

Problem 1.

A large number N of identical bosonic atoms, all in the same spin state, are confined inside a cube of length L . The orbitals are labelled by quantum numbers n_x, n_y, n_z with values $1, 2, 3, \dots$. The orbital energies are $\varepsilon_{n_x, n_y, n_z} = (n_x^2 + n_y^2 + n_z^2)h^2/(8mL^2)$.

(A) What is the lowest-energy orbital and what is its energy ε_0 ?

$$n_x = 1, n_y = 1, n_z = 1$$

$$\varepsilon_0 = 3 \frac{h^2}{8mL^2}$$

(B) Describe the ground state of the system of N atoms. What is its total energy U ?

all N atoms in the lowest energy orbital: $n_x=1, n_y=1, n_z=1$

$$U = N\varepsilon_0 = N \frac{3h^2}{8mL^2}$$

Suppose the atoms are in equilibrium at temperature T and chemical potential μ .

(C) What is the average number N_0 of atoms in the lowest-energy orbital.

$$N_0 = \frac{1}{e^{\beta(\varepsilon_0 - \mu)} - 1}, \quad \beta = \frac{1}{kT}$$

(D) Express the average total number \bar{N} of atoms as a sum over quantum numbers (with definite upper and lower limits).

$$\bar{N} = \sum_{n_x=1}^{\infty} \sum_{n_y=1}^{\infty} \sum_{n_z=1}^{\infty} \frac{1}{e^{\beta(\varepsilon_{n_x, n_y, n_z} - \mu)} - 1}$$

(E) Give the best possible upper bound on the chemical potential μ . What is this upper bound in the thermodynamic limit?

$$\mu < \varepsilon_0$$

thermodynamic limit: $L \rightarrow \infty \Rightarrow \varepsilon_0 \rightarrow 0$

$$\mu \leq 0$$

(F) Define the thermodynamic limit in terms of the variables N , L , and T .

$$N \rightarrow \infty$$

$$L \rightarrow \infty \text{ with } \frac{N}{L^3} \text{ fixed}$$

$$T \text{ fixed}$$

Now focus on the thermodynamic limit for this system of N identical bosonic atoms, all in the same spin state, confined to a volume $V = L^3$.

(G) Express the orbital energy ϵ as a function of the momentum.

$$\epsilon = \frac{p^2}{2m}$$

(H) Determine the critical temperature T_c for Bose-Einstein condensation as a function of the number density N/V . (The integrals at the bottom of the page may be useful.)

$$N = \frac{V}{h^3} \int d^3p \frac{1}{e^{\beta p^2/2m} - 1} = \frac{V}{h^3} \zeta\left(\frac{3}{2}\right) (2\pi m kT)^{3/2}$$

$$kT_c = \frac{h^2}{2\pi m} \left(\frac{1}{\zeta(3/2)} \frac{N}{V} \right)^{2/3}$$

(I) For temperatures $T > T_c$, express N and the total energy U in terms of momentum integrals.

$$N = \frac{V}{h^3} \int d^3p \frac{1}{e^{\beta(\epsilon - \mu)} - 1}$$

$$\epsilon = \frac{p^2}{2m}$$

$$U = \frac{V}{h^3} \int d^3p \frac{\epsilon}{e^{\beta(\epsilon - \mu)} - 1}$$

(J) For temperatures $T < T_c$, express N and U in terms of momentum integrals (with μ eliminated in favor of N_0).

$$N = N_0 + \frac{V}{h^3} \int d^3p \frac{1}{e^{\beta\epsilon} - 1}$$

$$\epsilon = \frac{p^2}{2m}$$

$$U = N_0 \epsilon_0 + \frac{V}{h^3} \int d^3p \frac{\epsilon}{e^{\beta\epsilon} - 1}$$

(K) For $T < T_c$, determine N_0 as a function of T . Express the condensate fraction N_0/N as a function of T/T_c only.

$$N = N_0 + \frac{V}{h^3} \zeta\left(\frac{3}{2}\right) (2\pi m kT)^{3/2} \quad N_0 = N - \frac{V}{h^3} \zeta\left(\frac{3}{2}\right) (2\pi m kT)^{3/2}$$

$$\text{divide by } N = \frac{V}{h^3} (2\pi m kT_c)^{3/2}$$

$$\frac{N_0}{N} = 1 - \left(\frac{T}{T_c}\right)^{3/2}$$

(L) For $T < T_c$, determine U as a function of T .

$$U = N_0 \epsilon_0 + \frac{V}{h^3} \frac{3}{2\pi} \zeta\left(\frac{5}{2}\right) (2\pi m kT)^{5/2}$$

$$\int d^3p \frac{1}{\exp(p^2/2mkT) - 1} = \zeta\left(\frac{3}{2}\right) (2\pi m kT)^{3/2}, \quad \int d^3p \frac{p^2}{\exp(p^2/2mkT) - 1} = \frac{3}{2\pi} \zeta\left(\frac{5}{2}\right) (2\pi m kT)^{5/2}$$

Problem 2.

The differential power in electromagnetic radiation emitted by a black body of surface area A and temperature T is

$$dP = A \frac{2\pi h}{c^2} \frac{f^3}{\exp(hf/kT) - 1} df.$$

(A) Stefan's Law states that the total power radiated by the black body is $P = \sigma T^4 A$, where σ is a constant. Express σ in terms of physical constants and an integral over a dimensionless variable x .

$$P = A \frac{2\pi h}{c^2} \int_0^\infty df \frac{f^3}{e^{hf/kT} - 1} = A \frac{2\pi h}{c^2} \left(\frac{kT}{h}\right)^4 \int_0^\infty dx \frac{x^3}{e^x - 1}$$

$$\sigma = \frac{2\pi k^4}{c^2 h^3} \int_0^\infty dx \frac{x^3}{e^x - 1}$$

(B) Wien's Law states that the frequency f_{\max} at which the maximum power is radiated is proportional to T . The law can be expressed as $f_{\max} = xkT/h$, where x is a numerical constant. Derive an equation for x .

$$\text{maximum power} \Rightarrow \frac{d}{df} \frac{2\pi h}{c^2} \frac{f^3}{e^{hf/kT} - 1} = 0$$

$$\frac{3f^2}{e^{hf/kT} - 1} - \frac{f^3}{(e^{hf/kT} - 1)^2} e^{hf/kT} \frac{h}{kT} = 0$$

$$\text{multiply by } (h/kT)^2: \frac{3x^2}{e^x - 1} - \frac{x^3 e^x}{(e^x - 1)^2} = 0$$

The electromagnetic energy inside a cavity of volume V in a conductor with temperature T can be expressed as an integral over the photon energy ϵ :

$$U = V \frac{8\pi}{h^3 c^3} \int_0^\infty \frac{\epsilon^3}{\exp(\epsilon/kT) - 1} d\epsilon.$$

(C) State as precisely as possible the relation between the shape of the frequency spectrum emitted by a black body of temperature T and the shape of the photon energy distribution in a spherical cavity at the same temperature T .

frequency spectrum has same shape as energy distribution if we set $f = \epsilon/h$

(D) Suppose the black body is a sphere of radius R and the cavity is also a sphere of radius R . State as precisely as possible the relation between the total power P emitted by the black body and the total energy U in the cavity.

$$P = 4\pi R^2 \cdot \frac{2\pi}{c^2 h^3} (kT)^4 \int_0^\infty dx \frac{x^3}{e^x - 1}$$

$$U = \frac{4}{3}\pi R^3 \cdot \frac{8\pi}{h^3 c^3} (kT)^4 \int_0^\infty dx \frac{x^3}{e^x - 1}$$

$$P = \frac{3c}{4R} U$$

(E) Suppose a small cylindrical hole of radius r is drilled into the conductor so that electromagnetic radiation can leak out from the cavity through the hole. What is the power radiated from the hole?

$$P = \sigma T^4 \cdot (\pi r^2)$$

In a 2-dimensional universe, electromagnetic (EM) waves have only one transverse polarization mode. Consider a 2-dimensional cavity in a conductor in such a universe.

If the cavity is a square of area L^2 , the normal modes of standing EM waves can be labelled by integers $n_x, n_y = 1, 2, 3, \dots$. The amplitude $a(t)$ of a normal mode oscillates with frequency $f = \sqrt{n_x^2 + n_y^2} c / 2L$.

(E) Suppose the amplitude of each normal mode can be treated as a classical harmonic oscillator that is in equilibrium at temperature T . Use the equipartition theorem to determine the average energy \bar{E} in a single standing wave of frequency f .

1 oscillator \Rightarrow 2 quadratic degrees of freedom
equipartition theorem: $\bar{E} = 2 (\frac{1}{2} kT) = kT$

(F) Use the equipartition theorem to determine the average total energy \bar{U} in all the EM standing waves. Is there anything unphysical about the result for \bar{U} ?

$$\bar{U} = \sum_{n_x=1}^{\infty} \sum_{n_y=1}^{\infty} kT = \infty \cdot kT$$

ultraviolet catastrophe: \bar{U} is divergent

(G) Suppose the energy in a oscillator of frequency f is quantized with quantum hf : $E = nhf$, $n = 0, 1, 2, \dots$. If the oscillator is in equilibrium at temperature T , what is its average energy \bar{E} ?

$$\bar{E} = \frac{\sum_{n=0}^{\infty} nhf e^{-\beta n h f}}{\sum_{n=0}^{\infty} e^{-\beta n h f}} = \frac{hf}{e^{\beta h f} - 1}$$

(H) Under the same assumptions, what is the average total energy \bar{U} in all the EM standing waves?

$$\bar{U} = \sum_{n_x=1}^{\infty} \sum_{n_y=1}^{\infty} \frac{h \sqrt{n_x^2 + n_y^2} c / 2L}{e^{\beta h \sqrt{n_x^2 + n_y^2} c / 2L} - 1}$$

Now consider this system in the thermodynamic limit. Suppose the electromagnetic energy is in the form of photons, which are massless bosons with energy-momentum relation $\epsilon = \sqrt{p_x^2 + p_y^2} c$ and chemical potential 0.

(I) What is the average energy per area?

$$U = \frac{A}{h^2} \int d^2 p \frac{\epsilon}{e^{\beta \epsilon} - 1}, \quad \epsilon = pc \quad \frac{U}{A} = \frac{2\pi}{h^2} \int_0^{\infty} p dp \frac{pc}{e^{\beta pc} - 1}$$

(J) What is the average number of photons per area?

$$N_{\text{photons}} = \frac{A}{h^2} \int d^2 p \frac{1}{e^{\beta \epsilon} - 1}, \quad \epsilon = pc \quad \frac{N_{\text{photons}}}{A} = \frac{2\pi}{h^2} \int_0^{\infty} p dp \frac{1}{e^{\beta pc} - 1}$$

Problem 3.

A single orbital with orbital energy ϵ is in both thermal and chemical equilibrium with a reservoir of temperature T and chemical potential μ .

Suppose the only possible occupation numbers n of the orbital are 0 and 1.

(A) What is the ground state partition function Z_1 for the single orbital?

$$Z_1 = \sum_{n=0}^1 e^{-\beta(\epsilon-\mu)n} = 1 + e^{-\beta(\epsilon-\mu)}$$

(B) What are the probabilities $P(0)$ and $P(1)$ for the orbital to have occupation numbers 0 and 1?

$$P(0) = \frac{1}{1 + e^{-\beta(\epsilon-\mu)}} \quad P(1) = \frac{e^{-\beta(\epsilon-\mu)}}{1 + e^{-\beta(\epsilon-\mu)}}$$

(C) What is the average occupation number \bar{n} ?

$$\bar{n} = \frac{1}{e^{\beta(\epsilon-\mu)} + 1}$$

(D) What is the average energy \bar{E} of the orbital?

$$\bar{E} = \frac{\epsilon}{e^{\beta(\epsilon-\mu)} + 1}$$

Suppose the occupation number n of the orbital can be any nonnegative integer: 0,1,2,3,...

(E) What is the ground state partition function Z_1 for the single orbital?

$$Z_1 = \sum_{n=0}^{\infty} e^{-\beta(\epsilon-\mu)n} = \frac{1}{1 - e^{-\beta(\epsilon-\mu)}}$$

(F) What is the probability distribution $P(n)$ for the occupation number n ?

$$P(n) = \frac{1}{Z_1} e^{-\beta(\epsilon-\mu)n}$$

(G) What is the average occupation number \bar{n} ?

$$\bar{n} = \frac{1}{e^{\beta(\epsilon-\mu)} - 1}$$

(H) What is the energy E in the orbital if its occupation number is n ? What is the average energy \bar{E} of the orbital?

$$E = n\epsilon$$

$$\bar{E} = \frac{\epsilon}{e^{\beta(\epsilon-\mu)} - 1}$$

A system consists of many identical fermions, all in the same spin state, with temperature T and chemical potential μ . The fermions have K distinct orbitals labelled by $k = 1, 2, \dots, K$ with energies $\varepsilon_1, \varepsilon_2, \dots, \varepsilon_K$. Express each of the following in terms of a sum over the orbitals.

(I) the average number \bar{N} of fermions.

$$\bar{N} = \sum_{k=1}^K \frac{1}{e^{\beta(\varepsilon_k - \mu)} + 1}$$

(J) the average total energy \bar{U} .

$$\bar{U} = \sum_{k=1}^K \frac{\varepsilon_k}{e^{\beta(\varepsilon_k - \mu)} + 1}$$

(K) the logarithm of the grand partition function \mathcal{Z} .

$$\log \mathcal{Z} = \sum_{k=1}^K \log (1 + e^{-\beta(\varepsilon_k - \mu)})$$

Suppose the system consists of exactly N identical noninteracting fermions, ^{in a volume V} all in the same spin state, at temperature T . In the thermodynamic limit, it can be described using the grand canonical ensemble with temperature T and some chemical potential μ .

(K) Express the orbital energy as a function of the momentum of the fermion. Express the sum over orbitals as a momentum integral.

$$\varepsilon = \frac{p^2}{2m}$$

$$\sum_{\text{orbital}} = \frac{V}{h^3} \int d^3p$$

(L) Write down the equation whose solution determines the chemical potential μ . (It should involve a sum over orbitals.)

$$N = \bar{N}$$

$$N = \frac{V}{h^3} \int d^3p \frac{1}{e^{\beta(p^2/2m)} + 1}$$