

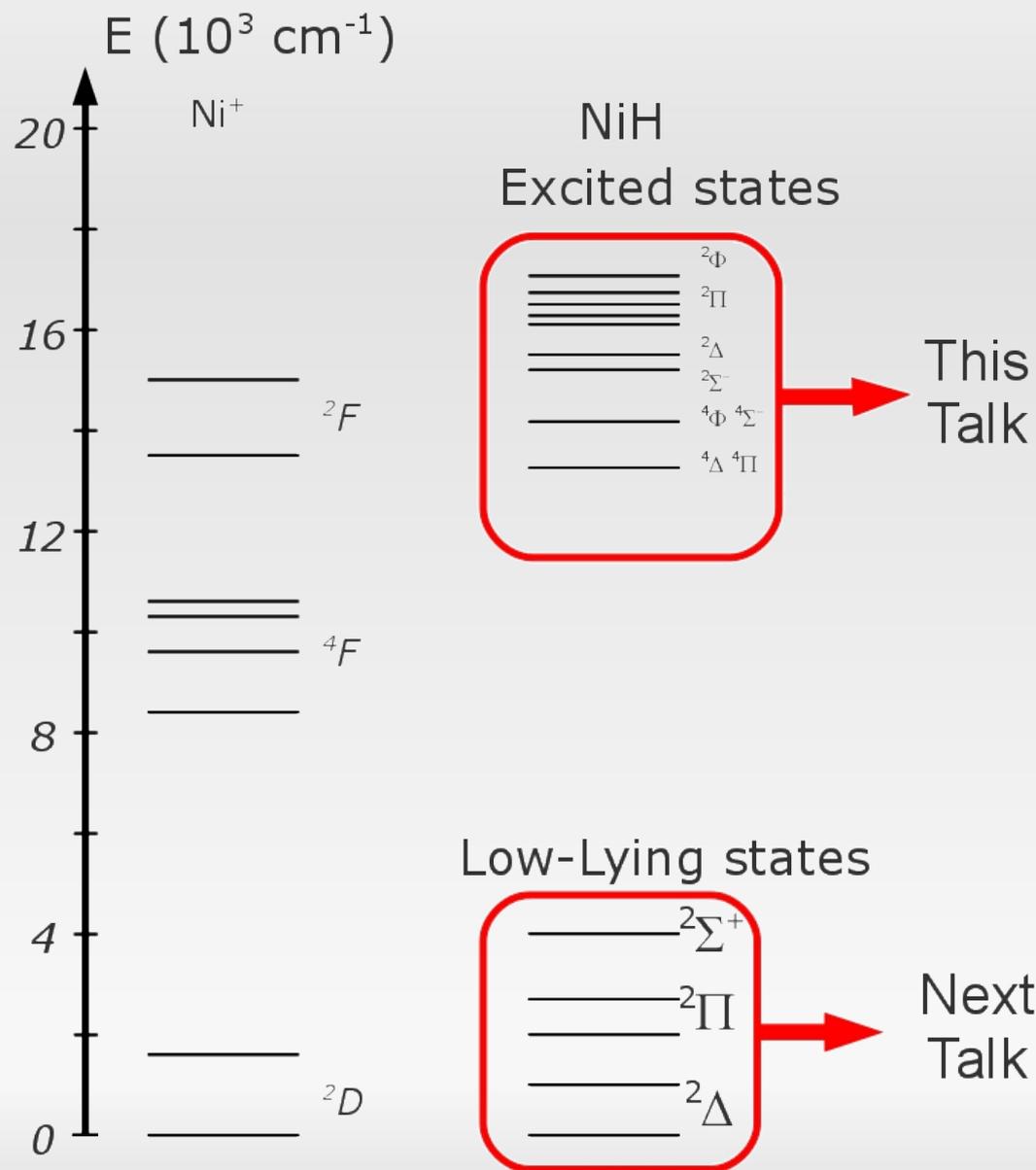
LASER EXCITATION SPECTROSCOPY OF ^{58}NiH IN A MAGNETIC FIELD

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NiH in a magnetic field

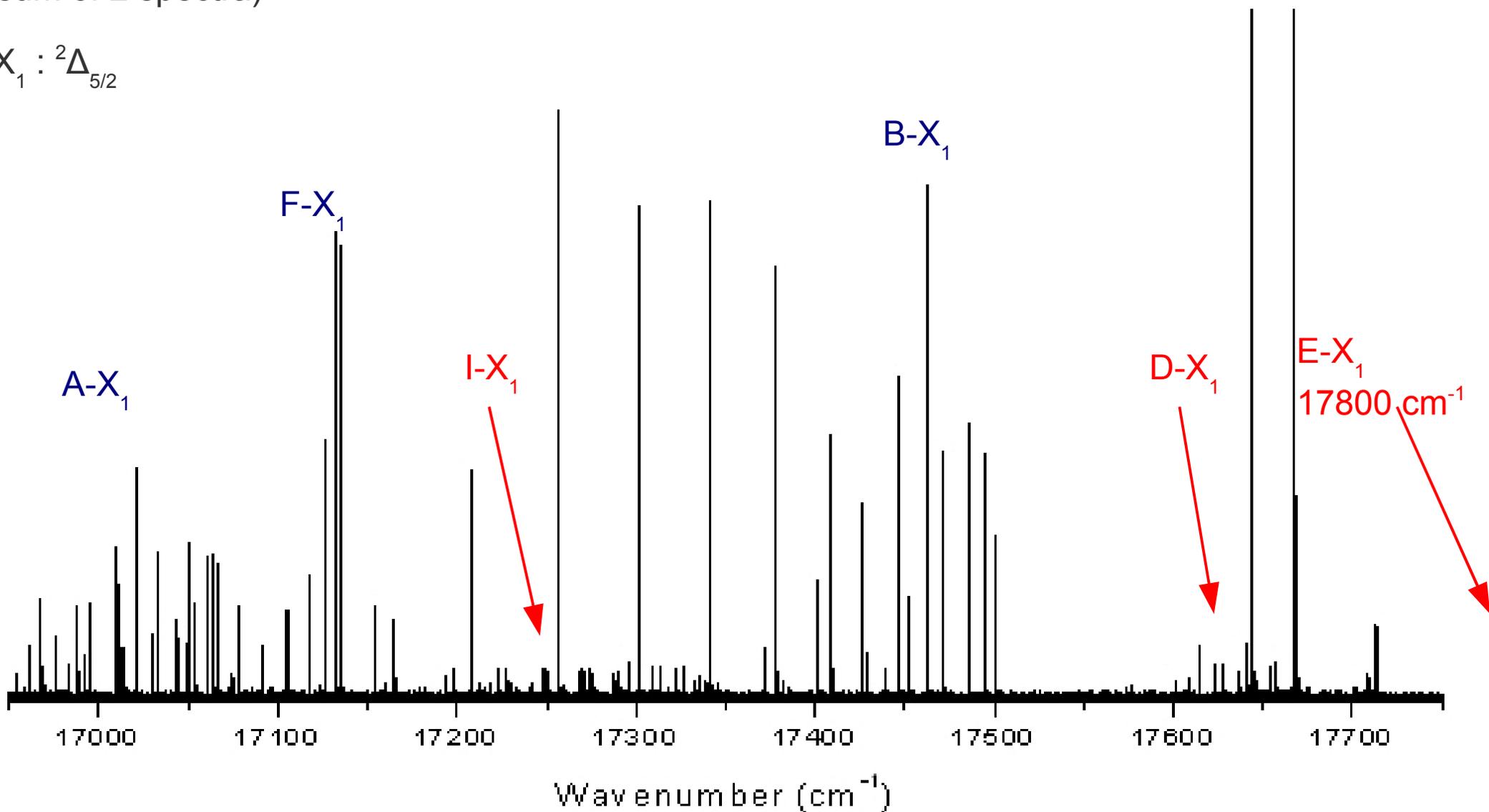
- Sputter source inserted between permanent magnets to investigate magnetic response in fields up to 1 Tesla
- NiH is used to optimize experimental conditions
- Studies are a prelude to work on FeH $\sim 1 \mu\text{m}$

Electronic spectrum of NiH ~ 570 nm

Dispersed fluorescence by Fourier Transform spectroscopy
(sum of 2 spectra)

Laser lines

$X_1 : ^2\Delta_{5/2}$

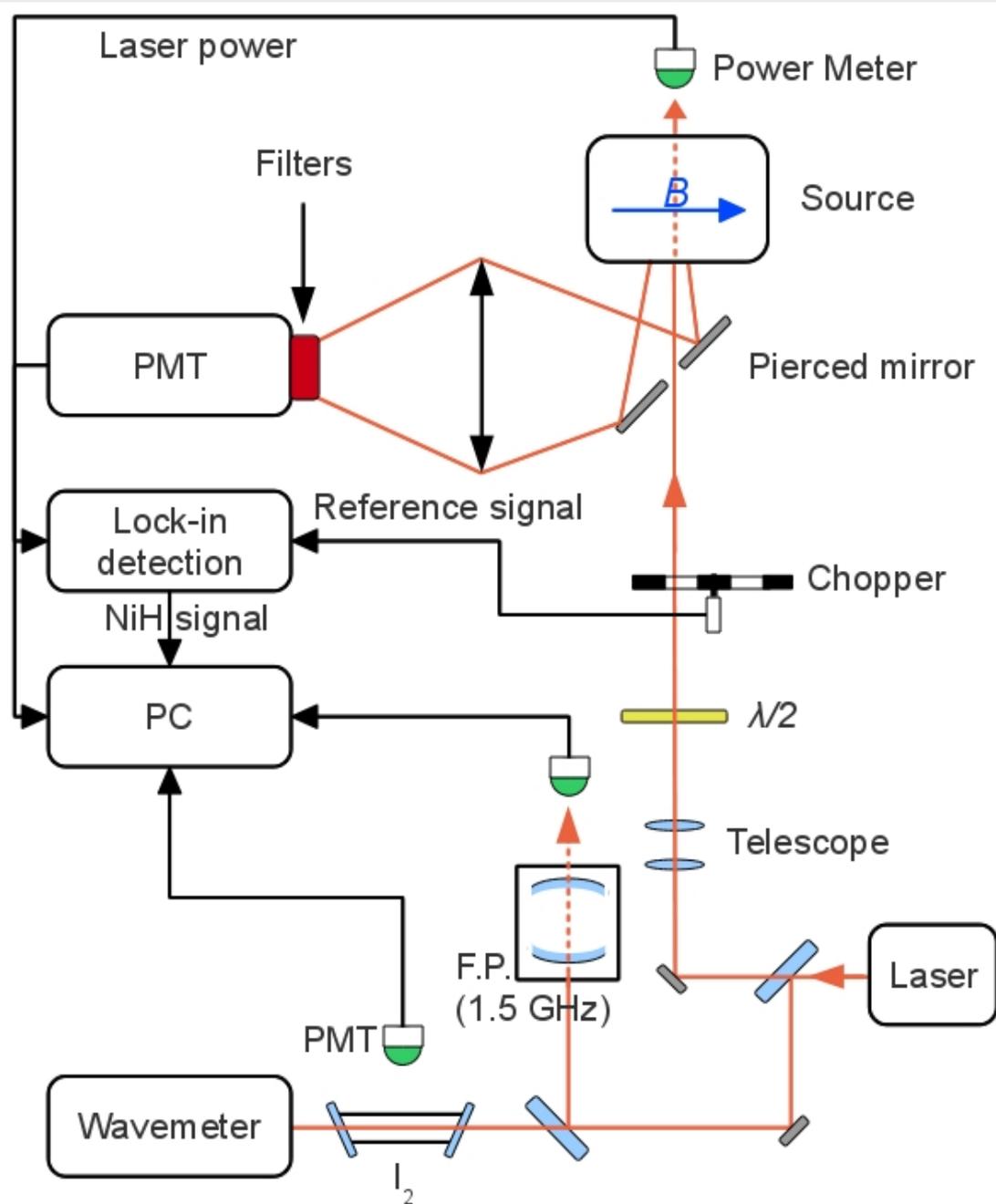


Aims

- FT dispersed fluorescence suggests $\Omega=3/2$ states ($I \rightarrow W_1 (^2\Pi_{3/2})$) gives well resolved magnetic response at high J
 - Study excited states with $\Omega=3/2$
- We plan to study $\Lambda_{3/2} \leftarrow X$ transitions to find upper state Landé factors.

Transitions from $X (^2\Delta_{5/2})$ are WEAK

Experimental setup



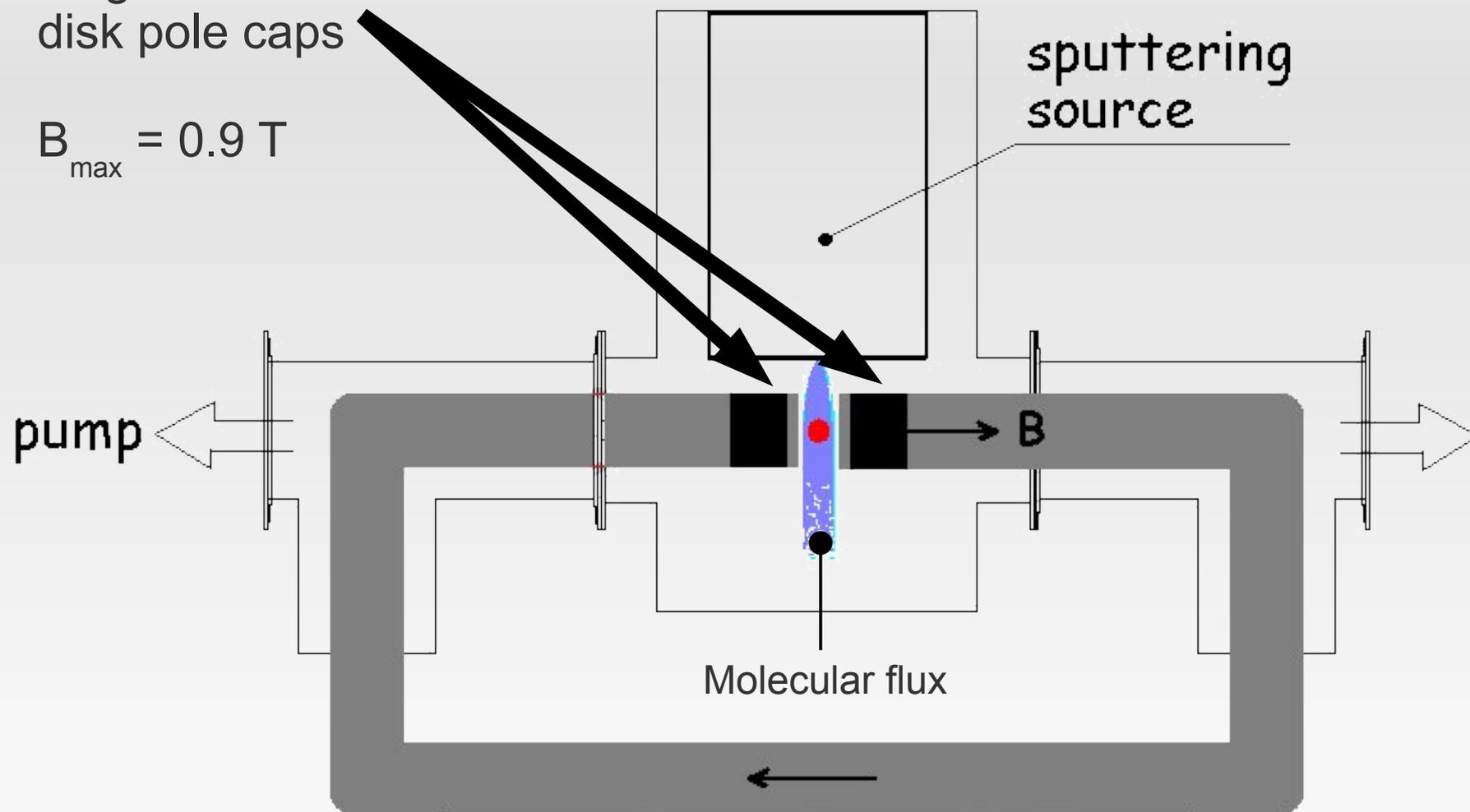
- Spectrum recorded in 1 cm^{-1} scans with a ring dye laser, ($17000 - 17800 \text{ cm}^{-1}$)
- Lock-in detection $\tau \sim 30 \text{ ms}$
- Check for mode hops (FP fringes)
- Calibrate with I_2 spectra

Experimental setup

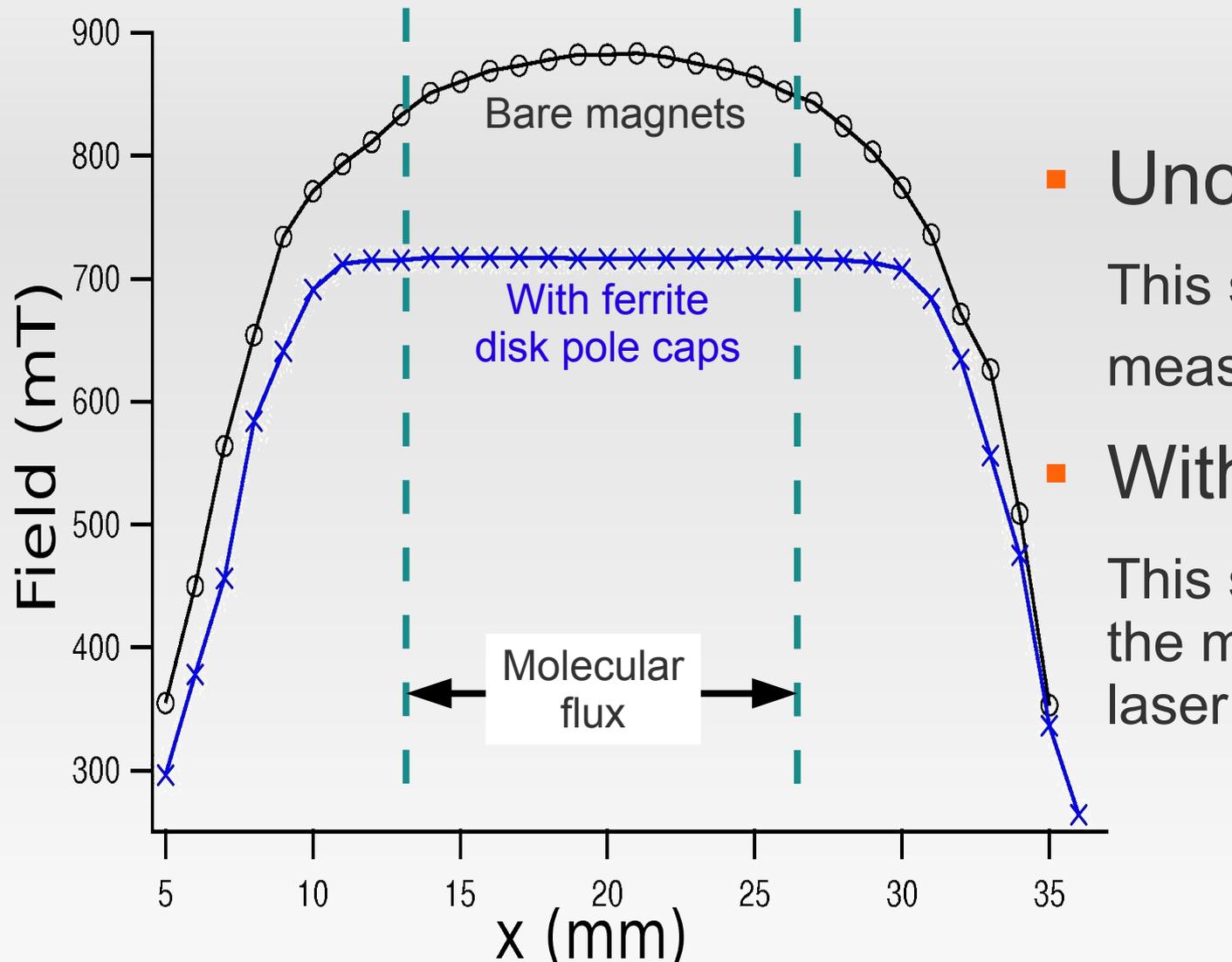
Permanent magnets with disk pole caps

$$B_{\text{max}} = 0.9 \text{ T}$$

sputtering source



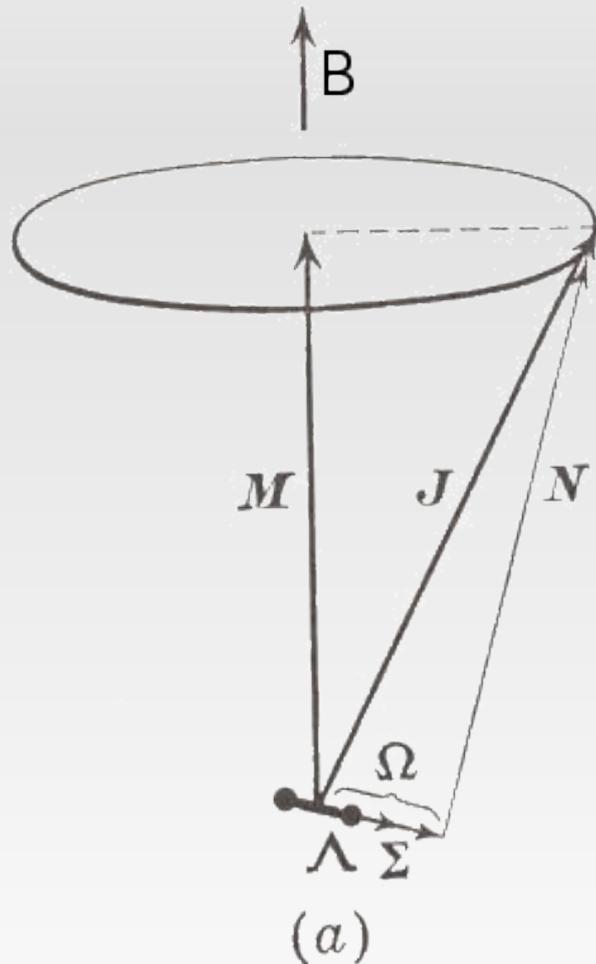
Magnetic field between the pole caps



- **Uncovered magnets**
This shows non-uniform B_0 measured on a Hall probe
- **With Disk Pole Caps**
This shows a flat profile of the magnetic field along the laser beam trajectory

Axis corresponds to the laser beam trajectory

Zeeman effect



- Energy splitting :

$$E_0 + \frac{\mu_B B M_J \Omega}{J(J+1)} g_{eff}$$

with $g_{eff} = \Lambda + g_s \Sigma$ in Hund's case (a)

Example for Hund's case (a) :

$${}^2\Sigma : g_{eff} = 1,$$

$${}^2\Pi_{1/2} : g_{eff} = 0$$

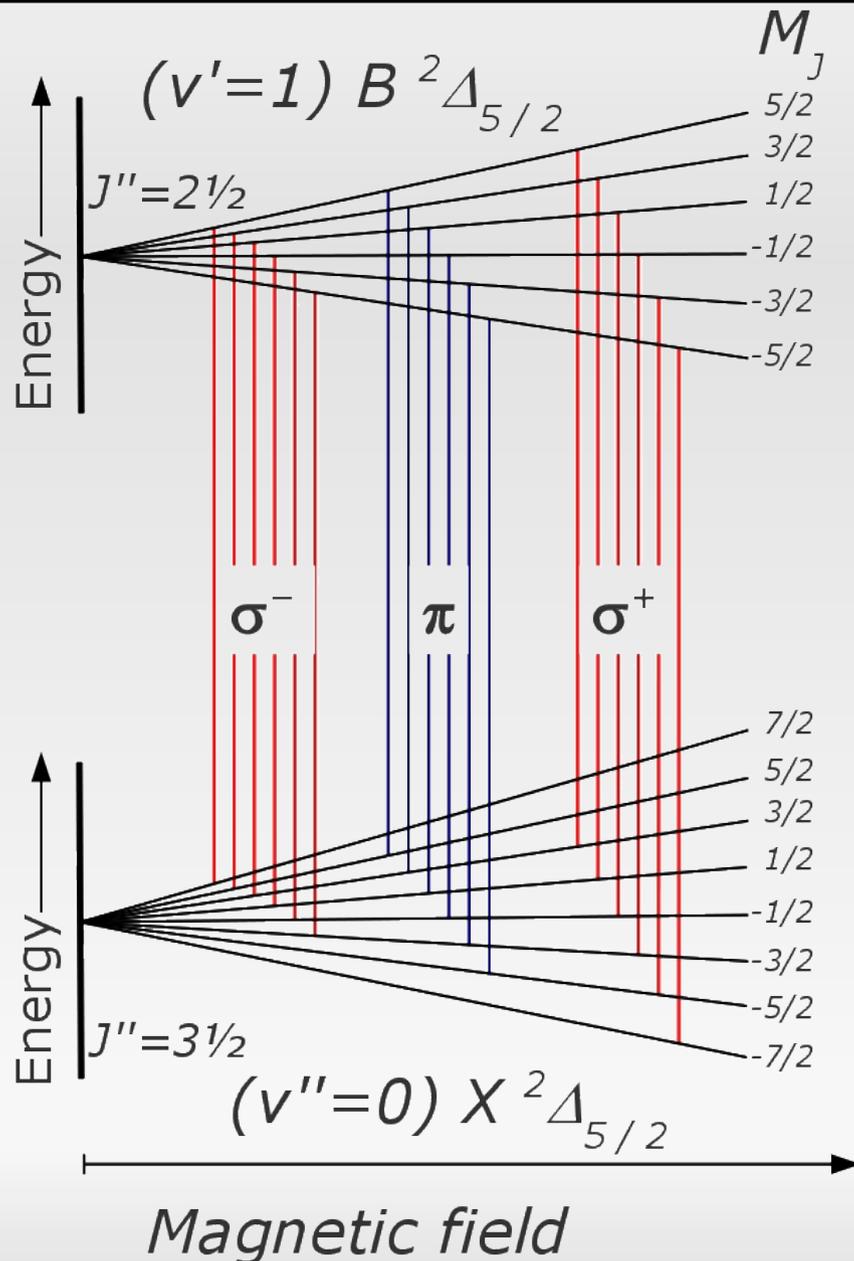
$${}^2\Pi_{3/2} : g_{eff} = 2$$

$${}^2\Delta_{3/2} : g_{eff} = 1$$

$${}^2\Delta_{5/2} : g_{eff} = 3$$

We know Ω for the excited states from observation of first lines ($\Omega = J_{\min}$), but Λ , Σ and S are less obvious

Zeeman spectroscopy



$\Delta M_J = +1 : \sigma^+$ transition

$\Delta M_J = -1 : \sigma^-$ transition

Laser field is perpendicular to the external field

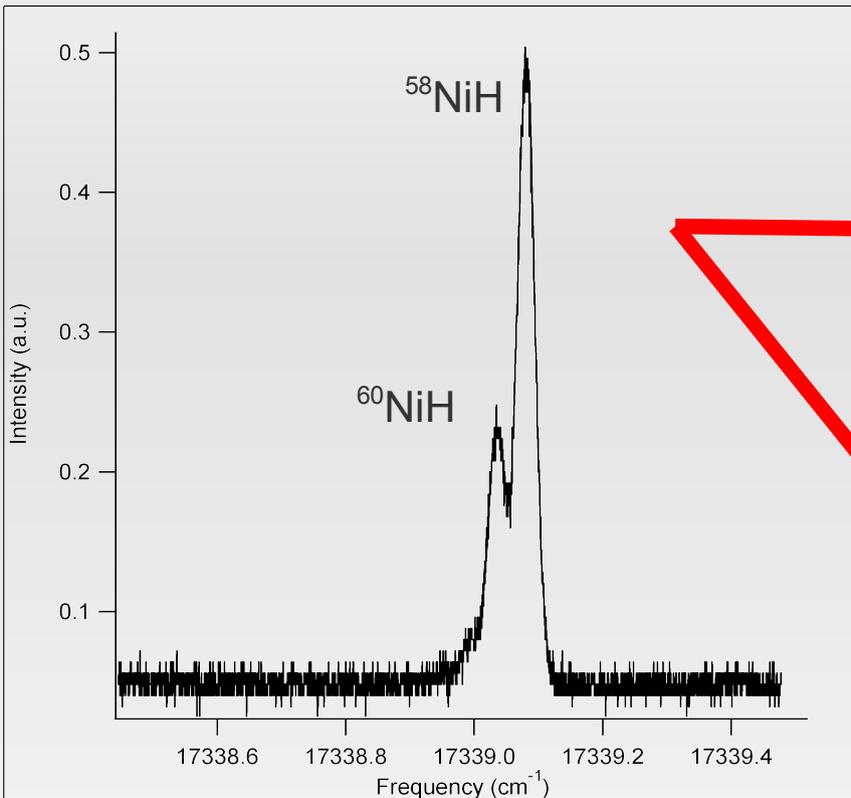
$\Delta M_J = 0 : \pi$ transition

Laser is parallel polarized

P($3\frac{1}{2}$) line

Examples of $I \leftarrow X_1$ spectra

Note isotopic overlap even in zero field conditions



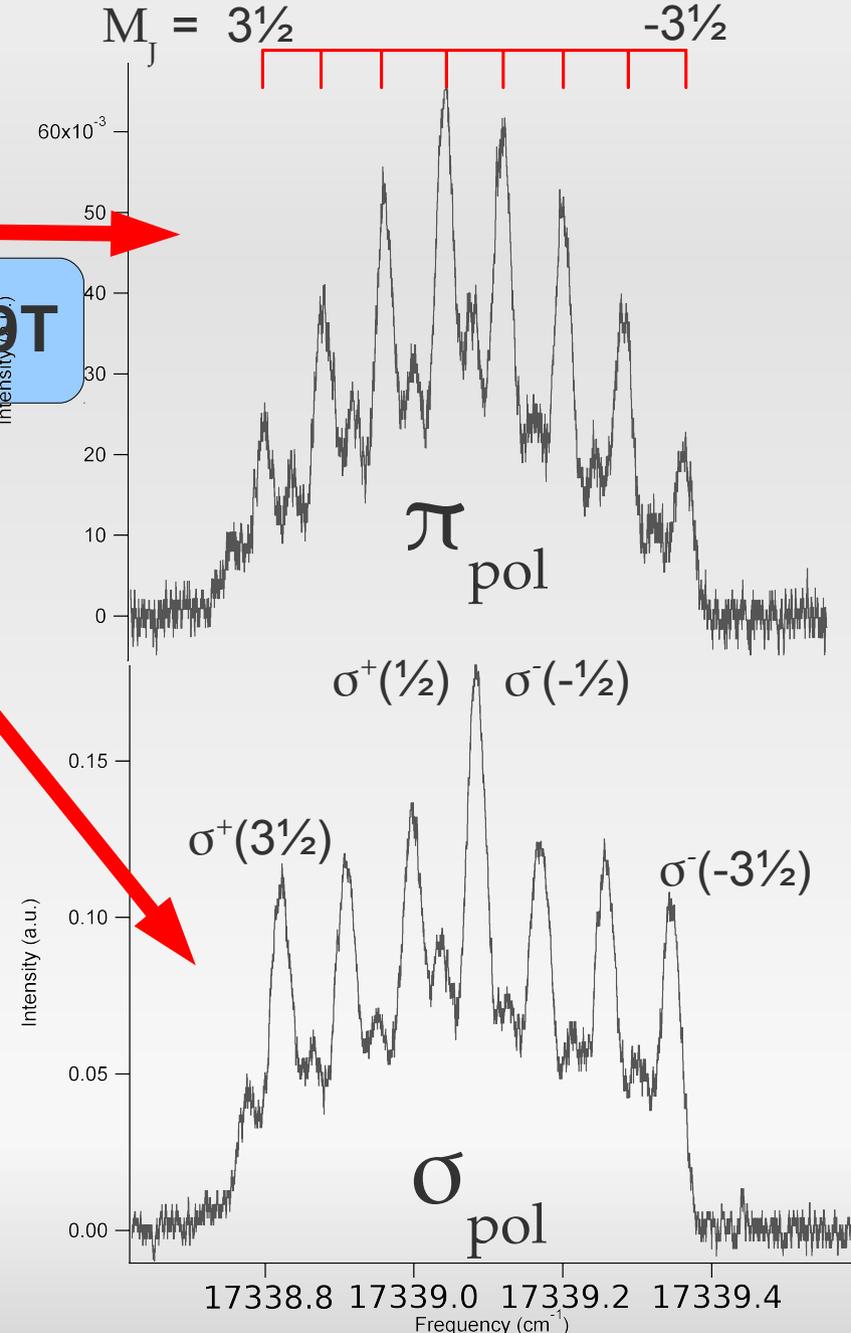
$R_e(3\frac{1}{2}) 0 - 0 I-X_1$

Zero field :

^{58}NiH 17339.090 cm⁻¹

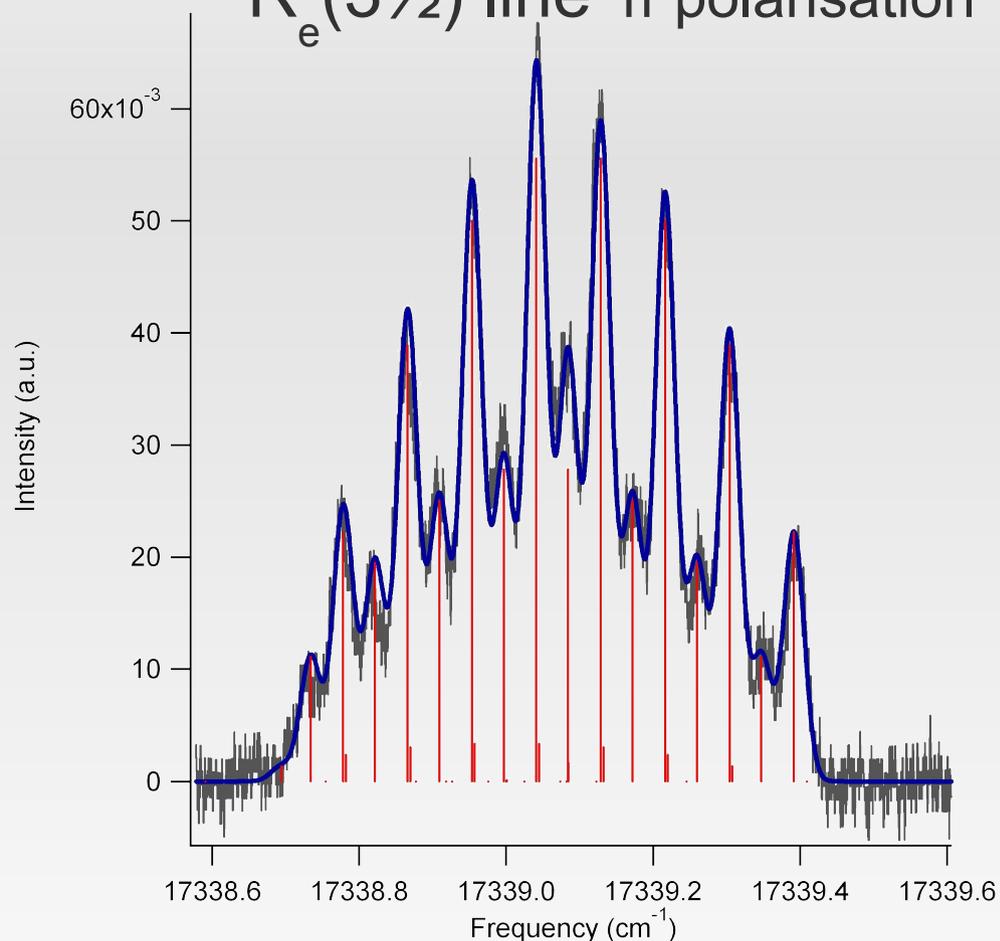
^{60}NiH 17339.050 cm⁻¹

B=0.59T

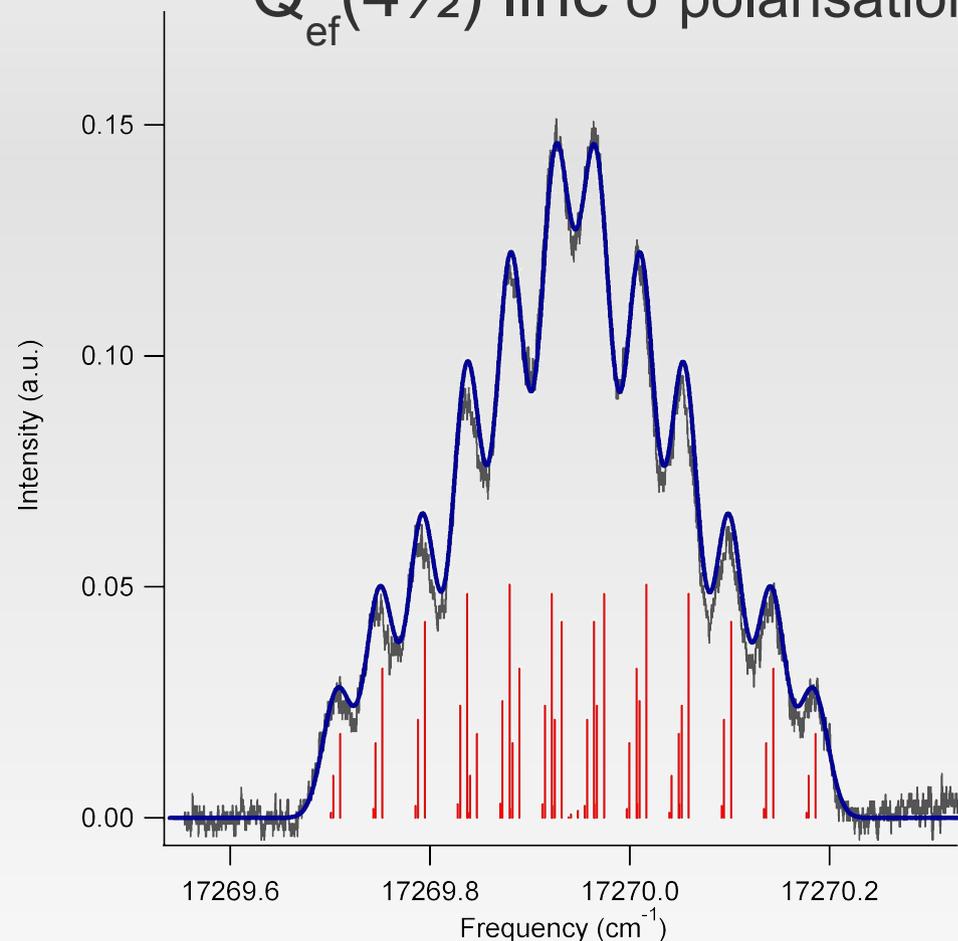


Intensity issues

$R_e(3\frac{1}{2})$ line π polarisation



$Q_{ef}(4\frac{1}{2})$ line σ polarisation



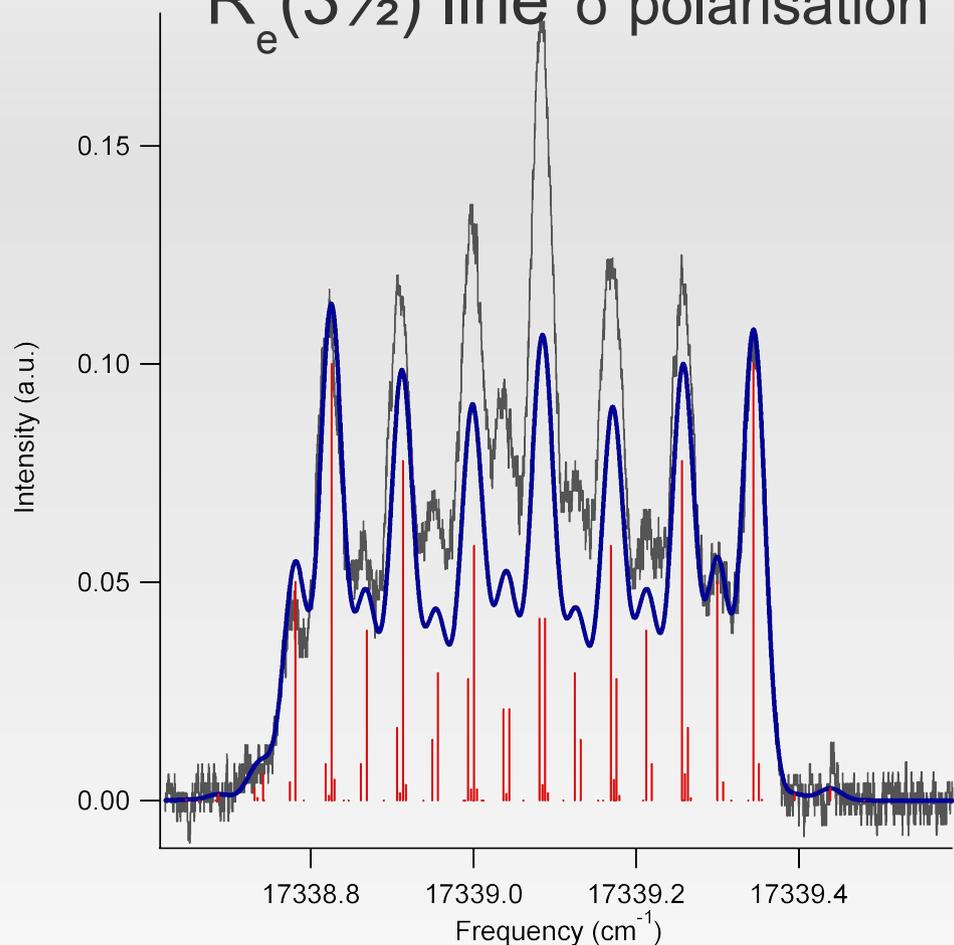
Model with Hönl-London formulae

MATCH

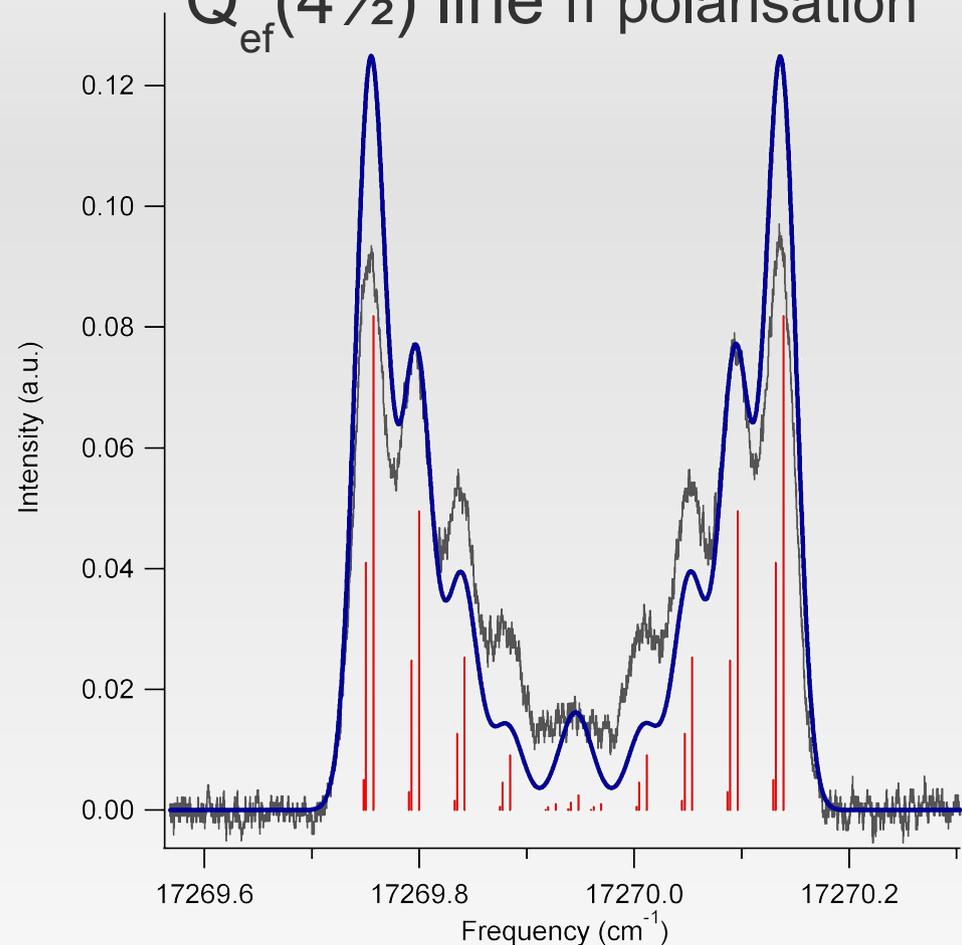
No isotopic problem
with Q line

Intensity issues

$R_e(3\frac{1}{2})$ line σ polarisation



$Q_{ef}(4\frac{1}{2})$ line π polarisation

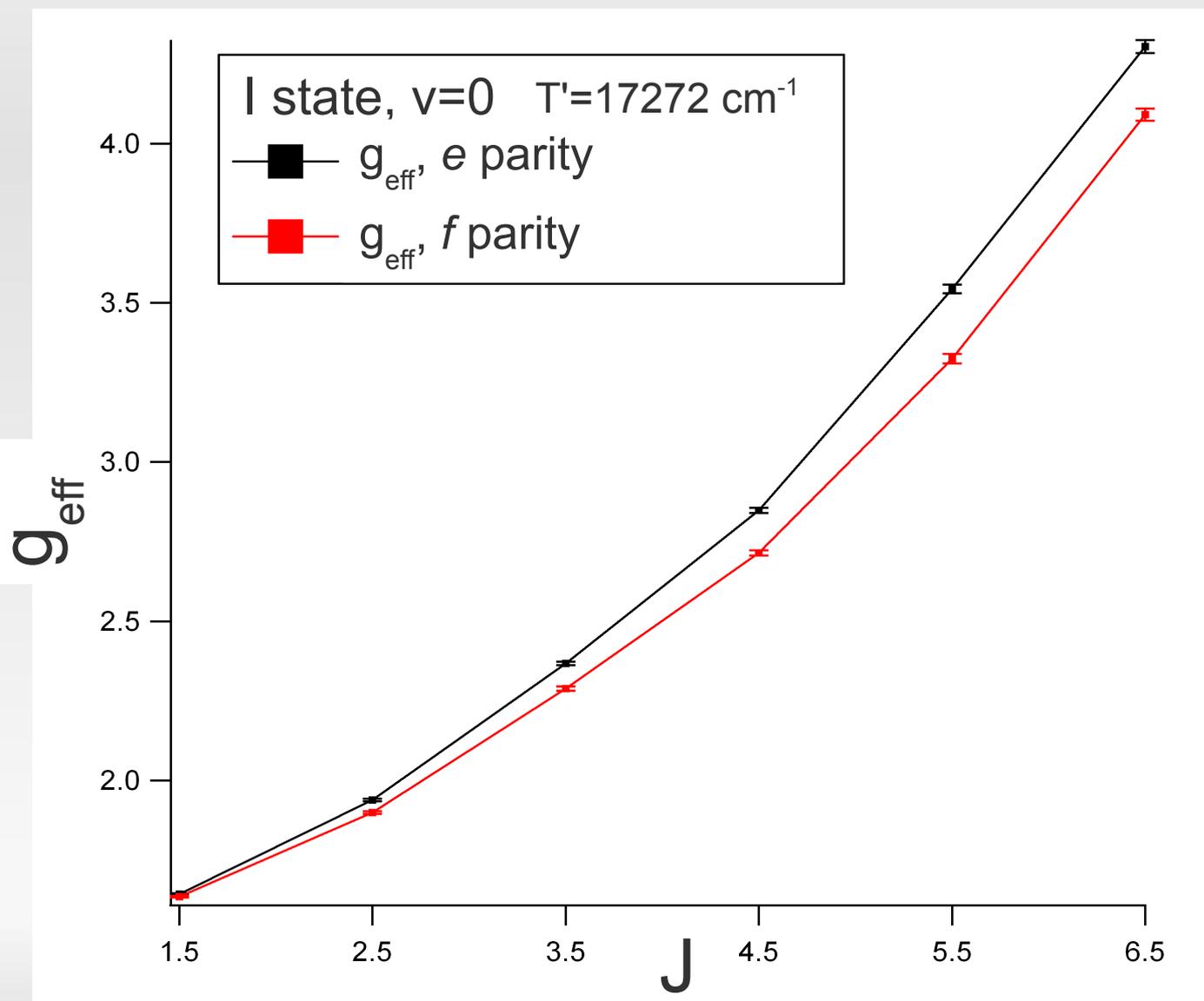


POOR INTENSITY MATCH
systematic discrepancy

Data Reduction

- Data have been recorded for I-X and E-X systems, $J < 8\frac{1}{2}$
- Experiments were performed for 11 magnetic field strengths
Deconvolution was necessary in some cases
- $g_{\text{eff}}(J)$ Landé factors are obtained from a least squares fit of
~ 2000 lines
- Typical uncertainties on $g_{\text{eff}}(J)$ are ~ 0.02
- Weighted RMS error for the fit : $6.58 \times 10^{-3} \text{ cm}^{-1}$

Landé factor of I($\Omega=3/2$)



- Rapid variation of g_{eff} in the rotational levels

- At low J :

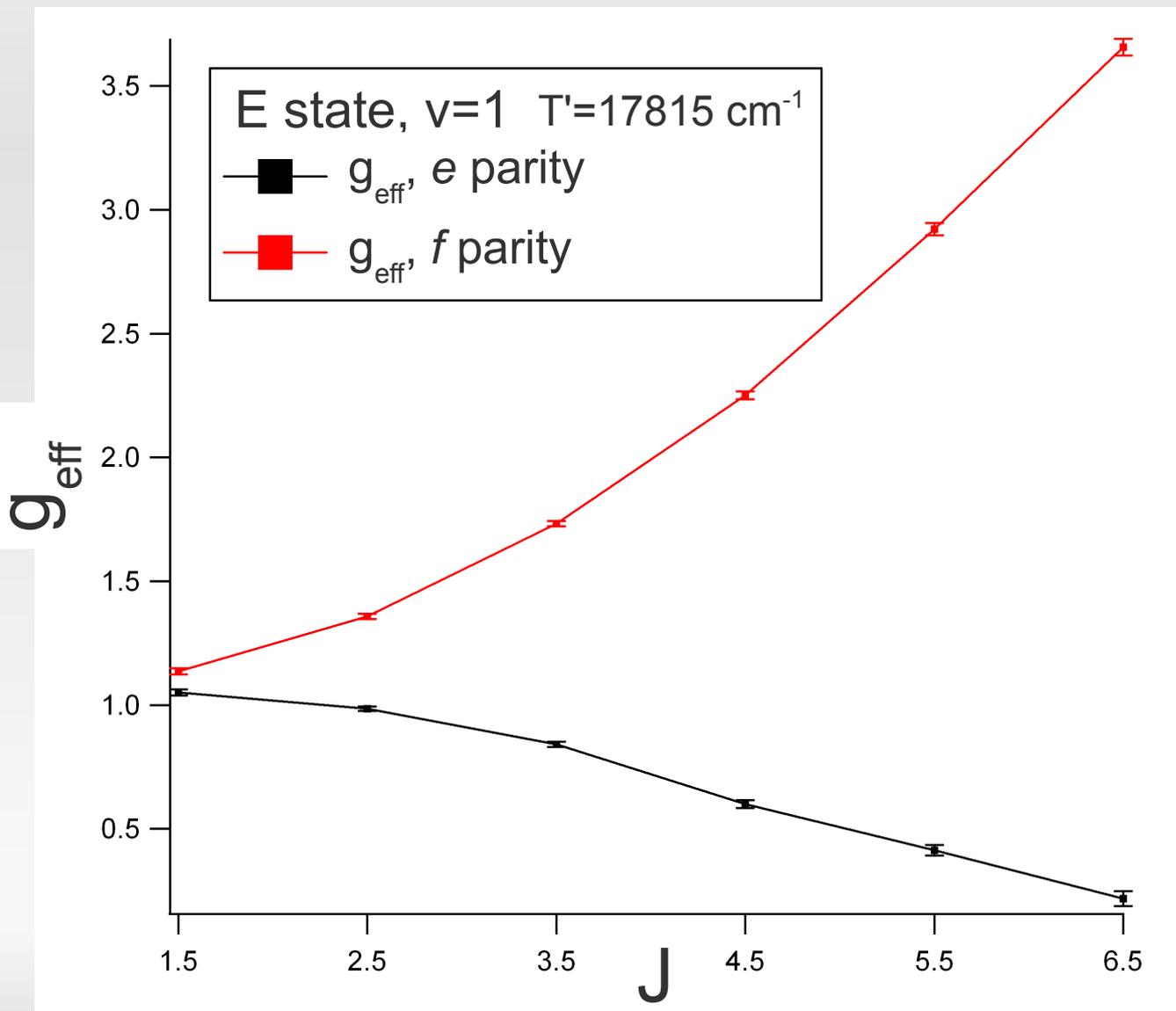
$$\Lambda_{\text{eff}} \approx g_s \Omega - g_{\text{eff}}$$

$$g_e(1/2) \approx g_f(1/2) \approx 1.64$$

$$\Lambda_{\text{eff}} \approx 1.36$$

probably Π state

Landé factor of E($\Omega=3/2$)



- Difference in variation of g_{eff} with e/f parity

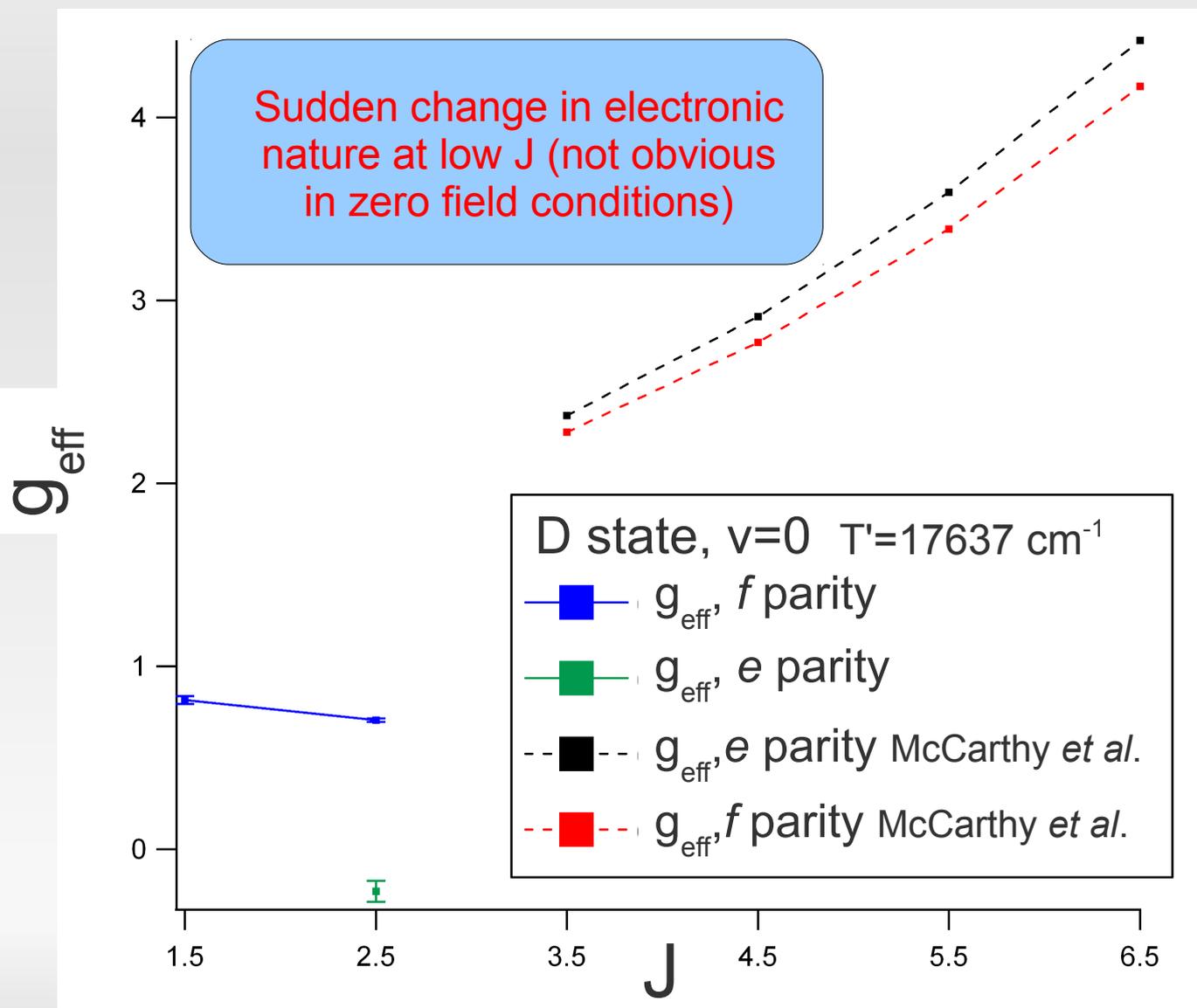
- Evidence for very different electronic character

$$g_e(1^{1/2}) \approx 1.052, g_f(1^{1/2}) \approx 1.136$$

$$\Lambda_{\text{eff}} \approx 2$$

probably Δ state

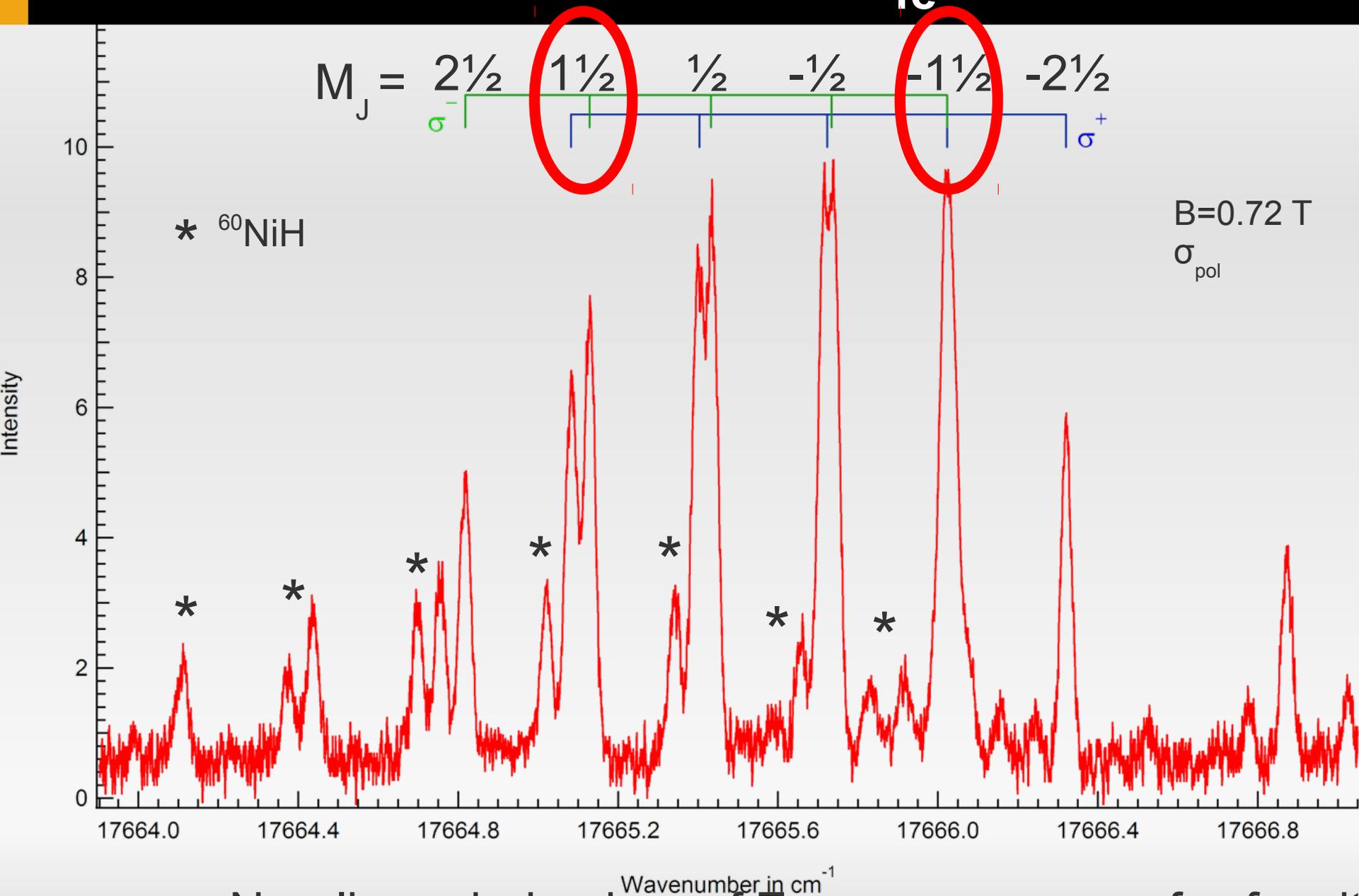
Landé factor of D($\Omega=3/2$)



Work in progress ...

McCarthy *et al.* JCP. **107**,
4179-4188, (1997)

Zeeman profile of $Q_{\text{Fe}}(2^{1/2}) : D \leftarrow X_1$



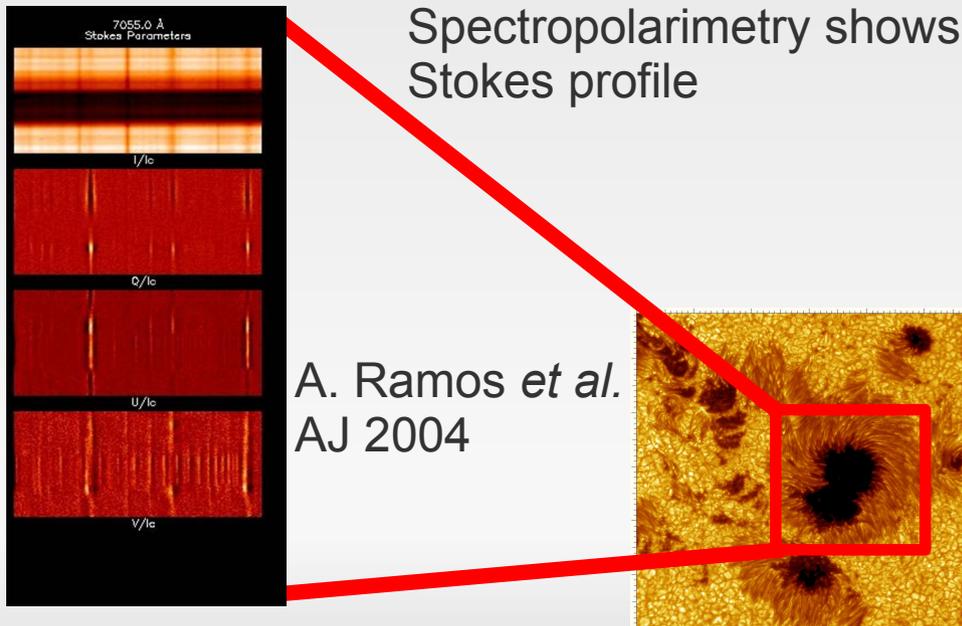
Non-linear behaviour of Zeeman response for f parity

Perspectives and Conclusions

Zeeman studies easily reveal changes in the electronic structure of these MH species.

Modeling the transitions is difficult for the excited states (predictions are unreliable)

For stellar spectropolarimetry, intensities are important too!



MH are possible probes of magnetic fields in sun- or star spots

Introduction to study FeH

ACKNOWLEDGEMENTS

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- Thank you for listening

