

Last time: talked more about chiral symmetry

$$N_f = 2$$

$$\mathcal{L} = \bar{q}_L i\gamma^5 \partial_\mu q_L + \bar{q}_R i\gamma^5 \partial_\mu q_R$$

$\Rightarrow \text{SU}(2)_L \otimes \text{SU}(2)_R$  chirally symmetric.

Conserved currents:

$$j_L^{i\mu} = \bar{q}_L \gamma^\mu \frac{\sigma^i}{2} q_L$$

$$j_R^{i\mu} = \bar{q}_R \gamma^\mu \frac{\sigma^i}{2} q_R$$

Conserved charges:

$$Q_{L,R}^i(t) = \int d^3x j_{L,R}^i(\vec{x}, t)$$

$$\frac{dQ_{L,R}^i}{dt} = 0$$

The charges are generators  
of  $\text{SU}(2)_L \otimes \text{SU}(2)_R$ :

$$[Q_L^i, Q_L^j] = i\varepsilon^{ijk} Q_L^k$$

$$[Q_R^i, Q_R^j] = i\varepsilon^{ijk} Q_R^k$$

$$[Q_L^i, Q_R^j] = 0$$

Problem:

$$(N_f = 3)$$

$$\text{SU}(3)_L \otimes \text{SU}(3)_R$$

$$m_u = m_d = m_s \neq 0$$

$\Rightarrow$  need SSB here

(spontaneous symmetry  
breaking)

Both symmetries  
are equally  
broken, but  
 $\text{SU}(3)$  is still  
observed in  
nature, while  
 $\text{SU}(3)_L \otimes \text{SU}(3)_R$   
is not!

$$\text{SU}(3)$$

$$m_u \neq m_d \neq m_s \neq 0$$

Nothing

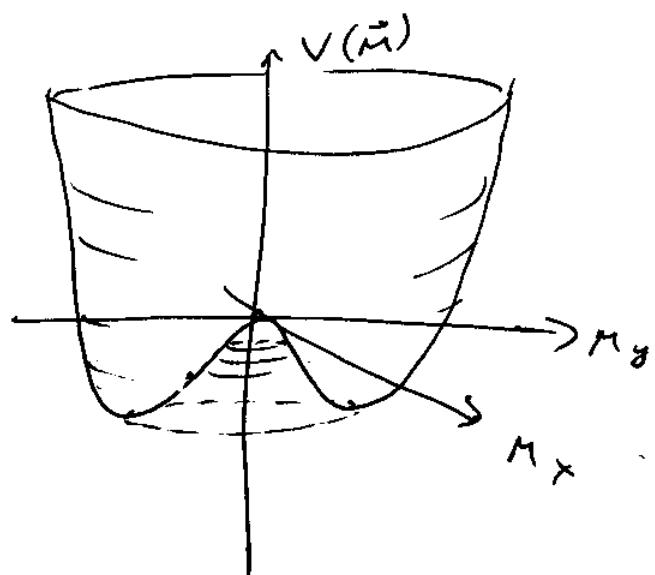
Def. SSB : a symmetry manifest in  $\mathcal{L}, H$ , but not respected by ground state.

Example : Ising model ,  $\begin{matrix} \uparrow & \uparrow & \uparrow \\ \uparrow & \uparrow & \uparrow \\ \uparrow & \uparrow & \uparrow \end{matrix}$  vs  $\begin{matrix} \downarrow & \downarrow & \downarrow \\ \downarrow & \downarrow & \downarrow \\ \downarrow & \downarrow & \downarrow \end{matrix}$ .

Landau - Ginzburg theory of ferromagnetism  
(cont'd)

$$H = \int d^3x \left[ (\nabla_i M_j)^2 + \underbrace{\mu^2 (T - T_c) \vec{M}^2}_{V} + \lambda (\vec{M}^2)^2 \right]$$

$$\lambda, \mu^2 > 0, T \sim \text{temperature}$$



for  $T < T_c$

$$M^2 \rightarrow M'^2 = R^{ij} M_i$$

SO(3) rotational symmetry

Ground state:

$$|\vec{M}_{\text{gr.}}| = \sqrt{\frac{\mu^2 (T_c - T)}{2\lambda}} \neq 0$$

Direction of  $\vec{M}_{\text{vac}}$  is random  $\approx \text{SSB!}$

$$\vec{x}'_i = R_{ij} \vec{x}_j \Rightarrow |\vec{x}'|^2 = |\vec{x}|^2 \Rightarrow R_{ij} R_{ik} \vec{x}_j \cdot \vec{x}_k = \vec{x}_i \cdot \vec{x}_i$$

$$\Rightarrow R_{ij} R_{ik} = \delta_{jk} \Rightarrow R \cdot R^T = R^T R = \mathbb{1} \Rightarrow \text{forget}$$

reflections  $\Rightarrow$  require  $\det R = +1 \Rightarrow SO(3)$

$\sim$  a group of special ( $\det = +1$ )  $\xrightarrow{\text{real}}$  orthogonal ( $RR^T = R^T R = \mathbb{1}$ )  $3 \times 3$  matrices.

$\Rightarrow$  for  $T < T_c$  the ground state is at the minima

$$\Rightarrow \mu^2 (T - T_c) 2 |\vec{M}| + 4\lambda |\vec{M}|^3 = 0$$

$$\Rightarrow |\vec{M}_{\text{vac}}| = \sqrt{\frac{\mu^2 (T_c - T)}{2\lambda}}$$

$\Rightarrow$  however, direction of  $\vec{M}$  is chosen spontaneously!

$$\text{Say, } M_{\text{vac}} = \sqrt{\frac{\mu^2 (T_c - T)}{2\lambda}} \hat{x} = |0\rangle$$

define generators of  $SO(3)$ :  $L_1 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & -i \\ 0 & i & 0 \end{pmatrix}$ ,  $L_2 = \begin{pmatrix} 0 & 0 & i \\ 0 & 0 & 0 \\ -i & 0 & 0 \end{pmatrix}$ ,

$L_3 = \begin{pmatrix} 0 & -i & 0 \\ i & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \Rightarrow e^{-i\vec{\alpha} \cdot \vec{L}}$  is a rotation by angle

$i\vec{\alpha}$  around  $\vec{\alpha}$ -direction.

$$\Rightarrow H \text{ is invariant under } \vec{M} \rightarrow \vec{M}' = e^{-i\vec{\alpha} \cdot \vec{L}} \vec{M}.$$

$\Rightarrow$  ground state is not rotationally symmetric:

$R|0\rangle \neq |0\rangle \Rightarrow$  if  $R = e^{i\vec{\alpha} \cdot \vec{Q}}$ ,  $\vec{Q}$  ~ conserved

charges of symmetry  $\Rightarrow Q^i |0\rangle \neq 0 \quad \begin{cases} \text{equivalently} \\ \langle 0 | \vec{M} | 0 \rangle \neq 0 \end{cases}$

## General Discussion

Imagine a system with Hamiltonian  $H$  and conserved symmetry charges  $Q^i$ :  $[H, Q^i] = 0$ .

Act on vacuum:  $H|0\rangle = 0$  (can choose vacuum to be 0-energy state)

$$H Q^i |0\rangle = \underbrace{[H, Q^i]}_{=0} |0\rangle + Q^i \underbrace{H|0\rangle}_{=0} = 0$$

$$\Rightarrow H Q^i |0\rangle = 0 \Rightarrow \text{either}$$

(i)  $Q^i |0\rangle = 0 \sim \text{no broken symmetries, vacuum is invariant under } Q^i: e^{i\vec{E} \cdot \vec{Q}} |0\rangle = |0\rangle$ .

(ii)  $Q^i |0\rangle \neq 0 \Rightarrow \text{vacuum is degenerate, more than one state such that } H|4_0\rangle = 0$ .

(e.g. rotating ground state in L-G model would give other possible ground states)

$\Rightarrow$  if the system spontaneously chooses one of these  $|4_0\rangle$  states for its ground state  $\Rightarrow$  spontaneous symmetry breaking.

## The Nambu - Goldstone Theorem.

Theorem Spontaneous breakdown of a continuous symmetry implies existence of massless spinless particles. (Nambu - Goldstone bosons)

(Nambu '60, Goldstone '61)

Proof  $j^M$  is a conserved current  $\partial_\mu j^M = 0$ ,  
 $Q(t) = \int d^3x j_0(\vec{x}, t)$  is the conserved charge.

For generic field  $\varphi(x)$ :

$$\varphi(x) \rightarrow \varphi'(x) = e^{i\alpha Q} \varphi(x) e^{-i\alpha Q} = \varphi(x) + i\alpha [Q, \varphi] + \dots$$

$$\Rightarrow 0 = \int d^3x [\partial_\mu j^M(\vec{x}, t), \varphi(0)] = \partial_0 \int d^3x [j^0(\vec{x}, t, \varphi(0))]$$

+ spatial surface term

$$\Rightarrow \frac{d}{dt} [Q(t), \varphi(0)] = 0 \Rightarrow \langle 0 | [Q(t), \varphi(0)] | 0 \rangle = v \neq 0$$

with  $v$  time-independent (constant) quantity

$$v = \langle 0 | [Q(t), \varphi(0)] | 0 \rangle = \langle 0 | Q(t) \varphi(0) | 0 \rangle - \langle 0 | \varphi(0) Q(t) | 0 \rangle = \int d^3x \left[ \langle 0 | j_0(t, \vec{x}) \varphi(0) | 0 \rangle \right]$$

$$- \langle 0 | \psi(0) j_0(t, \vec{x}) | 0 \rangle ] .$$

Insert a complete set of intermediate states

$$1 = \sum_n |n\rangle \langle n| \Rightarrow \text{get}$$

$$\nu = \sum_n \int d^3x \left[ \langle 0 | j_0(t, \vec{x}) | n \rangle \langle n | \psi(0) | 0 \rangle - \langle 0 | \psi(0) | n \rangle \langle n | j_0(t, \vec{x}) | 0 \rangle \right]$$

Now, in Heisenberg picture one can write

$$j_0(t, \vec{x}) = j_0(x) = e^{i\hat{p} \cdot x} j_0(0) e^{-i\hat{p} \cdot x}$$

where  $\hat{p}^\mu =^{(H, p)}$  is the 4-momentum operator,  $\hat{p}^\mu |0\rangle = 0$

$$\Rightarrow \nu = \int d^3x \sum_n \left[ \langle 0 | e^{i\hat{p} \cdot x} j_0(0) e^{-i\hat{p} \cdot x} | n \rangle \langle n | \psi(0) | 0 \rangle - \langle 0 | \psi(0) | n \rangle \langle n | e^{i\hat{p} \cdot x} j_0(0) e^{-i\hat{p} \cdot x} | 0 \rangle \right]$$

$$\text{Take } \langle 0 | e^{i\hat{p} \cdot x} j_0(0) e^{-i\hat{p} \cdot x} | n \rangle = \langle 0 | j_0(0) | n \rangle \cdot e^{-ip_n \cdot x}$$

$$\Rightarrow \nu = (2\pi)^3 \sum_n \delta^3(\vec{p}_n) \left[ e^{-iE_n t} \langle 0 | j_0(0) | n \rangle \cdot \langle n | \psi(0) | 0 \rangle - \langle 0 | \psi(0) | n \rangle \langle n | j_0(0) | 0 \rangle e^{iE_n t} \right]$$

$$\text{LHS } \nu = \text{time-independent (constant)}, \nu \neq 0 \text{ as SSB makes it } \neq 0$$

integrate over time:  $\lim_{T \rightarrow \infty} \int_{-T}^T dt \Rightarrow$

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$$\Rightarrow \lim_{T \rightarrow \infty} (2T) \cdot v = (2\pi)^4 \sum_n \delta^3(\vec{p}_n) \delta(E_n) [\langle 0 | j_0(0) | n \rangle -$$

$$\langle n | \varphi(0) | 0 \rangle - \langle 0 | \varphi(0) | n \rangle \langle n | j_0(0) | 0 \rangle]$$

as  $\lim_{T \rightarrow \infty} \int_{-T}^T dt = 2\pi \delta(0)$

$$\Rightarrow 2\pi \delta(0) v = (2\pi)^4 \sum_n \delta^3(\vec{p}_n) \delta(E_n) [\langle 0 | j_0(0) | n \rangle -$$

$$\langle n | \varphi(0) | 0 \rangle - \langle 0 | \varphi(0) | n \rangle \langle n | j_0(0) | 0 \rangle]$$

$\Rightarrow$  at  $\vec{p}_n = 0$  have a spectrum of states  $E_n$

$\Rightarrow$  for equation to hold need to have at least one state with  $E_n = 0 \Rightarrow$  get  $\delta(0)$  too. (all other  $E_n$ 's give zero contributions)

$\Rightarrow$  there must be a state with  $E_n = 0, \vec{p}_n = 0$

$\Rightarrow$  a massless particle (Goldstone boson)

(or Nambu-Goldstone boson)

$\Rightarrow \langle n | \varphi(0) | 0 \rangle \neq 0, \langle 0 | j_0(0) | n \rangle \neq 0$  for state  $|n\rangle$

$\Rightarrow \varphi \sim$  scalar field  $\Rightarrow$  boson.

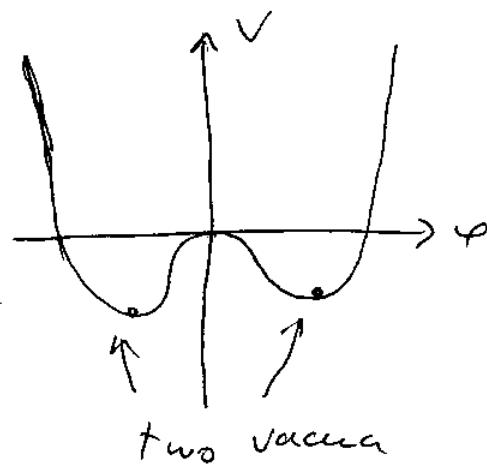
Example 1:  $\varphi$  - real scalar field

$$\mathcal{L} = \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi + \underbrace{\frac{m^2}{2} \varphi^2 - \frac{\lambda}{4} \varphi^4}_{-V(\varphi)} \Rightarrow \varphi \rightarrow -\varphi \text{ symmetric}$$

$m^2 > 0 \Rightarrow$  symmetry is broken.

Vacuum:  $m^2 \cdot \varphi - \frac{\lambda}{4} \varphi^3 = 0$

$$\Rightarrow \varphi = \pm v = \pm m \cdot \sqrt{\frac{1}{\lambda}}$$



$\Rightarrow$  once the system picks  $+v$  or  $-v$

the  $\varphi \rightarrow -\varphi$  is spontaneously broken.

$\Rightarrow$  Near the vacuum at  $v$  write  $\varphi = v + \varphi'$

$$\begin{aligned} \Rightarrow \mathcal{L} &= \frac{1}{2} \partial_\mu \varphi' \partial^\mu \varphi' + \frac{m^2}{2} (v + \varphi')^2 - \frac{\lambda}{4} (v + \varphi')^4 = \\ &= (\text{drop constants}) = \frac{1}{2} \partial_\mu \varphi' \partial^\mu \varphi' + \left( \frac{m^2}{2} - \frac{\lambda}{4} v^2 \right) \varphi'^2 \end{aligned}$$

$$\# - \underbrace{\frac{\lambda}{4} \cdot 6 v^2 \varphi'^2}_{m^2 \frac{3}{2}} - \frac{\lambda}{4} \cdot 4 v \varphi'^3 + \frac{m^2}{2} \varphi'^2 = \frac{\lambda}{4} \varphi'^4 =$$

$$= \frac{1}{2} \partial_\mu \varphi' \partial^\mu \varphi' - m^2 \varphi'^2 - \frac{\lambda}{4} v \varphi'^3 - \frac{\lambda}{4} \varphi'^4$$

$\Rightarrow \varphi'$  has mass  $= \sqrt{m^2 + \frac{\lambda v^2}{4}}$   $\neq$  massless!

$\Rightarrow$  Is Goldstone theorem wrong? No, it's just that  $\varphi \rightarrow -\varphi$  symmetry is discrete!  
(G. thm is about continuous symmetries.)

## Example 2 : Abelian O-Model:

$$\mathcal{L} = \frac{1}{2} \partial_\mu \sigma \partial^\mu \sigma + \frac{1}{2} \partial_\mu \bar{\sigma} \partial^\mu \bar{\sigma} + \underbrace{\frac{\mu^2}{2} (\sigma^2 + \bar{\sigma}^2)}_{-V} - \frac{\lambda}{4} (\sigma^2 + \bar{\sigma}^2)^2$$

with  $\mu^2 > 0, \lambda > 0$  (constants).

$\sigma, \bar{\sigma}$  ~ real fields

$\mathcal{L}$  is invariant under rotations:

$$\begin{pmatrix} \sigma \\ \bar{\sigma} \end{pmatrix} \rightarrow \begin{pmatrix} \sigma' \\ \bar{\sigma}' \end{pmatrix} = \begin{pmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{pmatrix} \begin{pmatrix} \sigma \\ \bar{\sigma} \end{pmatrix}, \quad \alpha \text{~real #}$$

$\Rightarrow O(2)$  symmetry ( $= U(1)$ ).

$\Rightarrow$  get "Mexican hat" potential again:

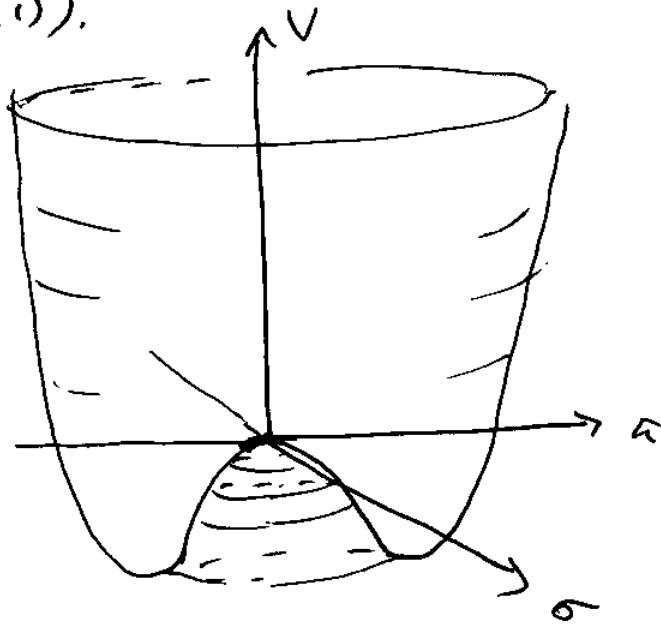
$\Rightarrow$  the minimum is at

$$\sigma^2 + \bar{\sigma}^2 = v^2$$

$$\Rightarrow \left( \frac{\mu^2}{2} \cdot v^2 - \frac{\lambda}{4} v^4 \right)'_v = 0$$

$$\mu^2 \cdot v - \lambda v^3 = 0$$

$$\Rightarrow v = \mu \sqrt{\frac{1}{\lambda}}$$



$\Rightarrow$  direction in  $(\sigma, \bar{\sigma})$  space is random  $\Rightarrow$

$\Rightarrow$  pick the vacuum to be at  $\langle 0 | \sigma | 0 \rangle = v, \langle 0 | \bar{\sigma} | 0 \rangle = 0$ .

Expand  $\sigma$  near the vacuum:  $\sigma = v + \sigma'$

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$$\begin{aligned} \Rightarrow \mathcal{L} &= \frac{1}{2} \partial_\mu \sigma' \partial^\mu \sigma' + \frac{1}{2} \partial_\mu \bar{\pi} \partial^\mu \bar{\pi} + \frac{m^2}{2} [ (v + \sigma')^2 + \bar{\pi}^2 ] \\ &- \frac{\lambda}{4} [ (v + \sigma')^2 + \bar{\pi}^2 ]^2 = \frac{1}{2} \partial_\mu \sigma' \partial_\mu \sigma' + \frac{1}{2} \partial_\mu \bar{\pi} \partial^\mu \bar{\pi} + \underbrace{\text{constant}}_{\text{drop}} \\ &+ \sigma' \left[ \frac{m^2 v - \lambda \cdot 4 v^3}{4} \right] \rightarrow 0 + \sigma'^2 \left[ \frac{m^2}{2} - \frac{\lambda}{4} \cdot (2v^2 + 4v^2) \right] + \\ &+ \bar{\pi}^2 \left[ \frac{m^2}{2} - \frac{\lambda}{4} \cdot 2v^2 \right] \rightarrow 0 - \frac{\lambda}{4} [ 4 \sigma' v (\sigma'^2 + \bar{\pi}^2) + (\sigma'^2 + \bar{\pi}^2)^2 ] \\ &= \frac{1}{2} \partial_\mu \sigma' \partial_\mu \sigma' + \frac{1}{2} \partial_\mu \bar{\pi} \partial^\mu \bar{\pi} - m^2 \sigma'^2 - \lambda v \sigma' (\sigma'^2 + \bar{\pi}^2) \\ &- \frac{\lambda}{4} (\sigma'^2 + \bar{\pi}^2)^2. \end{aligned}$$

$\Rightarrow$  now  $\bar{\pi}$ 's have no  $\bar{\pi}^2$  term  $\Rightarrow$   $\bar{\pi}$  field is massless in agreement with Goldstone thm!

### Non-Abelian $\sigma$ -Model

Let's illustrate how the chiral  $SU(3)_c \otimes SU(3)_R$  symmetry is broken in QCD. As an example consider breaking of  $SU(2)_L \otimes SU(2)_R$  symmetry.

Start with the Lagrangian:

$$\begin{aligned} \mathcal{L} &= \frac{1}{2} (\partial_\mu \sigma \partial^\mu \sigma + \partial_\mu \vec{\pi} \cdot \partial^\mu \vec{\pi}) + \frac{m^2}{2} (\sigma^2 + \vec{\pi}^2) - \frac{\lambda}{4} [\sigma^2 + \vec{\pi}^2]^2 \\ m^2, \lambda > 0, \quad \sigma, \vec{\pi} &= \underbrace{(\pi_1, \pi_2, \pi_3)}_{\text{iso triplet, pions}} \sim \text{real fields} \end{aligned}$$

Define a  $\overset{2 \times 2}{\sigma}$  matrix field  $\Sigma = \sigma \mathbb{1} + i \vec{\tau} \cdot \vec{n}$

$\tau^1, \tau^2, \tau^3 \sim$  Pauli matrices (we use  $\vec{\tau}$  to not confuse them with  $\sigma$ )

$$\Rightarrow \text{tr} [\Sigma \Sigma^+] = \text{tr} [\sigma^2 \mathbb{1} + i \vec{\tau} \cdot \vec{n} (-i) \vec{\tau} \cdot \vec{n}] \\ = 2 \sigma^2 + 2 \vec{n}^2 \quad \text{as } \text{tr} \tau^i \tau^j = 2 \delta^{ij}$$

$$\Rightarrow \text{tr} [\partial_\mu \Sigma \partial^\mu \Sigma^+] = 2 [\partial_\mu \sigma \partial^\mu \sigma + \partial_\mu \vec{n} \partial^\mu \vec{n}]$$

$$\Rightarrow \mathcal{L}_\Sigma = \frac{1}{4} [\text{tr} \partial_\mu \Sigma \partial^\mu \Sigma^+] + \frac{M^2}{4} \text{tr} [\Sigma \Sigma^+] - \frac{\lambda}{16} (\text{tr} [\Sigma \Sigma^+])^2$$

Now add "quarks": (originally they were protons and neutrons):  $q = \begin{pmatrix} u \\ d \end{pmatrix}$  or  $\begin{pmatrix} p \\ n \end{pmatrix} = q^N$

$$\mathcal{L} = \bar{q}^N i \gamma \cdot \partial q^N - g \bar{q}^N [\sigma \mathbb{1} + i \vec{\tau} \cdot \vec{n} \gamma_5] q^N + \mathcal{L}_\Sigma$$

such that

$$\mathcal{L} = \bar{q}^N i \gamma \cdot \partial q^N - g \bar{q}^N [\sigma \mathbb{1} + i \vec{\tau} \cdot \vec{n} \gamma_5] q^N + \\ + \frac{1}{2} (\partial_\mu \sigma \partial^\mu \sigma + \partial_\mu \vec{n} \partial^\mu \vec{n}) + \frac{M^2}{2} (\sigma^2 + \vec{n}^2) - \frac{\lambda}{4} (\sigma^2 + \vec{n}^2)^2$$

full Lagrangian for  $SU(2)_L \otimes SU(2)_R$  6-model.

(Gell-Mann & Levy, 1960)

As usual write  $q^N = q_L^N + q_R^N \Rightarrow$

$$\bar{q}^N i \gamma \cdot \partial q^N = \bar{q}_L^N i \gamma \cdot \partial q_L^N + \bar{q}_R^N i \gamma \cdot \partial q_R^N$$