

Last time | We constructed the Lagrangian density for the electromagnetic field. Combined with the interaction term for charges, we had:

$$\mathcal{L} = -\frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu} - \frac{1}{c} J_{\mu} A^{\mu}$$

We derived the EOM for this Lagrangian density:

Maxwell equations:

$$\partial_{\mu} F^{\mu\nu} = \frac{4\pi}{c} J^{\nu}$$

$$\partial_{\mu} \tilde{F}^{\mu\nu} = 0$$

↪ EOM proper

↪ always true for $\tilde{F}^{\mu\nu} = \frac{1}{2} \epsilon^{\mu\nu\rho\sigma} F_{\rho\sigma}$

Maxwell Equations

$$\nabla = 0$$

in the 1st eq'n:

$$\vec{\nabla} \cdot \vec{E} = 4\pi \rho$$

Coulomb's Law

(Gauss's Law)

$$\nabla = i$$

— — — :

$$\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{J} + \frac{1}{c} \frac{\partial \vec{E}}{\partial t}$$

Ampere-

-Maxwell Law

Maxwell Equations

Let us rewrite the obtained equations for $F_{\mu\nu}$ & $\tilde{F}_{\mu\nu}$ as equations for \vec{E} & \vec{B} fields.

Start with $\partial_\mu F^{\mu\nu} = \frac{4\pi}{c} J^\nu$.

	$F^{0i} = -E^i$
	$F^{ij} = -\epsilon^{ijk} B^k$

$\nu=0$ $\partial_\mu F^{\mu 0} = \frac{4\pi}{c} J^0$

\Rightarrow as $J^0 = c\rho$ and $\partial_\mu F^{\mu 0} = \partial_i F^{i0} = \partial_i E^i = \vec{\nabla} \cdot \vec{E}$
(since $\partial_i = \nabla^i$) \Rightarrow get

$\vec{\nabla} \cdot \vec{E} = 4\pi\rho$

Coulomb's Law
ca. 1785 (aka Gauss's Law)

$\nu=i$ $\partial_\mu F^{\mu i} = \frac{4\pi}{c} J^i$, $i=1,2,3$

$\Rightarrow \partial_0 F^{0i} + \partial_j F^{ji} = \frac{4\pi}{c} J^i$

$-\partial_0 E^i - \partial_j \epsilon^{jik} B^k = \frac{4\pi}{c} J^i$

$\underbrace{\epsilon^{ijk} \nabla^j B^k}_{(\vec{\nabla} \times \vec{B})^i} = \partial_0 E^i + \frac{4\pi}{c} J^i$

\Rightarrow $\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{J} + \frac{1}{c} \frac{\partial \vec{E}}{\partial t}$

1826 1861
Ampere-Maxwell
Law

$\tilde{F}^{\mu\nu}$ has the following components:

$$\tilde{F}^{0i} = \frac{1}{2} \varepsilon^{0ijk} F_{jk} = \frac{1}{2} \varepsilon^{ijk} (-) \varepsilon^{jkl} B^l$$

\Rightarrow use $\varepsilon^{ijk} \varepsilon^{i'j'k'} = \delta^{jj'} \delta^{kk'} - \delta^{jk'} \delta^{kj'}$ to write

$$\varepsilon^{ijk} \varepsilon^{ij'k'} = 2 \delta^{kk'} \Rightarrow \tilde{F}^{0i} = -\frac{1}{2} \cdot 2 \delta^{il} B^l = -B^i$$

$$\Rightarrow \boxed{\tilde{F}^{0i} = -B^i}$$

$$\tilde{F}^{ij} = \frac{1}{2} \varepsilon^{ij\rho\sigma} F_{\rho\sigma} = \varepsilon^{ij0k} F_{0k} = -\varepsilon^{ijk} F^{0k} =$$

$$= -\varepsilon^{ijk} (-E^k) = \varepsilon^{ijk} E^k \Rightarrow \boxed{\tilde{F}^{ij} = \varepsilon^{ijk} E^k}$$

\Rightarrow let's study $\partial_\mu \tilde{F}^{\mu\nu} = 0$

$$\boxed{V=0} \quad \partial_\mu \tilde{F}^{\mu 0} = 0 \Rightarrow \partial_i \tilde{F}^{i0} = 0 \Rightarrow \partial_i B^i = 0$$

$\Rightarrow \boxed{\vec{\nabla} \cdot \vec{B} = 0}$ \sim law stating the absence of magnetic monopoles

$$\boxed{V=i} \quad \partial_\mu \tilde{F}^{\mu i} = 0 \Rightarrow \partial_0 \tilde{F}^{0i} + \partial_j \tilde{F}^{ji} = 0$$

$$\Rightarrow -\partial_0 B^i + \partial_j \varepsilon^{jik} E^k = 0 \Rightarrow \underbrace{\varepsilon^{jik} \nabla^j E^k}_{(\vec{\nabla} \times \vec{E})^i} + \underbrace{\partial_0 B^i}_{\frac{1}{c} \frac{\partial B^i}{\partial t}} = 0$$

$$\Rightarrow \boxed{\vec{\nabla} \times \vec{E} + \frac{1}{c} \frac{\partial \vec{B}}{\partial t} = 0} \quad \underline{\text{Faraday's Law, ca. 1831}}$$

To summarize, in 3-vector notation we have (62)
 the following Maxwell equations (ca. 1865):

$\vec{\nabla} \cdot \vec{E} = 4\pi\rho$ $\vec{\nabla} \cdot \vec{B} = 0$	$\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{J} + \frac{1}{c} \frac{\partial \vec{E}}{\partial t}$ $\vec{\nabla} \times \vec{E} + \frac{1}{c} \frac{\partial \vec{B}}{\partial t} = 0$	Gaussian Units
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$\vec{\nabla} \cdot \vec{E} = 4\pi\rho \sim$ Coulomb's Law (Gauss's Law)

$\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{J} + \frac{1}{c} \frac{\partial \vec{E}}{\partial t} \sim$ Ampere-Maxwell Law

$\vec{\nabla} \cdot \vec{B} = 0 \sim$ no monopoles

$\vec{\nabla} \times \vec{E} + \frac{1}{c} \frac{\partial \vec{B}}{\partial t} = 0 \sim$ Faraday's Law

For future use, let us rewrite Maxwell equations in SI (Le Système International d'unités):

(see page 781 of Jackson for a useful table)

$\vec{\nabla} \cdot \vec{E} = \rho / \epsilon_0$ $\vec{\nabla} \cdot \vec{B} = 0$	$\vec{\nabla} \times \vec{B} = \mu_0 \vec{J} + \frac{1}{c^2} \frac{\partial \vec{E}}{\partial t}$ $\vec{\nabla} \times \vec{E} + \frac{\mu_0}{4} \frac{\partial \vec{B}}{\partial t} = 0$	Maxwell equations, SI units
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$\frac{1}{\mu_0 \epsilon_0} = c^2$, $\epsilon_0 \approx 8.854 \times 10^{-12} \text{ F/m}$, $1 \text{ Farad} = 1 \frac{\text{Coulomb}^2}{\text{N}\cdot\text{m}}$

$\mu_0 = 4\pi \times 10^{-7} \frac{\text{Newtons}}{\text{Ampere}^2}$, $1 \text{ Ampere} = 1 \frac{\text{Coulomb}}{\text{Second}}$

We have obtained Maxwell equations by assuming

- Lorentz-invariance of physics
 - existence of a vector field $A^\mu(x)$
 - gauge-invariance of physics
(charge conservation)
- ⊕ Superposition principle for $A^\mu(x)$.

Conservation Laws and Energy-Momentum (64)

Tensor

Noether's theorem | Every continuous symmetry of the action corresponds to a conservation law.

Space-time is homogeneous \Rightarrow physics is invariant under space-time translations

$$x^M \rightarrow x^M - a^M, \quad a^M = \text{const}$$

\Rightarrow energy & momentum are conserved

\Rightarrow let's find energy and momentum of an E & M - interacting system

Conservation Laws:

(65)

Energy conservation in mechanics:

$\mathcal{S} = \int dt L(q_i, \dot{q}_i, t) \Rightarrow$ EL equations read

$$\frac{\partial L}{\partial q_i} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right) = 0.$$

$$\text{Now, } \frac{dL}{dt} = \sum_i \frac{\partial L}{\partial q_i} \dot{q}_i + \sum_i \frac{\partial L}{\partial \dot{q}_i} \ddot{q}_i + \frac{\partial L}{\partial t} = (\text{EL eqns})$$

$$= \sum_i \left[\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right) \dot{q}_i + \frac{\partial L}{\partial \dot{q}_i} \ddot{q}_i \right] + \frac{\partial L}{\partial t} =$$

$$= \frac{d}{dt} \left(\sum_i \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i \right) + \frac{\partial L}{\partial t}$$

$$\Rightarrow \frac{d}{dt} \left(\underbrace{\sum_i \frac{\partial L}{\partial \dot{q}_i} \dot{q}_i - L}_{\text{''}} \right) = - \frac{\partial L}{\partial t}$$

''
E energy of the system (cf. Hamiltonian)

$$\text{If } L = L(q_i, \dot{q}_i) \Rightarrow \frac{\partial L}{\partial t} = 0 \Rightarrow \boxed{\frac{dE}{dt} = 0}$$

\Rightarrow energy is conserved!

Imagine a ^{field} theory with

$$\mathcal{L} = \mathcal{L}(\phi, \partial_\mu \phi) \quad (\text{no } x^\mu\text{-dependence in } \mathcal{L})$$

Imagine an infinitesimal space-time shift:

$$x^\mu \rightarrow x^\mu - \delta a^\mu = x'^\mu \Rightarrow x^\mu = x'^\mu + \delta a^\mu$$

ϕ is inv. $\phi'(x') = \phi(x)$

$$\Rightarrow \phi(x) \rightarrow \phi(x'^\mu + \delta a^\mu) \approx \phi(x'^\mu) + \overbrace{\delta a^\mu \partial_\mu \phi(x')}^{\delta \phi}$$

$$\delta \mathcal{L} = \frac{\delta \mathcal{L}}{\delta \phi} \delta \phi + \left(\frac{\delta \mathcal{L}}{\delta(\partial_\mu \phi)} \delta(\partial_\mu \phi) \right) = \left[\frac{\delta \mathcal{L}}{\delta \phi} - \partial_\mu \left(\frac{\delta \mathcal{L}}{\delta(\partial_\mu \phi)} \right) \right] \delta \phi$$

" (EOM)

$$+ \partial_\mu \left(\frac{\delta \mathcal{L}}{\delta(\partial_\mu \phi)} \delta \phi \right) = \partial_\mu \left(\frac{\delta \mathcal{L}}{\delta(\partial_\mu \phi)} \delta \phi \right)$$

$$\Rightarrow \delta \mathcal{L} = \partial_\mu \left(\frac{\delta \mathcal{L}}{\delta(\partial_\mu \phi)} \delta a^\nu \partial_\nu \phi \right)$$

On the other hand \mathcal{L} is scalar $\Rightarrow \mathcal{L} = \mathcal{L}(x) \Rightarrow$

$$\mathcal{L} \rightarrow \mathcal{L}' = \mathcal{L} + \delta a^\mu \partial_\mu \mathcal{L} \Rightarrow \delta \mathcal{L} = \delta a^\nu \partial_\nu (\delta \mathcal{L})$$

Equating two $\delta \mathcal{L}$'s we get

$$\delta a^\nu \partial_\nu \left[\frac{\delta \mathcal{L}}{\delta(\partial_\mu \phi)} \partial_\mu \phi - \delta \mathcal{L} \right] = 0 \quad \text{for any } \delta a^\nu$$

Def: Energy-momentum tensor

$$T^{\mu}_{\nu} = \frac{\delta \mathcal{L}}{\delta(\partial_{\mu}\phi)} \partial_{\nu}\phi - \delta^{\mu}_{\nu} \mathcal{L}$$

$\Rightarrow \partial_{\mu} T^{\mu}_{\nu} = 0$ conserved!