1 \left(g_{\mu\nu} - \frac{g_{\mu}^{\alpha} g_{\nu}^{\beta}}{g^2} \right) + \frac{W_2}{m_p^2} \right]$$

$$\left(p_{\mu} p_{\nu} - \frac{p \cdot g}{g^2} \left(p_{\mu} g_{\nu}^{\alpha} + p_{\nu} g_{\mu}^{\alpha} \right) + \left(\frac{p \cdot g}{g^2} \right)^2 g_{\mu}^{\alpha} g_{\nu}^{\beta} \right)$$

$$= -l^{\mu}_{\mu} W_1 + \frac{W_2}{m_p^2} p_{\mu} p_{\nu} l^{\mu\nu} = \left[\text{as } l^{\mu\nu} = 2(k^{\mu} k'^{\nu} + k^{\nu} k'^{\mu} - g^{\mu\nu} k \cdot k') \right]$$

$$= -2(2k \cdot k' - 4k \cdot k') W_1 + \frac{W_2}{m_p^2} 2(2p \cdot k p \cdot k' - p^2 k \cdot k')$$

$$= 4k \cdot k' W_1 + 2 \frac{W_2}{m_p^2} (2p \cdot k p \cdot k' - m_p^2 k \cdot k')$$

remember: $k^{\mu} = (\epsilon, 0, 0, \epsilon)$, $k'^{\mu} = (\epsilon', \epsilon' \sin \theta, 0, \epsilon' \cos \theta)$
 $p^{\mu} = (m_p, \vec{0})$

$$\Rightarrow k \cdot k' = 2\epsilon\epsilon' \sin^2(\theta/2); \quad p \cdot k = m_p \epsilon, \quad p \cdot k' = m_p \epsilon'$$

$$L_{\mu\nu} W^{\mu\nu} = 4 \epsilon \epsilon' \left[2 W_1 \sin^2 \frac{\theta}{2} + W_2 \cos^2 \frac{\theta}{2} \right]$$

$$\frac{d\sigma}{d^3k'} = \frac{4 d\epsilon d\epsilon'}{Q^4} \left[2 W_1 \sin^2 \frac{\theta}{2} + W_2 \cos^2 \frac{\theta}{2} \right]$$

By varying the angle θ can separate W_1 & W_2 contributions in experiments. \Rightarrow Rosenbluth separation

Usually one defines $F_1(x, Q^2) = W_1(x, Q^2)$, $F_2(x, Q^2) = \nu W_2(x, Q^2)$

The Parton Model.

Sterman ch 14, Peskin 17.5
Y.K. & Levin, ch. 2.2

Go to Infinite Momentum Frame:

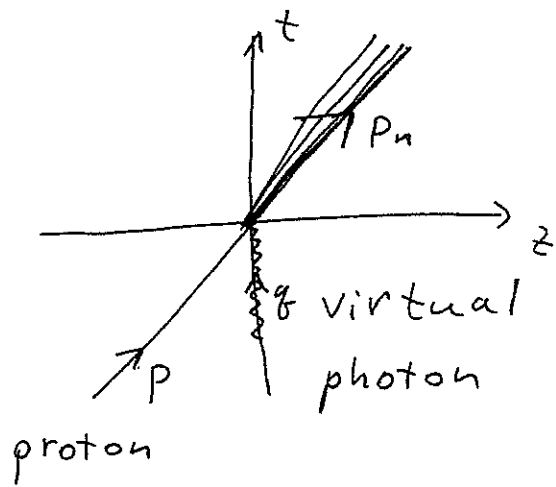
$$p_\mu \approx \left(p + \frac{m^2}{2p}, 0, 0, p \right),$$

$q = (q^0, q^1, q^2, q^3) \sim$ 2d vector in transverse plane

$$q_\mu = \left(q_0, \underline{q}, 0 \right),$$

Q^2 and x are 2 invariants
 \leftarrow large, $Q \gg \Lambda_{QCD}$

$$p \cdot q = m \nu = q_0 \cdot p$$



$\Rightarrow q_0 = \frac{m \nu}{p} \sim$ small as p goes large, $p \gg Q$

$$\Rightarrow \text{as } \nu = \frac{Q^2}{2m_p x} \Rightarrow q_0 = \frac{Q^2}{2xp}$$

$$\Rightarrow Q^2 = -q^2 = \underline{q}^2$$

$\Rightarrow q_0 \ll Q$ since $xp \gg Q$.

$$F_1(x, Q^2) = m_p W_1(x, Q^2)$$

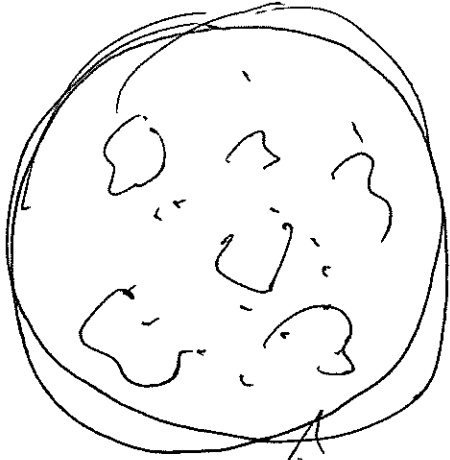
(86)

$$F_2(x, Q^2) = \nu W_2(x, Q^2) = \frac{Q^2}{2m_p x} W_2(x, Q^2)$$

$$\frac{d\sigma}{d^3k'} = \frac{4d_{EM}^2}{Q^4} \left[2 \cdot \frac{1}{m_p} F_1(x, Q^2) \sin^2 \frac{\theta}{2} + \frac{2m_p x}{Q^2} F_2(x, Q^2) \cdot \cos^2(\theta/2) \right]$$

F_1, F_2 are dimensionless

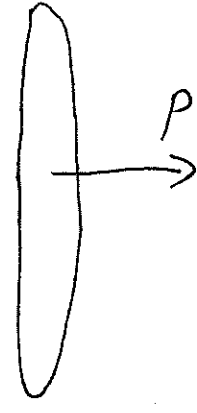
proton at rest



quantum fluctuations

infinite momentum frame (IMF)

ultra-boost
→



$$t_{fluct.} = \gamma \tau_{fluct} = \gamma \frac{1}{\Lambda_{QCD}}$$

$$\gamma = \frac{P}{m_p}$$

$$\Rightarrow t_{fluct} = \frac{P}{m_p} \frac{1}{\Lambda_{QCD}}$$

time scale

$$\tau_{fluct.} \approx \frac{1}{\Lambda_{QCD}}$$

← only scale in the problem

$$\frac{P}{m_p} = 7000 \text{ at LHC} \\ (\sqrt{s} = 14 \text{ TeV})$$

⇒ fluctuations live much longer in IMF ⇒
⇒ can probe them

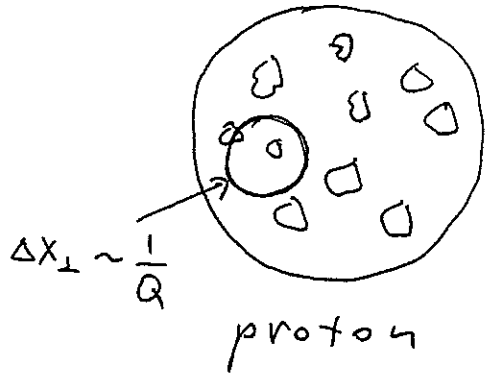
⇒ IMF "freezes" time for quantum fluctuations, allowing us to measure them.

$Q^2 = q^2 \Rightarrow$ photon acts like a microscope

in transverse plane:

$$\Delta x_{\perp} \cdot q_{\perp} \sim 1 \quad (h = 1)$$

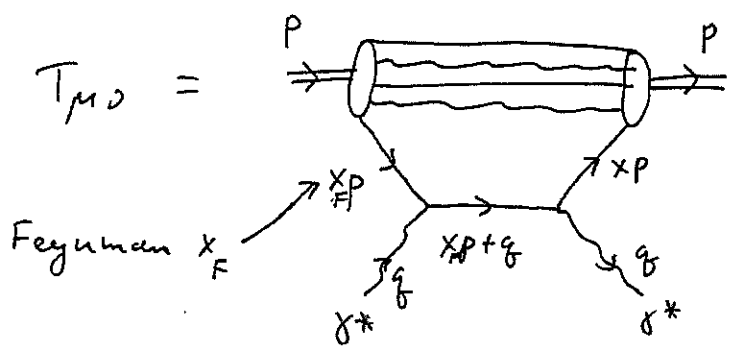
$$\Delta x_{\perp} \sim \frac{1}{q_{\perp}} \sim \frac{1}{Q}$$



large $Q \sim$ resolve just 1 quark

Define $T_{\mu\nu} = \frac{1}{4\pi m_p} \int d^4x e^{iq \cdot x} \frac{1}{2} \sum_{\sigma} \langle p, \sigma | T j_{\mu}(x) j_{\nu}(0) | p, \sigma \rangle$

$W_{\mu\nu} = 2 \text{Im} (i T_{\mu\nu})$ (optical theorem)



"Forward Amplitude"
 (Def) Feynman - x: the fraction of proton's longitudinal momentum carried by struck quark

typical interaction time in proton's rest frame

$\sim \frac{1}{\Lambda_{QCD}} \Rightarrow$ boost to get $\frac{P}{m} \frac{1}{\Lambda} \equiv \tau_{\Lambda}$

int. time of DIS is $\tau_{DIS} \approx \frac{1}{q^0}$, where $q^0 \approx \frac{Q^2}{2xP}$ is struck quark's velocity: $\tau_{DIS} \approx \frac{2xP}{Q^2}$

time-ordered product: (denoted T) (88)

$$T j_\mu(x) j_\nu(y) \equiv \theta(x^0 - y^0) j_\mu(x) j_\nu(y) + \theta(y^0 - x^0) j_\nu(y) j_\mu(x)$$

Note: currents do not commute with each other in general \Rightarrow not a trivial object.

$$2 \text{Im}(i T_{\mu\nu}) = 2 \text{Im} \left[i \cdot \frac{1}{4\pi m_p} \int d^4x e^{i q \cdot x} \langle p | \theta(x^0) j_\mu(x) j_\nu(0) + \theta(x^0) j_\nu(0) j_\mu(x) | p \rangle \right]$$

$$= 2 \cdot \frac{1}{4\pi m_p} \sum_n \text{Re} \left\{ \int d^4x e^{i q \cdot x + i p \cdot x - i p_n \cdot x} \theta(x^0) \langle p | j_\mu(0) | n \rangle \langle n | j_\nu(0) | p \rangle + \int d^4x e^{i q \cdot x + i p_n \cdot x - i p \cdot x} \theta(-x^0) \langle p | j_\nu(0) | n \rangle \langle n | j_\mu(0) | p \rangle \right\}$$

$$= 2 \frac{1}{4\pi m_p} \sum_n (2\pi)^3 \delta(\vec{q} + \vec{p} - \vec{p}_n) \left(\frac{-1}{i(q^0 + p^0 - p_n^0 + i\epsilon)} \langle p | j_\mu(0) | n \rangle \langle n | j_\nu(0) | p \rangle + (2\pi)^3 \delta(\vec{q} + \vec{p}_n - \vec{p}) \cdot \text{Re} \left(\frac{1}{i(q^0 + p_n^0 - p^0 - i\epsilon)} \langle p | j_\nu(0) | n \rangle \langle n | j_\mu(0) | p \rangle \right) \right)$$

not physical \Rightarrow drop (after including $\delta(q^0 + p_n^0 - p^0)$)

$$= 2 \frac{1}{4\pi m_p} \sum_n (2\pi)^3 \delta(\vec{q} + \vec{p} - \vec{p}_n) (-) \left(\text{Im} \frac{1}{q^0 + p^0 - p_n^0 + i\epsilon} \right) \cdot \langle p | j_\mu(0) | n \rangle \langle n | j_\nu(0) | p \rangle$$

$$\langle n | j_\nu(0) | p \rangle = \int dx \text{Im} \frac{1}{x + i\epsilon} = -\pi \delta(x) = \frac{1}{4\pi m_p} \sum_n (2\pi)^4 \delta^4(q + p - p_n)$$

$\langle p | j_\mu(0) | n \rangle \langle n | j_\nu(0) | p \rangle = W_{\mu\nu}$ as desired.
 (can prove that $\langle p | j_\mu(0) | n \rangle \langle n | j_\nu(0) | p \rangle$ is real)