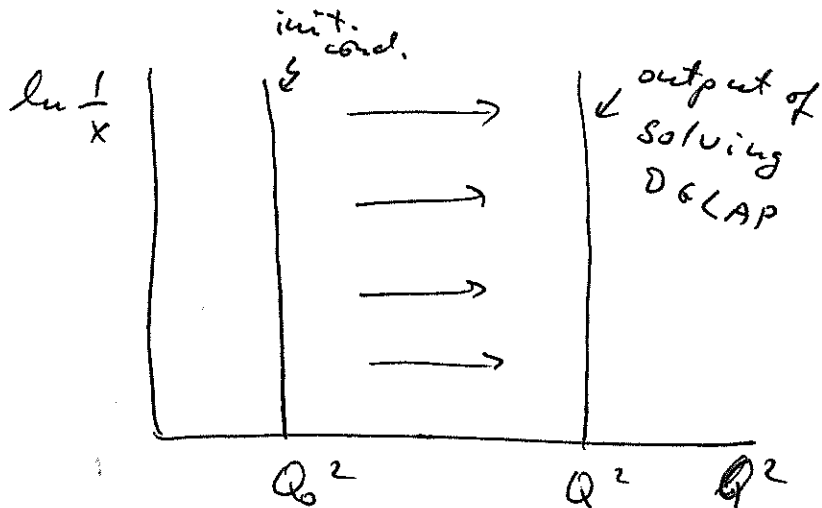


Last time

talked about DGLAP equation

~ usually one sets initial conditions at some $Q^2 = Q_0^2$ & "evolves" the PDFs with DGLAP to $Q^2 > Q_0^2$:



DGLAP at small x

Glucos dominate at small $x \Rightarrow$ drop quarks

$$\frac{\partial}{\partial \ln Q^2} G(x, Q^2) = \frac{\alpha_s(Q^2)}{2\pi} \int_x^1 \frac{dx_1}{x_1} P_{GG}\left(\frac{x}{x_1}\right) G(x_1, Q^2)$$

Def. moments: $G_n(Q^2) = \int_0^1 dx \cdot x^{n-1} \cdot G(x, Q^2)$

anomalous dimension: $\gamma_{GG}^{(n)} = \int_0^1 dz \cdot z^{n-1} P_{GG}(z)$

$$P_{GG}(z) \approx \frac{2N_c}{z} \text{ at small } -x.$$

In moments space, the solution is easy:

$$G_n(Q^2) = e^{\int_{Q_0^2}^{Q^2} \frac{dQ'^2}{Q'^2} \frac{\alpha_s(Q'^2)}{2\pi} \delta_{GG}^{(n)}} G_n(Q_0^2)$$

such that

$$G(x, Q^2) = \int \frac{dn}{2\pi i} x^{-n} e^{\frac{\delta_{GG}^{(n)}}{2\pi\beta_2} \ln\left(\frac{\ln Q^2/\Lambda^2}{\ln Q_0^2/\Lambda^2}\right)} G_n(Q_0^2)$$

with $\delta_{GG}^{(n)} = \frac{2N_c}{n-1}$

Performing saddle-point approximation we get

$$x G(x, Q^2) \sim e^{2\sqrt{\frac{N_c}{\pi\beta_2} \ln \frac{1}{x} \ln\left(\frac{\ln Q^2/\Lambda^2}{\ln Q_0^2/\Lambda^2}\right)}}$$

$\Rightarrow xG$ grows at small x , but slower than a power of x

\Rightarrow the data shows that $xG \sim \left(\frac{1}{x}\right)^\lambda$, with $\lambda \approx 0.3$ (a number). DGLAP evolution cannot give us this result \Rightarrow

\Rightarrow another evolution equation which resums logarithms of x (powers of $\alpha_s \ln \frac{1}{x}$) is needed to elucidate the small- x behaviour.

This equation exists and is known as the BFKL equation (Balitsky, Fadin, Kuraev, Lipatov)

Particle Production in High Energy

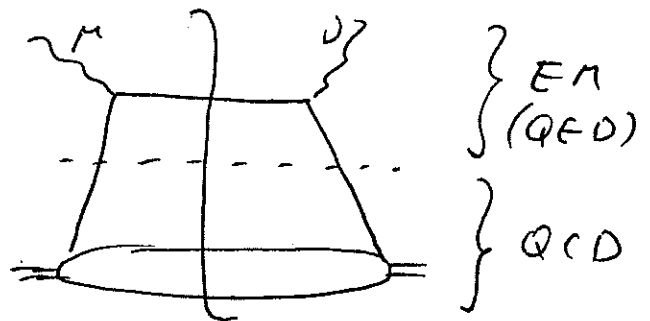
Hadronic Collisions.

Collinear Factorization

When we considered DIS above, we have factorized EM & QCD parts of the diagram:

We got

$$F_2(x, Q^2) = \sum_f e_f^2 x q_f^+(x)$$



⇒ this is an example of collinear factorization

In general one writes:

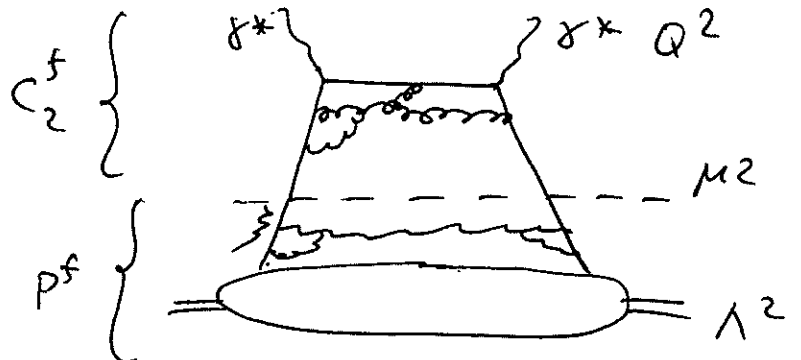
$$F_2(x, Q^2) = \sum_{f, \bar{f}, \text{gluons}} \int_0^1 d\xi C_2^f\left(\frac{x}{\xi}, Q^2, M^2\right) P^f(\xi, M^2) + O\left(\frac{m^2}{Q^2}\right)$$

higher twists
↓

$C_2^f \sim$ coefficient fun

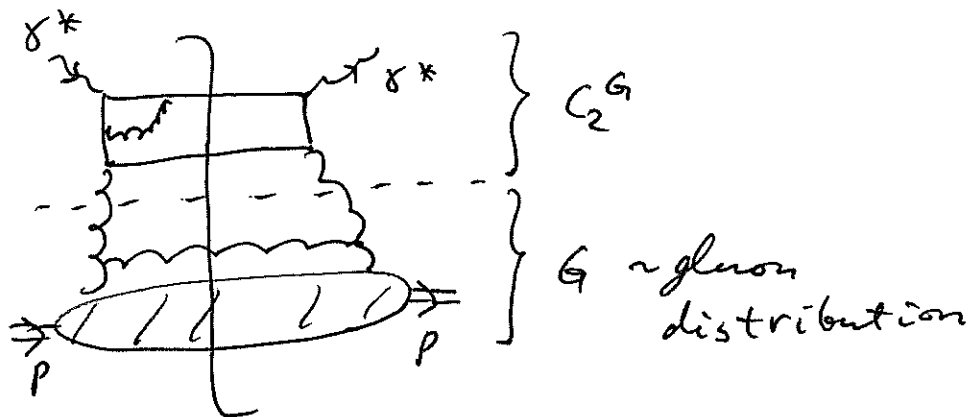
(may contain QCD

corrections at higher orders)



$P^f = \{q_f^+, q_f^-, G\}$; ξ - momentum fraction of the parton in P^f .

May also have



μ^2 is called factorization scale.

(112)

Note: C_2^f is perturbatively calculable,

P^f is not (though one has DGLAP for P^f)

$\Lambda^2 \lesssim \mu^2 \lesssim Q^2 \Rightarrow$ but F_2 does not depend on μ^2 , it is arbitrary $\Rightarrow \mu^2 \frac{d}{d\mu^2} F_2(x, Q^2) = 0$

Write $F_2 = C_2^f \otimes P^f$

$$\Rightarrow \mu^2 \frac{d}{d\mu^2} F_2 = 0 = \left(\mu^2 \frac{d}{d\mu^2} C_2^f \right) \otimes P^f + C_2^f \otimes \mu^2 \frac{d}{d\mu^2} P^f$$

\Rightarrow what happens (separation of variables, C_2^f depends on Q^2 , only P^f depends on μ^2):

$$\mu^2 \frac{d}{d\mu^2} P^f = \gamma(\alpha_s) \otimes P^f$$

\sim DGLAP evolution
for splitting
function.

$$\mu^2 \frac{d}{d\mu^2} C_2^f = -\gamma(\alpha_s) \otimes C_2^f$$

$$\Rightarrow \mu^2 \frac{d}{d\mu^2} F_2 = -\gamma \otimes C_2^f \otimes P^f + C_2^f \otimes \gamma \otimes P^f = 0.$$

as desired.

\Rightarrow can "place" corrections into PDF or coefficient function

Collinear factorization in DIS is a theorem which can be proven \Rightarrow must be right! (at large- Q^2 only!)

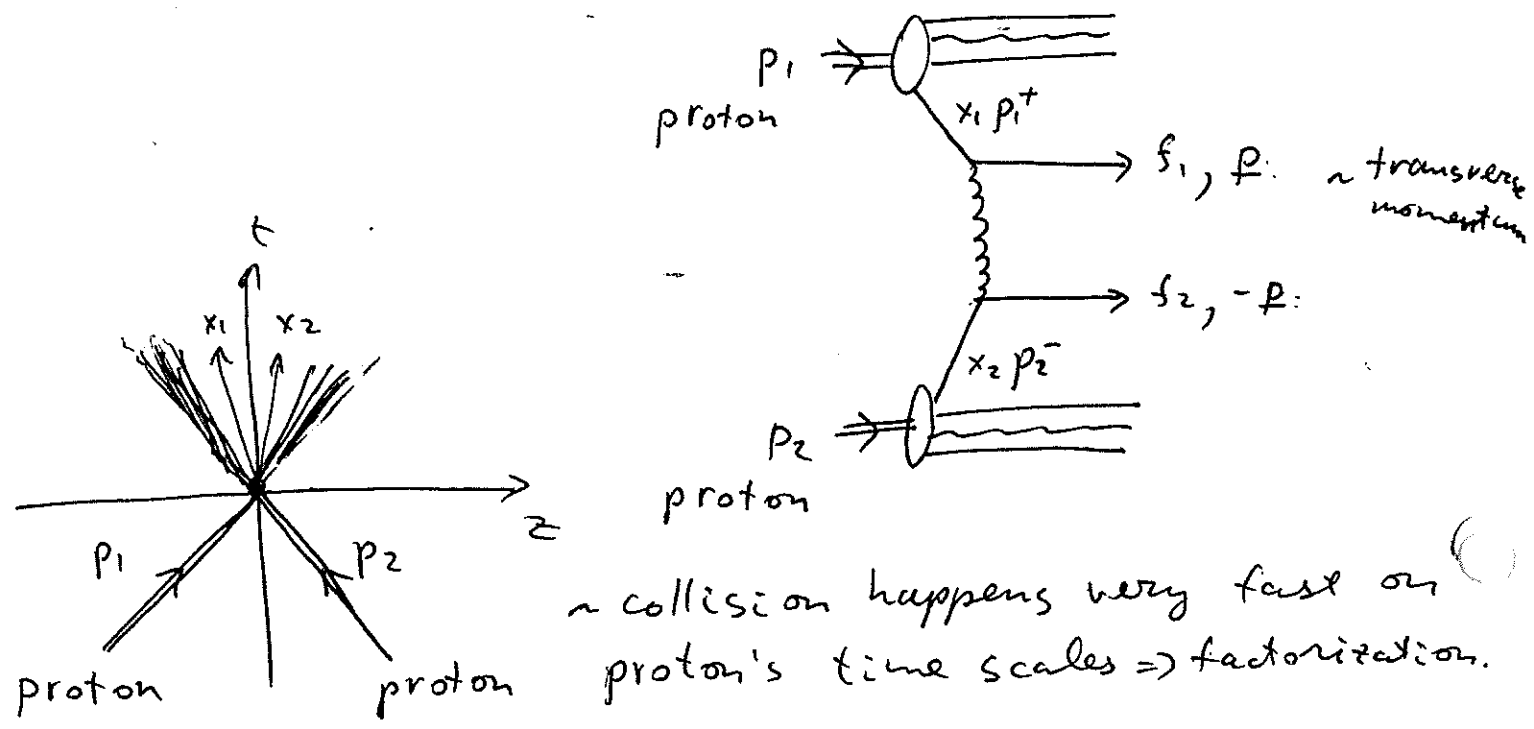
\sim at LO have $C_2^f = S(\frac{x}{\xi} - 1) e_f^2$, $f = \text{quarks only}$

$$\Rightarrow F_2(x, Q^2) = \sum_f \int_0^1 d\xi \underbrace{S(\frac{x}{\xi} - 1)}_X e_f^2 g^f(\xi)$$

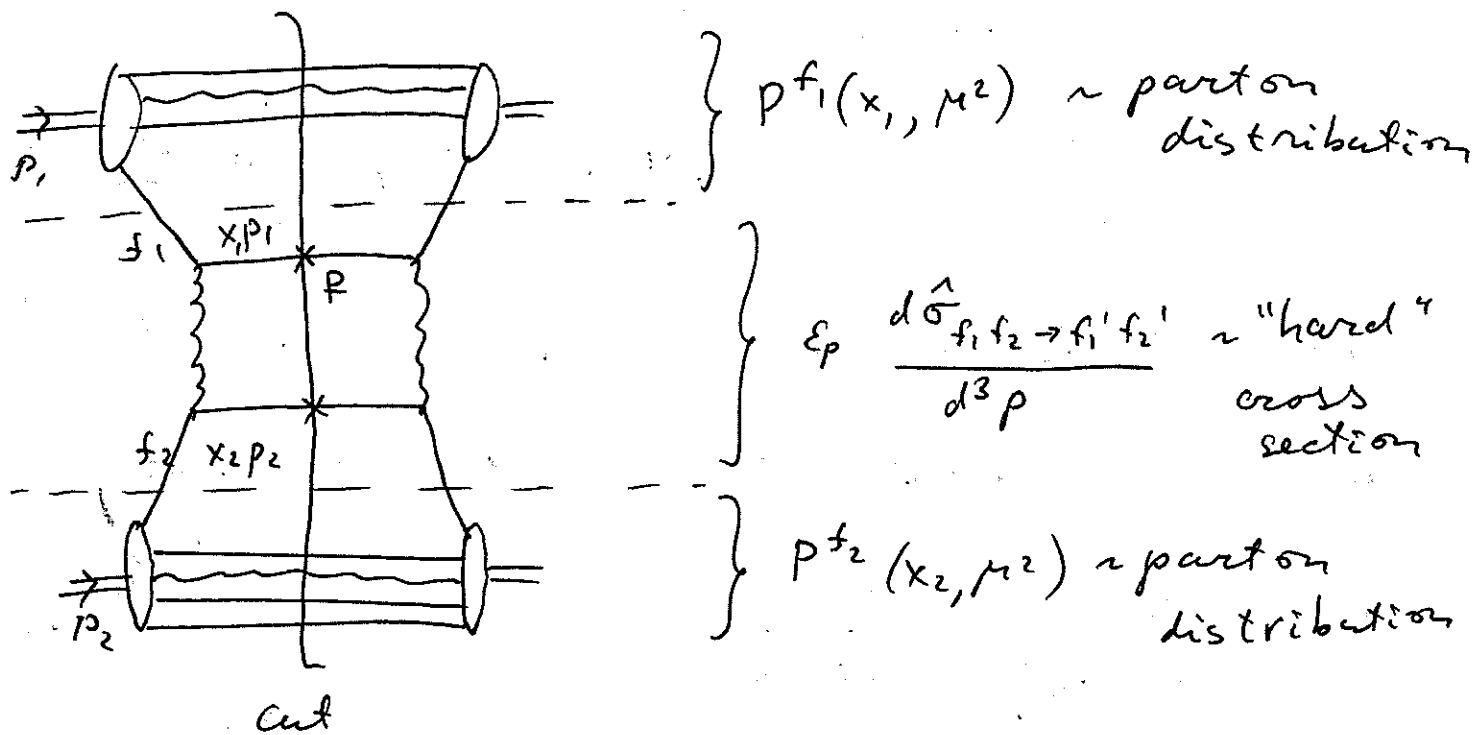
$$= \sum_f e_f^2 x g^f(x) \text{ as expected!}$$

Jet Production in Hadronic Collisions.

Collinear factorization also applies to hadron-hadron collisions. Consider quark production:



Square the diagram:



The collinear factorization formula then reads:

$$\mathcal{E}_p \frac{d\sigma}{d^3 p} = \sum_{i,j} \int_0^1 dx_1 \int_0^1 dx_2 P^{f_i}(x_1, \mu^2) \cdot \mathcal{E}_p \frac{d\hat{\sigma}_{f_i f_j \rightarrow f_i' f_j'}}{d^3 p} \cdot P^{f_j}(x_2, \mu^2)$$

Usually put $\mu^2 = p_T^2$ for large \$p_T\$ jets (or hadrons)

after the collision quarks (gluons) that are produced get dressed by further emissions.

But the flow of energy is not likely to

be modified much by those. (Still people construct other IR-safe observables insensitive to late-time emissions: (— + — + — ...))

Example Quark jet production (coming from q)
 quarks). Replace $p_{fi} \rightarrow q^f \Rightarrow$ write

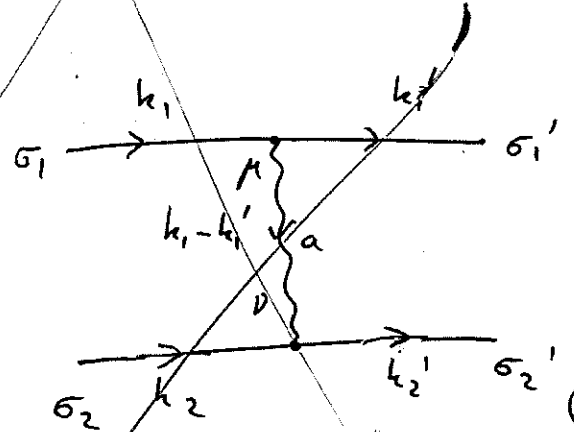
$$\epsilon_p \frac{d\sigma}{d^3p} = \sum_{f_1, f_2} \int_0^1 dx_1 dx_2 q^{f_1}(x_1, p_T^2) \epsilon_p \frac{d\hat{\sigma}_{f_1 f_2 \rightarrow f_1' f_2'}}{d^3p} q^{f_2}(x_2, p_T^2)$$

$q^f \sim$ to be found from DGLAP (PDF data)

We can calculate the hard cross section:

$$d\hat{\sigma} = \frac{1}{2\epsilon_1 2\epsilon_2 \cdot 2} \frac{d^3k_1'}{(2\pi)^3 2\epsilon_1'}$$

" "
 $|\vec{v}_1 - \vec{v}_2|$



$$\frac{d^3k_2'}{(2\pi)^3 2\epsilon_2'} (2\pi)^4 \delta^{(4)}(k_1 + k_2 - k_1' - k_2')$$

$$\cdot |M|^2 \left[\delta^3(\vec{k}_1' - \vec{p}) + \delta^3(\vec{k}_2' - \vec{p}) \right] \cdot d^3p$$

measured jet can be either quark!

$$\Rightarrow \epsilon_p \frac{d\hat{\sigma}}{d^3p} = ?$$

$\delta^{(3)}$ kills d^3k_2'

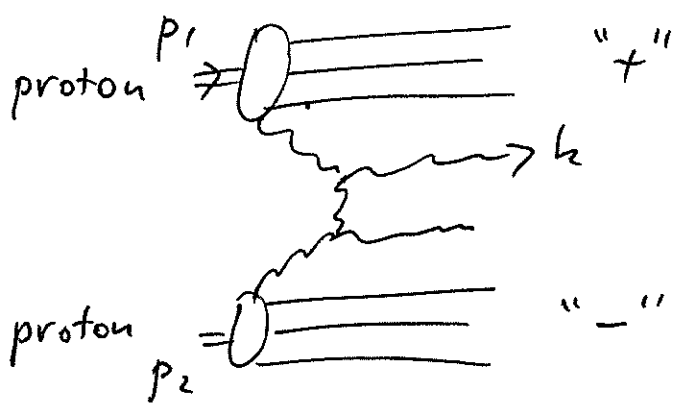
kills d^3k_1'

$$\frac{d^3k_1'}{(2\pi)^3 2\epsilon_1'} \frac{d^3k_2'}{(2\pi)^3 2\epsilon_2'} (2\pi)^4 \delta^{(4)}(k_1 + k_2 - k_1' - k_2') \delta^3(\vec{k}_1' - \vec{p}) =$$

$$= \frac{1}{(2\pi)^2} \frac{1}{4\epsilon_p \epsilon_2'} \delta(E_1 + E_2 - \epsilon_p - \epsilon_2')$$

Rapidity in p+p collisions

Def. Rapidity variable $\eta \equiv \frac{1}{2} \ln \frac{k^+}{k^-}$ for an outgoing particle with momentum 4-vector k^M .



at a collider, one proton comes in with

$$p_1^M = (p_1^+, \frac{m^2}{2p_1^+}, \underline{0}),$$

another one has

$$p_2^M = (\frac{m^2}{2p_2^-}, p_2^-, \underline{0}).$$

p_1^+, p_2^- are very large, $m = \text{proton mass}$

$$s = (p_1 + p_2)^2 \approx 2p_1 \cdot p_2 = 2p_1^+ p_2^-.$$

Proton rapidities: $y_1 = \frac{1}{2} \ln \frac{p_1^+}{m^2/2p_1^+} = \ln \frac{p_1^+ \sqrt{2}}{m}.$

$$y_2 = \frac{1}{2} \ln \frac{m^2/2p_2^-}{p_2^-} = - \ln \frac{p_2^- \sqrt{2}}{m}.$$

y_1 is large and positive, y_2 is large and < 0 .

The net rapidity interval between the protons ()

is $y_1 - y_2 = \ln \frac{2p_1^+ p_2^-}{m^2} = \ln \frac{S}{m^2} \Rightarrow$ the higher is the energy, the wider is the interval.

$$y_1 - y_2 = \ln \frac{S}{m^2}$$

Rapidity interval between our particle h and the proton p_1 is $y_1 - \eta = \ln \frac{p_1^+ \sqrt{s}}{m} - \frac{1}{2} \ln \frac{k^+}{k^-} =$

$= \left\{ \begin{array}{l} \text{as } k^- = \frac{k^2}{2k^+} \text{ for } \\ \text{massless particle} \end{array} \right. = \ln \frac{p_1^+ \sqrt{s}}{m} - \ln \frac{k^+ \sqrt{s}}{k_\perp} = \ln \frac{p_1^+}{k^+} + \ln \frac{k_\perp}{m}$

for $k_\perp \approx m$ get $y_1 - \eta = \ln \frac{p_1^+}{k^+} = \ln \frac{1}{x_1}$, where $x_1 = \frac{k^+}{p_1^+}$

is the Bjorken- x in proton #1: $y_1 - \eta \approx \ln \frac{1}{x_1}$

Similarly, $\eta - y_2 \approx \ln \frac{1}{x_2}$ with $x_2 = \frac{k^-}{p_2^-}$ ~ Bjorken- x in proton #2

We thus have

