

Last time: worked out interactions of  $W^\pm$ ,  $Z$  &  $\gamma$  with quarks. We got:

(i) Charged current ( $W^\pm$  bosons coupling)

$$\mathcal{L}_{c.c.} = \frac{g}{2\sqrt{2}} \left\{ \bar{u} \gamma \cdot W (1-\gamma_5) [V_{ud} d + V_{us} s + V_{ub} b] + \bar{c} \gamma \cdot W (1-\gamma_5) [V_{cd} d + V_{cs} s + V_{cb} b] + \bar{t} \gamma \cdot W (1-\gamma_5) [V_{td} d + V_{ts} s + V_{tb} b] + h.c. \right\}$$

(ii) Neutral current: ( $Z$ -bosons & photons coupling)

$$\mathcal{L}_{n.c.} = \frac{g}{4\cos\theta_w} \left\{ \bar{u} \gamma \cdot Z \left[ (1-\gamma_5) \left( 1 - \frac{4}{3} \sin^2\theta_w \right) - (1+\gamma_5) \frac{4}{3} \sin^2\theta_w \right] u - \bar{D} \gamma \cdot Z \left[ (1-\gamma_5) \left( 1 - \frac{2}{3} \sin^2\theta_w \right) - (1+\gamma_5) \frac{2}{3} \sin^2\theta_w \right] D + \sum_f e_f \bar{f} \gamma \cdot A f \right\}$$

where we have defined

$$u = \begin{pmatrix} u \\ c \\ t \end{pmatrix}, \quad D = \begin{pmatrix} d \\ s \\ b \end{pmatrix}, \quad e_f = e \begin{cases} 2/3 & \text{for } u, c, t \\ -1/3 & \text{for } d, s, b \end{cases}$$

$f = u, d, s, c, b, t$  for different index values.

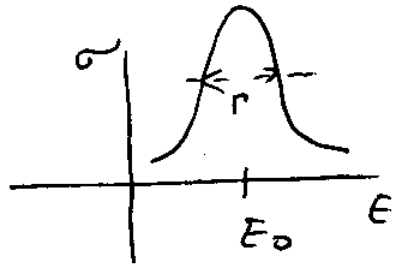
$$d\Gamma = \frac{1}{2m} \prod_{i=1}^N \frac{d^3 p_i}{(2\pi)^3 2E_i} |M(m; p_1, \dots, p_N)|^2 (2\pi)^4 \delta(k - \sum_{j=1}^N p_j)$$

$\sim$  decay rate

$$\Gamma = \frac{\# \text{ decays per unit time}}{\# \text{ of particles}}$$

Breit-Wigner formula for scattering amplitude:

$$f(E) \sim \frac{1}{E - E_0 + i\Gamma/2}$$



$$\Rightarrow \sigma(E) \propto |f(E)|^2 \sim \frac{1}{(E - E_0)^2 + \frac{\Gamma^2}{4}}$$

Width of resonance peak!

$\Rightarrow$  **(NB)** in calculating  $|M|^2$  sum over spins, etc in the final state, average over them in the initial state.

### Decay of the Z-boson.

$\Rightarrow$  let's calculate decay rate of the Z-boson

$\Rightarrow$  the Z-boson interaction Lagrangian is:

$$\mathcal{L}_Z = \frac{g}{4\cos\theta_W} \left\{ \bar{\nu}_e \gamma_\mu Z (1-\gamma_5) \nu_e + 2\sin^2\theta_W \bar{e} \gamma_\mu Z (1+\gamma_5) e + (2\sin^2\theta_W - 1) \bar{e} \gamma_\mu Z (1-\gamma_5) e + \bar{u} \gamma_\mu Z \left[ (1-\gamma_5) \left( 1 - \frac{4}{3}\sin^2\theta_W \right) - (1+\gamma_5) \left( \frac{4}{3}\sin^2\theta_W \right) \right] u - \bar{d} \gamma_\mu Z \left[ (1-\gamma_5) \left( 1 - \frac{2}{3}\sin^2\theta_W \right) - (1+\gamma_5) \left( \frac{2}{3}\sin^2\theta_W \right) \right] d \right\}$$

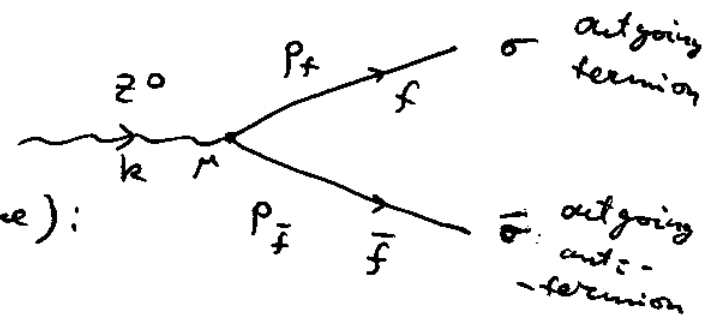
$Z \rightarrow W^+ W^-$ ,  $Z \rightarrow \eta Z$ ,  $Z \rightarrow W^+ W^- Z$ , ... are all prohibited by energy conservation!  $M_Z \approx 91 \text{ GeV} < 2M_W = 2 \cdot 80 \text{ GeV} \dots$

For each fermion species the Lagrangian

looks like:  $\mathcal{L} = \bar{\Psi} \gamma_0 z [a_f (1 - \gamma_5) - b_f (1 + \gamma_5)] \Psi$

with  $a_f, b_f$  coefficients being species-dependent.

Consider  $z$ -decay:



Amplitude (in  $z^0$  rest frame):

$$M = \underbrace{-i}_{\text{vertex}} \bar{u}_\sigma(p_f) \gamma^\mu \underbrace{[a_f (1 - \gamma_5) - b_f (1 + \gamma_5)]}_{\text{vertex}} v_{\bar{\sigma}}(p_{\bar{f}}) \cdot \epsilon_\mu^\lambda(k)$$

Need to find:

$$\frac{1}{3} \sum_{\lambda, \sigma, \bar{\sigma}} |M|^2 = \frac{1}{3} \sum_{\lambda, \sigma, \bar{\sigma}} \bar{u}_\sigma \gamma_0 \epsilon^\lambda [a_f (1 - \gamma_5) - b_f (1 + \gamma_5)]$$

↑ average over initial polarizations  
 ← sum over spins

$$v_{\bar{\sigma}} \underbrace{\bar{v}_{\bar{\sigma}} \gamma^0}_{v^\dagger \gamma^0 \gamma^0 = 1} [a_f^* (1 - \gamma_5^\dagger) - b_f^* (1 + \gamma_5^\dagger)] \gamma^\dagger \cdot \epsilon^{\lambda*} \gamma^0 u_\sigma$$

$\Rightarrow$  as  $(\gamma^0)^\dagger = \gamma^0$ ,  $(\gamma^i)^\dagger = -\gamma^i$ ,  $i = 1, 2, 3$

$\Rightarrow \gamma^0 \gamma_\mu^\dagger \gamma^0 = \begin{cases} \gamma^0 & \text{if } \mu = 0 \\ \gamma^i & \text{if } \mu = i \end{cases} \Rightarrow \gamma^0 (\gamma^\mu)^\dagger \gamma^0 = \gamma^\mu$

as  $\{\gamma^\mu, \gamma^5\} = 0 \Rightarrow \gamma^0 \gamma^5 (\gamma^\mu)^\dagger \gamma^0 = \begin{cases} -\gamma^5 \gamma^0, & \mu = 0 \\ -\gamma^5 \gamma^i, & \mu = i \end{cases}$

$\Rightarrow \gamma^5 \rightarrow -\gamma^5$

$$\Rightarrow \frac{1}{3} \sum_{\lambda, \sigma, \bar{\sigma}} |M|^2 = \frac{1}{3} \sum_{\lambda, \sigma, \bar{\sigma}} \bar{u}_{\sigma}(p_f) \gamma \cdot \epsilon^{\lambda} [a_f (1 - \delta_5) - b_f (1 + \delta_5)] v_{\bar{\sigma}}(p_f)$$

$$\cdot \bar{v}_{\bar{\sigma}}(p_{\bar{f}}) [a_f^* (1 + \delta_5) - b_f^* (1 - \delta_5)] \gamma \cdot \epsilon^{\lambda *} u_{\sigma}(p_f)$$

Now, use  $\sum_{\sigma} u_{\sigma}(p_f) \bar{u}_{\sigma}(p_f) = \not{p}_f + m_f$

$$\sum_{\bar{\sigma}} v_{\bar{\sigma}}(p_{\bar{f}}) \bar{v}_{\bar{\sigma}}(p_{\bar{f}}) = \not{p}_{\bar{f}} - m_f$$

$$\Rightarrow \frac{1}{3} \sum_{\lambda} |M|^2 = \frac{1}{3} \sum_{\lambda} \text{Tr} [(\not{p}_f + m_f) \gamma \cdot \epsilon^{\lambda} [a_f (1 - \delta_5) - b_f$$

$$\cdot (1 + \delta_5)] (\not{p}_{\bar{f}} - m_f) [a_f^* (1 + \delta_5) - b_f^* (1 - \delta_5)] \gamma \cdot \epsilon^{\lambda *} ] =$$

$$= \frac{1}{3} \sum_{\lambda} \text{Tr} [(\not{p}_f + m_f) \gamma \cdot \epsilon^{\lambda} [2(|a_f|^2 (1 - \delta_5) + |b_f|^2 (1 + \delta_5)) \not{p}_{\bar{f}} +$$

$$+ m_f (2 b_f a_f^* (1 + \delta_5) + 2 a_f b_f^* (1 - \delta_5))] \gamma \cdot \epsilon^{\lambda *} ] =$$

$$= \left( \text{as } m_{u,d} \ll M_Z \Rightarrow \text{neglect masses} \right) = \frac{2}{3} \sum_{\lambda} \text{Tr} [ \not{p}_f \gamma \cdot \epsilon^{\lambda}$$

$$\cdot (|a_f|^2 (1 - \delta_5) + |b_f|^2 (1 + \delta_5)) \not{p}_{\bar{f}} \gamma \cdot \epsilon^{\lambda *} ]$$

In the z-boson rest frame  $\epsilon_{\mu}^{(1)} = (0, 1, 0, 0)$ ,

$\epsilon_{\mu}^{(2)} = (0, 0, 1, 0)$ ,  $\epsilon_{\mu}^{(3)} = (0, 0, 0, 1)$  as  $k \cdot \epsilon = M \cdot \epsilon_0 = 0 \Rightarrow \epsilon_0 = 0$

$$\Rightarrow \sum_{\lambda} \epsilon_{\mu}^{\lambda} \epsilon_{\nu}^{\lambda *} = -g_{\mu\nu} \text{ for } \mu, \nu = 1, 2, 3 \quad (\emptyset \text{ otherwise})$$

$$\Rightarrow \frac{1}{3} \sum |M|^2 = \frac{2}{3} \text{Tr} \left[ \gamma^i \gamma \cdot p_f \gamma^i (|a_f|^2 (1-\gamma_5) + |b_f|^2 (1+\gamma_5)) \gamma \cdot p_{\bar{f}} \right]$$

$$\gamma^i \gamma \cdot p_f \gamma^i = \underbrace{\gamma^i \gamma^0 \gamma^i}_{3\gamma^0} p_f - \underbrace{\gamma^i \gamma^j \gamma^i}_{\gamma^i (-\delta^i \delta^i + \{\gamma^j, \gamma^i\})} p_f^j =$$

$$= +3\gamma^0 p_f - 2\gamma^i p_f^i = \gamma^j$$

$$= 3\gamma^0 p_f - \vec{\gamma} \cdot \vec{p}_f$$

$$\Rightarrow \frac{1}{3} |M|^2 = \frac{2}{3} \text{Tr} \left[ (3\gamma^0 p_f - \vec{\gamma} \cdot \vec{p}_f) (|a_f|^2 (1-\gamma_5) + |b_f|^2 (1+\gamma_5)) \right]$$

$$\gamma \cdot p_{\bar{f}} = \left[ \begin{array}{l} \text{as } \text{Tr}(\gamma^\mu \gamma^\nu \gamma^5) = 0 \\ \Rightarrow \text{drop } \gamma^5 \text{'s} \\ \text{tr } \gamma^\mu \gamma^\nu = 4g^{\mu\nu} \end{array} \right] = \frac{2}{3} (|a_f|^2 + |b_f|^2) \cdot [4 \cdot 3 \cdot p_f \cdot p_{\bar{f}} - 4 \cdot \vec{p}_f \cdot \vec{p}_{\bar{f}}]$$

$$- 4 \cdot \vec{p}_f \cdot \vec{p}_{\bar{f}} = \left[ \begin{array}{l} \text{Z CMS frame } \Rightarrow \\ \Rightarrow \vec{p}_{\bar{f}} = -\vec{p}_f, p_{\bar{f}} = p_f \end{array} \right] \Rightarrow = \frac{8}{3} (|a_f|^2 + |b_f|^2)$$

$$\cdot (3p_f^2 + \vec{p}_f^2) = \frac{32}{3} |\vec{p}_f|^2 (|a_f|^2 + |b_f|^2) = \frac{1}{3} \sum |M|^2$$

$\Rightarrow$  the decay rate:

$$d\Gamma = \frac{1}{2M_2} \frac{d^3 p_f}{(2\pi)^3 2E_f} \frac{d^3 p_{\bar{f}}}{(2\pi)^3 2E_{\bar{f}}} (2\pi)^4 \delta^{(4)}(k - p_f - p_{\bar{f}}) \frac{1}{3} \sum |M|^2$$

$$\Rightarrow \Gamma = \frac{1}{2M_2} \int \frac{d^3 p_f d^3 p_{\bar{f}}}{(2\pi)^6 4E_f E_{\bar{f}}} (2\pi)^4 \delta^{(4)}(k - p_f - p_{\bar{f}}) \frac{32}{3} |\vec{p}_f|^2$$

$$\cdot (|a_f|^2 + |b_f|^2) = \frac{1}{2M_z} \frac{g}{3} \frac{1}{(2\pi)^2} \int d^3 p_f \cdot \delta(M_z - 2p_f) \cdot$$

$$\cdot (|a_f|^2 + |b_f|^2) = \frac{4}{3M_z} \frac{1}{(2\pi)^2} \cdot \cancel{4\pi} \cdot \left(\frac{M_z}{2}\right)^2 \cdot \frac{1}{2} (|a_f|^2 + |b_f|^2)$$

$$= \frac{M_z}{6\pi} [ |a_f|^2 + |b_f|^2 ] \Rightarrow \text{finally, as } a_f, b_f \text{ are}$$

real  $\Rightarrow$  drop  $|\dots| \Rightarrow$

$$\Gamma_{z \rightarrow f\bar{f}} = \frac{M_z}{6\pi} [ a_f^2 + b_f^2 ]$$

a) Neutrinos:  $b_\nu = 0, a_\nu = \frac{g}{4\cos\theta_w}$

$$\Rightarrow \Gamma_{z \rightarrow \nu\bar{\nu}} = \frac{M_z}{6\pi} \frac{g^2}{16\cos^2\theta_w} = \frac{g^2 M_z}{96\pi \cos^2\theta_w}$$

b) Electrons:  $a_e = \frac{g}{4\cos\theta_w} (2\sin^2\theta_w - 1)$

$$b_e = \frac{-g}{4\cos\theta_w} 2\sin^2\theta_w$$

$$\Rightarrow \Gamma_{z \rightarrow e^+e^-} = \frac{g^2 M_z}{96\pi \cos^2\theta_w} [ (2\sin^2\theta_w - 1)^2 + 4\sin^4\theta_w ]$$

$$= \frac{g^2 M_z}{96\pi \cos^2\theta_w} \frac{1}{2} [ 1 + (1 - 4\sin^2\theta_w)^2 ] = \Gamma_{z \rightarrow \nu\bar{\nu}} \cdot \frac{1}{2} [ 1 + (1 - 4\sin^2\theta_w)^2 ]$$

c) u-quarks:  $a_u = \frac{g}{4\cos\theta_w} (1 - \frac{4}{3}\sin^2\theta_w)$

$$b_u = \frac{g}{4\cos\theta_w} \frac{4}{3}\sin^2\theta_w$$

$$\Gamma_{Z \rightarrow u\bar{u}} = \Gamma_{Z \rightarrow \nu\bar{\nu}} \times (3 \text{ colors}) \times \left[ \left(1 - \frac{4}{3} \sin^2 \theta_w\right)^2 + \left(\frac{4}{3} \sin^2 \theta_w\right)^2 \right] \quad (132)$$

$$= \Gamma_{Z \rightarrow \nu\bar{\nu}} \cdot \frac{3}{2} \cdot \left[ 1 + \left(1 - \frac{8}{3} \sin^2 \theta_w\right)^2 \right]$$

d) d-quarks

$$a_d = - \frac{g}{4 \cos \theta_w} \left(1 - \frac{2}{3} \sin^2 \theta_w\right)$$

$$b_d = - \frac{g}{4 \cos \theta_w} \frac{2}{3} \sin^2 \theta_w$$

$$\Rightarrow \Gamma_{Z \rightarrow d\bar{d}} = \Gamma_{Z \rightarrow \nu\bar{\nu}} \cdot 3 \left[ \left(1 - \frac{2}{3} \sin^2 \theta_w\right)^2 + \left(\frac{2}{3} \sin^2 \theta_w\right)^2 \right]$$

$$= \Gamma_{Z \rightarrow \nu\bar{\nu}} \cdot \frac{3}{2} \cdot \left[ 1 + \left(1 - \frac{4}{3} \sin^2 \theta_w\right)^2 \right]$$

The total Z-boson decay width:

$$\Gamma_Z = \frac{g^2 M_Z}{192 \pi \cos^2 \theta_w} \left\{ 2 N_\nu + \left[ 1 + (1 - 4 \sin^2 \theta_w)^2 \right] N_e + \right. \\ \left. + 3 \left[ 1 + \left(1 - \frac{8}{3} \sin^2 \theta_w\right)^2 \right] N_u + 3 \left[ 1 + \left(1 - \frac{4}{3} \sin^2 \theta_w\right)^2 \right] N_d \right\}$$

$$\text{as } M_Z \approx 91 \text{ GeV} \Rightarrow N_u = 2 \quad \left( \begin{array}{l} \text{u, c only,} \\ m_t \approx 170 \text{ GeV} \end{array} \right)$$

$$N_d = 3 \quad (d, s, b)$$

$$\text{know } N_e = 3 \quad (e, \mu, \tau)$$

$$\text{as } \sin^2 \theta_w \approx \frac{1}{4} \Rightarrow \Gamma_{Z \rightarrow \nu\bar{\nu}} : \Gamma_{Z \rightarrow e^+e^-} : \Gamma_{Z \rightarrow u\bar{u}} : \Gamma_{Z \rightarrow d\bar{d}} =$$

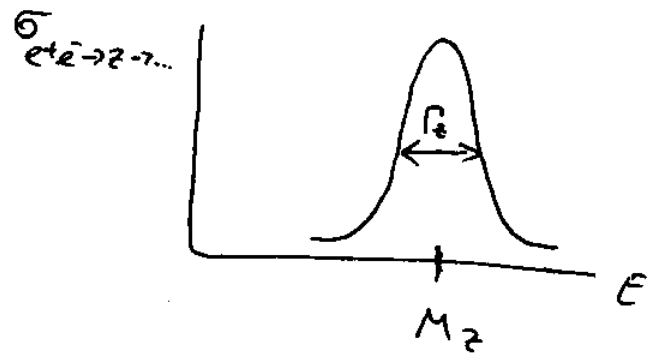
$$= 2 N_\nu : N_e : \frac{10}{3} N_u : \frac{13}{3} N_d$$

$$\Rightarrow \text{measure } \Gamma_Z \approx 2.5 \text{ GeV}, \Gamma_{Z \rightarrow e^+e^-}, \Gamma_{Z \rightarrow u\bar{u}}, \Gamma_{Z \rightarrow d\bar{d}} \Rightarrow$$

=> get  $N_D = 2.984 \pm 0.008$

(look at  $e^+e^- \rightarrow z \rightarrow \dots$  at LEP, SLAC...)

=> there are 3 neutrino generations!



### Neutrino Masses and Oscillations.

=> imagine that neutrinos have a mass (we know they do) & lepton # is not conserved ( $\nu_e \rightarrow \nu_\mu, \dots$ )

=> again mass eigenstate  $\neq$  weak eigenstates

$(\nu_1, \nu_2, \nu_3)$   $(\nu_e, \nu_\mu, \nu_\tau)$

just like quarks

=> consider 2 generations for simplicity:

$$\begin{pmatrix} \nu_\mu \\ \nu_e \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

↙ mixing angle

Maki et al '69  
 Pontecorvo '58, '68  
 Gribov & Pontecorvo '69

=> in EW process one produces  $\nu_e$  or  $\nu_\mu$ .

propagation:  $\nu_1(t) = \nu_1(0) e^{-iE_1 t}, \nu_2(t) = \nu_2(0) e^{-iE_2 t}$



$\Rightarrow$  write (at  $t=0$ ) that we have  $\nu_\mu(0)=1, \nu_e(0)=0$

$\sim$  a pure  $\mu$ -neutrino:

$$\Rightarrow \begin{cases} 1 = \nu_1(0) \cos \theta + \nu_2(0) \sin \theta \\ 0 = -\nu_1(0) \sin \theta + \nu_2(0) \cos \theta \end{cases} \Rightarrow \begin{cases} \nu_1(0) = \cos \theta \\ \nu_2(0) = \sin \theta \end{cases}$$

$$\begin{aligned} \Rightarrow \nu_\mu(t) &= \underbrace{e^{-iE_1 t}}_{\nu_1(t)} \nu_1(0) \cos \theta + \underbrace{e^{-iE_2 t}}_{\nu_2(t)} \nu_2(0) \sin \theta = \\ &= e^{-iE_1 t} \cos^2 \theta + e^{-iE_2 t} \sin^2 \theta \end{aligned}$$

$$\begin{aligned} \Rightarrow |\nu_\mu(t)|^2 &= \left| e^{-iE_1 t} \cos^2 \theta + e^{-iE_2 t} \sin^2 \theta \right|^2 = \\ &= \cos^4 \theta + \sin^4 \theta + \cos^2 \theta \cdot \sin^2 \theta \cdot 2 \cos((E_1 - E_2)t) \\ &= 1 + 2 \sin^2 \theta \cos^2 \theta \left[ \cos((E_1 - E_2)t) - 1 \right] = \\ &= 1 - \sin^2(2\theta) \cdot \sin^2 \left( \frac{E_1 - E_2}{2} t \right). \end{aligned}$$

For small masses:  $E = \sqrt{m^2 + p^2} \approx p + \frac{m^2}{2p}$

$$\Rightarrow E_1 - E_2 \approx \frac{m_1^2}{2E_1} - \frac{m_2^2}{2E_2} \approx \frac{m_1^2 - m_2^2}{2E} \quad \text{as } E_1 \approx E_2 \text{ at this order}$$

$$\Rightarrow P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2(2\theta) \cdot \sin^2 \left( \frac{1.27 \Delta m^2 L}{E} \right)$$

with  $\Delta m^2 \equiv m_2^2 - m_1^2$ ,  $L = t$  (meters),  $E$  is in MeV.  
(eV)<sup>2</sup>/c<sup>4</sup>

$$P_{\nu_{\mu} \rightarrow \nu_e} = \sin^2(2\theta) \cdot \sin^2\left(\frac{1.27 \Delta m^2 L}{E}\right)$$

$\Rightarrow \nu_{\mu}$  can turn into  $\nu_e$  & vice versa  $\Rightarrow$   
 $\Rightarrow$  neutrino oscillations!

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$$|\nu_{\mu}(0)\rangle = \cos\theta |\nu_1(0)\rangle + \sin\theta |\nu_2(0)\rangle$$

$\sim$  initially produced a true  $\nu_{\mu}$  state (EW state)

$$|\nu_e(0)\rangle = -\sin\theta |\nu_1(0)\rangle + \cos\theta |\nu_2(0)\rangle, \quad |\psi_i\rangle = |\nu_{\mu}(0)\rangle + \theta \cdot |\nu_e(0)\rangle$$

$\sim$  this is a true  $\nu_e$  state  $\perp$   $\nu_{\mu}$  state (not there initially)

$$|\nu_{\mu}(t)\rangle = e^{-iE_1 t} \cos\theta |\nu_1(0)\rangle + e^{-iE_2 t} \sin\theta |\nu_2(0)\rangle$$

$$|\langle \nu_{\mu}(0) | \nu_{\mu}(t) \rangle|^2 = \left| \cos^2\theta e^{-iE_1 t} + \sin^2\theta e^{-iE_2 t} \right|^2$$

$\Rightarrow$  the rest is like above

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$\sim$  solar neutrino problem: #  $\nu_e$ 's from the Sun was  $\sim 3$  times smaller than expected

(Ray Davies '68, John Bahcall '80)

$\sim$  SNO experiment in 2003 measured  $\nu_e$  and  $\nu_{\mu}$  from the sun: total # of neutrinos was just right, in agreement with solar models  $\Rightarrow$  oscillations!

~ also Super-Kamiokande, KamLAND.

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=> assuming 3 neutrino flavors get: (PDG):

$$\sin^2 \theta_{12} \simeq 0.86$$

$$\Delta m_{12}^2 = 8 \times 10^{-5} \text{ eV}^2$$

$$\sin^2 \theta_{23} \gtrsim 0.92$$

$$\Delta m_{32}^2 = 1.9 \div 3.0 \times 10^{-3} \text{ eV}^2$$

Note the large mixing angles

=> mass hierarchy of neutrinos has not been worked out either... experiments keep on going...

Can we "fix" Standard Model to include right-handed neutrinos? Sure we can; for instance do like for quarks: postulate right-handed neutrino

singlet  $\nu_R \Rightarrow \mathcal{L}_{R.H.V} = G_R \left[ \bar{L}_L \tilde{\phi} \nu_R + \text{c.c.} \right] + \dots$

$\downarrow \quad \downarrow \quad \downarrow$   
 $Y=+1 \quad Y=-1 \quad Y=0 \Rightarrow \text{ok.}$

=> VEV of  $\tilde{\phi}$  is  $\langle 0 | \tilde{\phi} | 0 \rangle = \begin{pmatrix} v/\sqrt{2} \\ 0 \end{pmatrix}$

=>  $\mathcal{L}_{R.H.V} = G_R \left[ (\bar{\nu}_L e_L) \begin{pmatrix} v/\sqrt{2} \\ 0 \end{pmatrix} \nu_R + \text{c.c.} \right] = G_R \left[ \bar{\nu}_L \nu_R + \bar{\nu}_R \nu_L \right] \frac{v}{\sqrt{2}}$

$= \frac{G_R v}{\sqrt{2}} \bar{\nu} \nu \Rightarrow m_\nu = \frac{G_R v}{\sqrt{2}} \Rightarrow \text{as } m_\nu \geq 0.04 \text{ eV,}$   
 $v \approx 289 \text{ GeV}$

=>  $G_R = \frac{m_\nu \sqrt{2}}{v} \approx 2 \times 10^{-13} \sim \text{too much fine-tuning ...}$   
(some say "why not?")

# "Naturalness" Problem:

$$\mathcal{L}_{\text{Higgs}} = \frac{1}{2} \partial_\mu \eta \partial^\mu \eta - m^2 \eta^2 - \lambda \eta^4 - \frac{\lambda}{4} \eta^4$$

=> have diagrams like



giving  $m_H^2 = m_{H,0}^2 + \# \cdot \lambda \cdot \int \frac{d^4 k}{(2\pi)^4} \frac{1}{k^2 - m_{H,0}^2 + i\epsilon}$

↑  
bare mass

diverges like  $\Lambda_{UV}^2$

$$\Rightarrow M_H^2 = m_{H,0}^2 + \# \cdot \lambda \cdot \Lambda_{UV}^2$$

What is  $\Lambda_{UV}$ ? Could be large, like  $M_{\text{Planck}} \approx 10^{19} \text{ GeV}$ ...

=> for that to work & give "finite"  $m_H$  need to fine-tune the bare mass  $m_{H,0}^2$  up to 30 decimal places!

=> not natural

=> likely SM is incomplete...

