

SYNCHROTRON FAR INFRARED SPECTROSCOPY: HIGHER RESOLUTION AND LONGER WAVELENGTHS AT THE CANADIAN LIGHT SOURCE

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In order to take full advantage of the high spectral resolution capability of the Bruker IFS125 Fourier transform spectrometer (MOPD = 9.4 m) at the Canadian Light Source far infrared beamline, we have interfaced a large 2 m gas cell which allows longer absorption paths, lower sample pressures, and hence reduced pressure broadening. We are using acrolein (CH₂CHCHO) for testing. Interesting in its own right, this molecule is also of importance in astrophysics, combustion chemistry, and human respiration (cigarette smoke, smog). Our initial CLS work^b on ν_{12} (~ 564 cm⁻¹) and ν_{17} (~ 593 cm⁻¹) has now been extended to ν_{18} (~ 158 cm⁻¹), $2\nu_{18}$ (~ 314 cm⁻¹), ν_{13} (~ 323 cm⁻¹), $3\nu_{18}$ (~ 469 cm⁻¹), and further vibrational levels.

We experience significant gains in source intensity and signal-to-noise ratio due to the brightness advantage of synchrotron radiation compared to conventional thermal far infrared sources, though source noise (due to mechanical vibrations) remains a concern. Even greater gains at long wavelengths (< 50 cm⁻¹) may be possible using coherent synchrotron radiation (CSR), achieved by shortening the electron bunch length in the synchrotron storage ring.^c Initial testing of CSR at CLS has already yielded a recognizable high-resolution spectrum of N₂O pure rotational transitions in the 10 - 20 cm⁻¹ range.

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^bA.R.W. McKellar, D.W. Tokaryk, L.-H. Xu, D.R.T. Appadoo, and T. May, *J. Mol. Spectrosc.* **241**, 31 (2007).

^cM. Abo-Bakr, J. Feikes, K. Hollmack, P. Kuske, W. B. Peatman, U. Schade, G. Wüstefeld, and H.-W. Hübers, *Phys. Rev. Letters* **90**, 094801 (2003).