

Last time: we reviewed 4-vectors:

contravariant:

$$A'^{\mu} = \frac{\partial x'^{\mu}}{\partial x^{\nu}} A^{\nu}$$

covariant:

$$B'_{\mu} = \frac{\partial x^{\nu}}{\partial x'^{\mu}} B_{\nu}$$

defined scalar product

$$A_{\mu} B^{\mu}, \text{ metric } g_{\mu\nu} = \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & -1 & \\ & & & 1 \end{pmatrix}$$

$$A^{\mu} = g^{\mu\nu} A_{\nu}, \quad A_{\mu} = g_{\mu\nu} A^{\nu}, \quad g_{\mu\alpha} g^{\alpha\nu} = \delta_{\mu}^{\nu}$$

defined derivatives  $\partial_{\mu} = \frac{\partial}{\partial x^{\mu}}, \quad \partial^{\mu} = \frac{\partial}{\partial x_{\mu}}$

started talking about classical field theory

for Free Scalar Field:  $\varphi$  - scalar field

the action

$$S = \int d^4x \mathcal{L}(\varphi, \partial_{\mu}\varphi)$$

$\mathcal{L}$  lagrangian density

least action principle  $\delta S = 0 \Rightarrow$

$$\frac{\delta \mathcal{L}}{\delta \varphi} - \partial_{\mu} \frac{\delta \mathcal{L}}{\delta (\partial_{\mu}\varphi)} = 0$$

Euler-Lagrange equations

for massive scalar field:

$$\mathcal{L} = \frac{1}{2} \partial_{\mu}\varphi \partial^{\mu}\varphi - \frac{m^2}{2} \varphi^2$$

$\Rightarrow$  EOM is

$$[\partial_{\mu} \partial^{\mu} + m^2] \varphi = 0$$

Klein-Gordon equation

or  $[\square + m^2] \varphi = 0$

$$\Rightarrow \boxed{[\partial_\mu \partial^\mu + m^2] \varphi = 0}$$

Klein-Gordon equation

or  $\boxed{[\square + m^2] \varphi = 0}$

To solve K-G equation write  $\varphi(x) = \int d^4k e^{-ik \cdot x} \tilde{\varphi}(k)$

with  $k \cdot x = k_\mu x^\mu = k^0 x^0 - \vec{k} \cdot \vec{x}$ .

$$[\square + m^2] \varphi = \int d^4k \tilde{\varphi}(k) (\square + m^2) e^{-ik \cdot x} = \int d^4k \tilde{\varphi}(k) \cdot$$

$$[-k^2 + m^2] = 0 \quad \text{with } k^2 = k_\mu k^\mu = (k^0)^2 - (\vec{k})^2$$

$$\Rightarrow [k^2 - m^2] \tilde{\varphi} = 0 \Rightarrow \text{as } \tilde{\varphi} \neq 0 \Rightarrow k^2 = m^2 \text{ or}$$

$$E_k^2 - \vec{k}^2 = m^2 \Rightarrow E_k = \pm \sqrt{\vec{k}^2 + m^2} \Rightarrow \text{define } \boxed{E_k = \sqrt{\vec{k}^2 + m^2}}$$

$$\Rightarrow \boxed{\varphi(x) = \int \frac{d^3k}{(2\pi)^3 2E_k} \left[ a_{\vec{k}} e^{-iE_k t + i\vec{k} \cdot \vec{x}} + a_{\vec{k}}^* e^{iE_k t - i\vec{k} \cdot \vec{x}} \right]}$$

most general solution.

Canonical Quantization

In your QM class you must have seen that if we treat K-G equation as the equation for a single-particle wave function  $\varphi(x)$  (just like

Schroedinger eqn, except now relativistic), then we have a host of problems:

(i) as the energy =  $\pm \sqrt{\vec{k}^2 + m^2 c^2} \Rightarrow$  can have free particles with negative energy!?

(ii) particle propagation  $\langle \vec{x} | e^{-it\hat{H}} | \vec{y} \rangle$  from point  $\vec{y}$  to pt.  $\vec{x}$  is acausal ~ can find the particle outside the light-cone  $\Rightarrow$  it would propagate "faster" than light...

~ in general we know that relativistic kinematics allows for a particle to decay into several particles  $\Rightarrow$  particle # is not conserved  $\Rightarrow$  should not have a single-particle wave function interpretation.

$\Rightarrow$  we quantize the system treating  $\psi(x)$  as a field!

Again, let us draw an analogy with Quantum mechanics. Start with a system with degrees of freedom  $q_i$  described by Lagrangian  $L(q_i, \dot{q}_i)$

$i=1, \dots, N$ . Define canonical momenta  $p_i = \frac{\partial L}{\partial \dot{q}_i}$

⇒ get Hamiltonian  $H = \sum_i \dot{q}_i p_i - L \Rightarrow H(q_i, p_i)$

⇒ quantize by "promoting"  $q_i, p_i$  to operators

$\hat{q}_i, \hat{p}_i$  such that ( $[A, B] \equiv AB - BA$  ~ commutator)

$[ \hat{q}_i, \hat{p}_j ] = i \delta_{ij}, [ \hat{q}_i, \hat{q}_j ] = [ \hat{p}_i, \hat{p}_j ] = 0$

Mechanics

Field Theory

$q_i \rightarrow \varphi(x)$

$i \rightarrow x^\mu$

$p_i = \frac{\partial L}{\partial \dot{q}_i} \rightarrow \pi(x) = \frac{\delta \mathcal{L}}{\delta \dot{\varphi}(x)}$

$[ \hat{q}_i, \hat{p}_j ] = i \delta_{ij} \rightarrow [ \varphi(\vec{x}, t), \pi(\vec{x}', t) ] = i \delta(\vec{x} - \vec{x}')$

$[ \hat{q}_i, \hat{q}_j ] = 0 \rightarrow [ \varphi(\vec{x}, t), \varphi(\vec{x}', t) ] = 0$

$[ \hat{p}_i, \hat{p}_j ] = 0 \rightarrow [ \pi(\vec{x}, t), \pi(\vec{x}', t) ] = 0$

⇒ Note that canonical quantization favors time direction and is therefore not relat. inv. (physics is indeed invariant).

$\hat{H}(\hat{q}_i, \hat{p}_i) \rightarrow \hat{H} = \int d^3x \cdot \mathcal{H}(\varphi, \pi)$

Hamiltonian gives time-evolution of the system

We have :

$$[\varphi(\vec{x}, t), \pi(\vec{x}', t)] = i \delta(\vec{x} - \vec{x}')$$

$$[\varphi(\vec{x}, t), \varphi(\vec{x}', t)] = [\pi(\vec{x}, t), \pi(\vec{x}', t)] = 0$$

where  $\varphi, \pi$  are operators,  $\pi(x) \equiv \frac{\delta \mathcal{L}}{\delta \dot{\varphi}(x)}$

Write

$$\varphi(x) = \int \frac{d^3k}{(2\pi)^3 2\epsilon_k} \left[ \hat{a}_{\vec{k}} e^{-ik \cdot x} + \hat{a}_{\vec{k}}^\dagger e^{ik \cdot x} \right] \quad (\text{now } \hat{a}, \hat{a}^\dagger \text{ are operators})$$

$$\Rightarrow \left[ \hat{a}_{\vec{k}}, \hat{a}_{\vec{k}'}^\dagger \right] = (2\pi)^3 2\epsilon_k \delta(\vec{k} - \vec{k}')$$

$$\left[ \hat{a}_{\vec{k}}, \hat{a}_{\vec{k}'} \right] = \left[ \hat{a}_{\vec{k}}^\dagger, \hat{a}_{\vec{k}'}^\dagger \right] = 0$$

(can show)

*for  $\mathcal{L} = \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi + \dots$   
 $\pi = \partial_0 \varphi = \dot{\varphi}$*

The Hamiltonian is  $H = \int d^3x \left[ \dot{\varphi}(x) \pi(x) - \mathcal{L} \right]$

$$\Rightarrow i \frac{\partial}{\partial t} \hat{O} = [\hat{O}, \hat{H}]$$

Heisenberg picture.

$$\Rightarrow \hat{O}(\vec{x}, t) = e^{i\hat{H}t} \hat{O}(\vec{x}, 0) e^{-i\hat{H}t}$$

Example :  $\mathcal{L} = \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi - \frac{m^2}{2} \varphi^2 \Rightarrow$

$$\Rightarrow \pi(x) = \frac{\delta \mathcal{L}}{\delta \dot{\varphi}} = \frac{\delta \mathcal{L}}{\delta (\partial_0 \varphi)} = \partial_0 \varphi \Rightarrow$$

$$H = \int d^3x \left[ \dot{\varphi}^2 - \frac{1}{2} \dot{\varphi}^2 + \frac{1}{2} (\vec{\nabla} \varphi)^2 + \frac{m^2}{2} \varphi^2 \right]$$

define  $\overleftrightarrow{\partial}_0$  by  $\varphi_1 \overleftrightarrow{\partial}_0 \varphi_2 = \varphi_1 \partial_0 \varphi_2 - \varphi_2 \partial_0 \varphi_1$

Note that  $\int d^3x e^{i\vec{k}\cdot\vec{x}} \overleftrightarrow{\partial}_0 e^{-i\vec{k}'\cdot\vec{x}} = \int d^3x [-i\varepsilon_k - i\varepsilon_{k'}]$

$$e^{i\vec{x}\cdot(\vec{k}-\vec{k}')} = -2i\varepsilon_k (2\pi)^3 \delta(\vec{k}-\vec{k}')$$

$$\Rightarrow \text{if } \varphi(x) = \int \frac{d^3k}{(2\pi)^3 2\varepsilon_k} \left[ \hat{a}_{\vec{k}} e^{-i\vec{k}\cdot\vec{x}} + \hat{a}_{\vec{k}}^\dagger e^{i\vec{k}\cdot\vec{x}} \right]$$

$$\begin{aligned} \Rightarrow \int d^3x e^{i\vec{k}'\cdot\vec{x}} \overleftrightarrow{\partial}_0 \varphi(x) &= \int \frac{d^3k}{(2\pi)^3 2\varepsilon_k} \hat{a}_{\vec{k}} (-2i\varepsilon_k) (2\pi)^3 \delta(\vec{k}-\vec{k}') \\ &= -i \hat{a}_{\vec{k}'} \end{aligned}$$

$$\text{as } \int d^3x e^{i\vec{k}\cdot\vec{x}} \overleftrightarrow{\partial}_0 e^{+i\vec{k}'\cdot\vec{x}} = 0 \quad (\text{why?})$$

$$\Rightarrow \hat{a}_{\vec{k}} = \int d^3x e^{i\vec{k}\cdot\vec{x}} i \overleftrightarrow{\partial}_0 \varphi(x)$$

Similarly  $\hat{a}_{\vec{k}}^\dagger = \int d^3x \varphi(x) i \overleftrightarrow{\partial}_0 e^{-i\vec{k}\cdot\vec{x}}$

$$\Rightarrow [\hat{a}_{\vec{k}}, \hat{a}_{\vec{k}'}^\dagger] = \int d^3x d^3y \left[ e^{i\vec{k}\cdot\vec{x}} i \overleftrightarrow{\partial}_0 \varphi(x), \varphi(y) i \overleftrightarrow{\partial}_0 e^{-i\vec{k}'\cdot\vec{y}} \right]$$

$$= \int d^3x d^3y e^{i\vec{k}\cdot\vec{x} - i\vec{k}'\cdot\vec{y}} (-) \left[ \bar{n}(x) - i\varepsilon_k \varphi(x), -i\varepsilon_{k'} \varphi(y) - \bar{n}(y) \right]$$

$$= \int d^3x d^3y e^{i\vec{k}\cdot\vec{x} - i\vec{k}'\cdot\vec{y}} (-i) \left\{ [\varphi(x), \bar{n}(y)]_{\vec{k}, t} + [\varphi(y), \bar{n}(x)]_{\vec{k}', t} \right\} =$$

$$= \int d^3x d^3y e^{i\vec{k}\cdot\vec{x} - i\vec{k}'\cdot\vec{y}} (\varepsilon_k + \varepsilon_{k'}) \delta(\vec{x} - \vec{y}) = e^{i(\varepsilon_k - \varepsilon_{k'})t} (\varepsilon_k + \varepsilon_{k'}) (2\pi)^3 \delta(\vec{k} - \vec{k}')$$

$$= 2 \epsilon_u (2\vec{u})^3 \delta(\vec{u} - \vec{u}') \text{ as advertised!}$$

(15")

$$\Rightarrow H = \int d^3x \left[ \frac{1}{2} (\partial_0 \varphi)^2 + \frac{1}{2} (\vec{\nabla} \varphi)^2 + \frac{m^2}{2} \varphi^2 \right].$$

plug in  $\varphi = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\epsilon_k} \left[ \hat{a}_{\vec{k}} e^{-ik \cdot x} + \hat{a}_{\vec{k}}^\dagger e^{ik \cdot x} \right]$

to get (after some algebra)

$$H = \int \frac{d^3k}{(2\pi)^3} \frac{\epsilon_k}{2} \left( \hat{a}_{\vec{k}}^\dagger \hat{a}_{\vec{k}} + \hat{a}_{\vec{k}} \hat{a}_{\vec{k}}^\dagger \right)$$

$\Rightarrow$  defining particle number operator  $\hat{N}(\vec{k}) = \hat{a}_{\vec{k}}^\dagger \hat{a}_{\vec{k}}$

write  $H = \int \frac{d^3k}{(2\pi)^3} \frac{\epsilon_k}{2} \left[ \hat{N}(\vec{k}) + \infty \right]$

(zero point energy: consider harm. osc. in a box:  $E_n = \hbar\omega(n + \frac{1}{2}) \Rightarrow E_0 = \frac{1}{2} \hbar\omega = \frac{1}{2} \hbar \frac{2\pi m}{L} \Rightarrow \sum_m E_0 = \frac{\hbar}{L} \hbar \sum_{m=0}^{\infty} m = \infty$ )   
 a constant, zero-point energy, not detectable experimentally

$$\Rightarrow H \propto \epsilon_k \cdot N(\vec{k})$$

$\uparrow$  energies of particles   
 $\uparrow$  # of particles

$\sim$  makes physical sense!

### Free Dirac Field

Def. Dirac  $\gamma$ -matrices  $\gamma^\mu$  are defined by

requiring  $\{ \gamma^\mu, \gamma^\nu \} = 2g^{\mu\nu}$



In Weyl representation one has

$$\gamma^0 = \begin{pmatrix} 0 & \mathbb{1} \\ \mathbb{1} & 0 \end{pmatrix}, \quad \gamma^i = \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix}, \quad i=1,2,3$$

where  $\sigma^i$ 's are Pauli matrices

$$\sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

**Def.**  $\sigma_{\mu\nu} = \frac{i}{2} [\gamma_\mu, \gamma_\nu]$  One can show that these  $\sigma_{\mu\nu}$  matrices are generators of Lorentz algebra.

**Def.**  $\psi_\alpha(x)$ ,  $\alpha=1,2,3,4$  is called a spinor if it transforms as

$$\psi(x) \rightarrow \psi'(x') = e^{-\frac{i}{4} \omega^{\mu\nu} \sigma_{\mu\nu}} \psi(x)$$

under Lorentz transformation.

$$\psi = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix}$$

(E.g.  $\omega^{0i} = \zeta^i \sim$  boosts,  $\omega^{12} = \theta^3 \sim$  rotations, etc.)

**Def.**  $\bar{\psi}_\alpha = \psi^\dagger_\beta (\gamma_0)_{\beta\alpha}$

One can show that  $\bar{\psi}(x) \rightarrow \bar{\psi}'(x') = \bar{\psi}(x) e^{\frac{i}{4} \omega^{\mu\nu} \sigma_{\mu\nu}}$ .

Therefore  $\bar{\psi} \psi \equiv \sum_{\alpha=1}^4 \bar{\psi}_\alpha \psi_\alpha$  is a Lorentz-invariant!