Physics 336K: Newtonian Dynamics

Homework 1: Solutions

When N point particles interact via gravity the two-particle interaction force is of the form

$$\mathbf{f}_{ij} = G \, m_i \, m_j \, \frac{(\mathbf{r}_j - \mathbf{r}_i)}{|\mathbf{r}_j - \mathbf{r}_i|^3},\tag{1}$$

whereas the two-particle potential energy is

$$U_{ij} = -\frac{G m_i m_j}{|\mathbf{r}_j - \mathbf{r}_1|}. (2)$$

The equation of motion of the ith particle then becomes

$$m_i \ddot{\mathbf{r}}_i = \sum_{j=1}^{j \neq i} G m_i m_j \frac{(\mathbf{r}_j - \mathbf{r}_i)}{|\mathbf{r}_j - \mathbf{r}_i|^3}.$$
 (3)

The total kinetic energy is

$$K = \frac{1}{2} \sum_{i} \dot{\mathbf{r}}_{i} \cdot \dot{\mathbf{r}}_{i}. \tag{4}$$

The total potential energy is

$$U = \frac{1}{2} \sum_{i,j}^{i \neq j} = -\frac{1}{2} \sum_{i,j}^{i \neq j} \frac{G \, m_i \, m_j}{|\mathbf{r}_j - \mathbf{r}_i|},\tag{5}$$

where the factor 1/2 is to compensate for double counting. We can form the scalar product of (3) with $\dot{\mathbf{r}}_i$, and then sum over all particles, to obtain

$$\sum_{i} m_{i} \dot{\mathbf{r}}_{i} \cdot \ddot{\mathbf{r}}_{i} = \sum_{i,j}^{j \neq i} G m_{i} m_{j} \frac{\dot{\mathbf{r}}_{i} \cdot (\mathbf{r}_{j} - \mathbf{r}_{i})}{|\mathbf{r}_{j} - \mathbf{r}_{i}|^{3}},$$
(6)

which is equivalent to

$$\frac{1}{2} \frac{d}{dt} \sum_{i} m_{i} \dot{\mathbf{r}}_{i} \cdot \dot{\mathbf{r}}_{i} = \sum_{i,j}^{j \neq i} G m_{i} m_{j} \frac{\dot{\mathbf{r}}_{i} \cdot (\mathbf{r}_{j} - \mathbf{r}_{i})}{|\mathbf{r}_{j} - \mathbf{r}_{i}|^{3}}, \tag{7}$$

or

$$\frac{dK}{dt} = \sum_{i,j} G m_i m_j \frac{\dot{\mathbf{r}}_i \cdot (\mathbf{r}_j - \mathbf{r}_i)}{|\mathbf{r}_j - \mathbf{r}_i|^3}, \tag{8}$$

where use has been made of (4). Swapping the dummy indices i and j on the right-hand side of the above expression yields

$$\frac{dK}{dt} = \sum_{j,i}^{i\neq j} G m_j m_i \frac{\dot{\mathbf{r}}_j \cdot (\mathbf{r}_i - \mathbf{r}_j)}{|\mathbf{r}_i - \mathbf{r}_j|^3} = -\sum_{i,j}^{j\neq i} G m_i m_j \frac{\dot{\mathbf{r}}_j \cdot (\mathbf{r}_j - \mathbf{r}_i)}{|\mathbf{r}_j - \mathbf{r}_i|^3}.$$
(9)

Forming half the sum of the previous two equations, we obtain

$$\frac{dK}{dt} = \sum_{i,j}^{j\neq i} G m_i m_j \frac{(\dot{\mathbf{r}}_i - \dot{\mathbf{r}}_j) \cdot (\mathbf{r}_j - \mathbf{r}_i)}{|\mathbf{r}_j - \mathbf{r}_i|^3} = \frac{1}{2} \frac{d}{dt} \sum_{i,j}^{j\neq i} \frac{G m_i m_j}{|\mathbf{r}_j - \mathbf{r}_i|} = -\frac{dU}{dt}, \quad (10)$$

where use has been made of (5). Hence,

$$\frac{d(K+U)}{dt} = 0, (11)$$

which implies that the total energy of the system, E = K + U, is a constant.

2. Suppose that

$$f(t x_1, t x_2, t x_3, \dots) = t^{\alpha} f(x_1, x_2, x_3, \dots)$$
 (12)

for all t. It follows that

$$\frac{d}{dt} f(t x_1, t x_2, t x_3, \dots) = \alpha t^{\alpha - 1} f(x_1, x_2, x_3, \dots).$$
 (13)

But, we can also write

$$\frac{d}{dt} f(t x_1, t x_2, t x_3, \dots) = \sum_{i} x_i \frac{\partial f(t x_1, t x_2, t x_3, \dots)}{\partial (t x_i)}.$$
 (14)

Hence,

$$\sum_{i} x_{i} \frac{\partial f(t x_{1}, t x_{2}, t x_{3}, \cdots)}{\partial (t x_{i})} = \alpha t^{\alpha - 1} f(x_{1}, x_{2}, x_{3}, \cdots).$$
 (15)

Setting t = 1, we get

$$\sum_{i} x_i \frac{\partial f(x_1, x_2, x_3, \cdots)}{\partial x_i} = \alpha f(x_1, x_2, x_3, \cdots).$$
 (16)

(2.3). Now,

$$\mathbf{f}_{ij} = k_i \, k_j \, \frac{(\mathbf{r}_j - \mathbf{r}_i)}{|\mathbf{r}_j - \mathbf{r}_i|^{n+1}} \tag{17}$$

is the appropriate expression for an attractive central two-particle force whose magnitude is $k_i k_j |\mathbf{r}_j - \mathbf{r}_i|^{-n}$. However,

$$k_i k_j \frac{(\mathbf{r}_j - \mathbf{r}_i)}{|\mathbf{r}_j - \mathbf{r}_i|^{n+1}} = -\frac{1}{n-1} \frac{\partial}{\partial \mathbf{r}_i} \left(\frac{k_i k_j}{|\mathbf{r}_j - \mathbf{r}_i|^{n-1}} \right), \tag{18}$$

provided that $n \neq 1$. In other words,

$$\mathbf{f}_{ij} = -\frac{\partial U_{ij}}{\partial \mathbf{r}_i},\tag{19}$$

where

$$U_{ij} = -\frac{1}{n-1} \frac{k_i \, k_j}{|\mathbf{r}_j - \mathbf{r}_i|^{n-1}} \tag{20}$$

is the two-particle potential energy. The total potential energy is

$$U = \frac{1}{2} \sum_{i,j}^{i \neq j} U_{ij}.$$
 (21)

However, if $\mathbf{r}_i \to t \, \mathbf{r}_i$ for all i then $U \to U/t^{n-1}$. Hence, U is a homogeneous function of degree 1-n. It follows, from the previous question, that

$$\sum_{i} \mathbf{r}_{i} \cdot \frac{\partial U}{\partial \mathbf{r}_{i}} = -(n-1) U. \tag{22}$$

Now, the equation of motion of the *i*th particle is

$$m_i \ddot{\mathbf{r}}_i = -\frac{\partial}{\partial \mathbf{r}_i} \sum_{j}^{j \neq i} U_{ij} = -\frac{1}{2} \frac{\partial}{\partial \mathbf{r}_i} \sum_{k,j}^{k \neq j} U_{kj}.$$
 (23)

The final step follows because only those terms in the sum for which either j = i or k = i survive partial differentiation with respect to \mathbf{r}_i , and also because $U_{ij} = U_{ji}$. Hence, using (21), the equation of motion becomes

$$m_i \ddot{\mathbf{r}}_i = -\frac{\partial U}{\partial \mathbf{r}_i}. (24)$$

Forming the scalar product with \mathbf{r}_i , and summing over all i, we get

$$\sum_{i} m_{i} \dot{\mathbf{r}}_{i} \cdot \ddot{\mathbf{r}}_{i} = -\sum_{i} \mathbf{r}_{i} \cdot \frac{\partial U}{\partial \mathbf{r}_{i}} = (n-1) U, \tag{25}$$

where use has been made of (22). Now,

$$\sum_{i} m_{i} \dot{\mathbf{r}}_{i} \cdot \ddot{\mathbf{r}}_{i} = \frac{1}{2} \frac{d^{2}}{dt^{2}} \sum_{i} m_{i} \mathbf{r}_{i} \cdot \mathbf{r}_{i} - \sum_{i} m_{i} \dot{\mathbf{r}}_{i} \cdot \dot{\mathbf{r}}_{i}, \tag{26}$$

which yields

$$\sum_{i} m_i \, \dot{\mathbf{r}}_i \cdot \ddot{\mathbf{r}}_i = \frac{1}{2} \, \ddot{I} - 2 \, K. \tag{27}$$

So, from (25),

$$\frac{1}{2}\ddot{I} = 2K + (n-1)U. \tag{28}$$

Now, in a steady state, $\ddot{I} = 0$, so that

$$U = -\frac{2K}{n-1}. (29)$$

In this case, the total energy becomes

$$E = K + U = \left(\frac{n-3}{n-1}\right)K. \tag{30}$$

We need E < 0 for a bound state (assuming n > 1, so that $U \to 0$ as the system disperses to infinity). Hence, we need

$$n < 3. (31)$$

In other words, there are no bound steady states for n > 3.

4. The virial equation, (28), for a gravitationally bound (i.e., n = 2) system, can be written

 $\frac{1}{2}\ddot{I} = 2K + U = 2(K + U) - U. \tag{32}$

However, the total energy, E = K + U, of an isolated system is constant in time. Hence, the above equation becomes

$$\ddot{I} = -2U + c, \tag{33}$$

where c=4E is a constant. Suppose that the system expands radially, in a uniform fashion, by some factor 1+u(t). If R is the outer radius of the star, and M its total mass, then $I \propto MR^2$ and $U \propto GM/R$. Hence, if $R \to R(1+u)$ at constant mass then $I \to I_0(1+u)^2$ and $U \to U_0/(1+u)$, where I_0 and U_0 are the unperturbed (i.e., u=0) moment of inertia and potential energy, respectively. It follows from (33) that, in the unperturbed (i.e., d/dt = 0) state,

$$0 = -2U_0 + c. (34)$$

Hence, (33) gives

$$\frac{d^2}{dt^2}[I_0(1+u)^2] = -2U_0/(1+u) + c. \tag{35}$$

Assuming that $|u| \ll 1$, so that the radial oscillations are of relatively small amplitude, we get

$$2I_0 \frac{d^2 u}{dt^2} \simeq -2U_0 + c + 2U_0 u = -2|U_0|u, \tag{36}$$

where use has been made of (34) and the fact that $U_0 < 0$. Hence, the radial oscillation equation takes the form

$$\frac{d^2u}{dt^2} = -\frac{|U_0|}{I_0}u. (37)$$

| Fitzpatrick, Chapter 2 |
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| Exercise 2.5 |
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| All the particle have the same change-to-massiatio: |
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| $\frac{2i}{mi} = \frac{2}{m}$ |
| |
| If the system is in a uniform magnetu field B, the equation of motion for the ith particle is |
| the equation of motion for the ith particle is |
| |
| $m_{i}\vec{r_{i}} = \sum_{j \neq i} \frac{k_{202j}}{ \vec{r_{i}} - \vec{r_{j}} ^{3}} (\vec{r_{i}} - \vec{r_{j}}) + 2i\vec{v_{o}} \times \vec{B}$ |
| |
| If we sum over i, the left side is the time derivative of the total momentum |
| of the total momentum |
| |
| Zimiri = du (Zimiri) = du P |
| |
| The resulting equation can be written |
| |
| $\frac{d\vec{r}}{dt} \vec{P} = \sum_{i} \frac{\langle 2, 2, r \rangle}{ \vec{r}_i - \vec{r}_j\rangle} (\vec{r}_i - \vec{r}_j) + \sum_{i} 2_i \vec{r}_i \times \vec{B}$ |
| |
| - 75 K2021 15 (- x) (- 57 , 17 =) 5 |

$$= \sum_{i \neq j} \frac{k_{2i} 2_{j}}{|\vec{r}_{i} - \vec{r}_{j}|} \frac{1}{1} (\vec{r}_{i} - \vec{r}_{j}) + (\vec{r}_{j} - \vec{r}_{j}) + (\vec{r}_{j} - \vec{r}_{j}) \times \vec{B}$$

$$= \frac{2}{m} \vec{p} \times \vec{B}$$

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This equation can be written
$$\frac{\partial}{\partial t} \vec{P} = -\Omega \vec{B} \times \vec{P},$$

$$\Omega = \frac{2}{m}B$$

It implies that P precesses around the direction B with angular foreguency I.

The total angular inomentum parallel to B is

The moment of inertia about an axis strough she origin parallel to Bu

$$I_{\parallel} = \sum_{i} m_{i} (\vec{r}_{i} \times \hat{B})^{2}$$

The time derivative of Ly is

$$\frac{d}{dt}L_{II} = \left[\frac{7}{2}m_{i}\left(\frac{1}{r_{o}} \times \frac{1}{r_{o}} + \frac{1}{r_{o}} \times \frac{1}{r_{o}} \right) \right] \cdot \hat{\mathbf{B}}$$

$$= \left(\sum_{i} \vec{r}_{i} \times (m_{i} \vec{r}_{i}) \right) \cdot \hat{\beta}$$

Inserting the equations of motion, the becomes $\frac{d}{dt} L_{\parallel} = \left(\sum_{i} \vec{r}_{i} \times (q_{i} \vec{r}_{i} \times \vec{B}) \right) \cdot \hat{B}$ $= \sum_{i} 2_{i} (\vec{r}_{i} \times \vec{B}) \times \vec{r}_{i} \cdot \hat{B}$

 $= -\frac{2}{m}B\sum_{i}m_{i}(\hat{r}_{i}\times\hat{B})\cdot(\hat{r}_{i}\times\hat{B})$

The time derivative of In is

 $\frac{d}{dt} I_{\parallel} = \sum_{\hat{r}} m_{\hat{r}} 2(\vec{r}_{o} \times \hat{B}) \cdot (\vec{r}_{c} \times \vec{B})$

Thus the time derivative of Ly + & RIH is

g (L1+ 12 II)

 $= -\frac{2}{m} B Z m_o(\vec{r}_o \times \hat{B}) \circ (\vec{r}_o \times \hat{B})$

+ 1 (2 B) 2 Zm, (r, xB). (r, xB)

=0

Thus L1 + 1 DI, is a constant of the motion.