

THE ROTATIONAL SPECTRA OF THE SILICON ISOTOPIC SPECIES OF SiCC

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Until this work, the rotational spectra of the silicon isotopic species of SiCC were based almost entirely on astronomical frequencies, because only the fundamental $1_{0,1} - 0_{0,0}$ transition $^{29}\text{SiC}_2$ and $^{30}\text{SiC}_2$ had been measured in the laboratory.^a We have now derived precise rotational and centrifugal distortion constants from laboratory measurements of 35 transitions of each isotopic species between 140 and 360 GHz with $J \leq 10$ and $K_a \leq 8$. The rotational spectra calculated with the laboratory measured constants are about two orders of magnitude more accurate than that of He *et al.*,^b who determined the spectroscopic constants from about 20 lines of $^{29}\text{SiC}_2$ and of $^{30}\text{SiC}_2$ in the wide-line source IRC+10216. The new laboratory measurements should aid assignment of the silicon isotopic species of SiCC in the spectral line survey of IRC+10216 with the SMA,^c and in future observations with ALMA.

^aR. D. Suenram, F. J. Lovas, and K. Matsumura, *Astrophys. Journ. Lett.* **342**, L103 (1989)

^bJ. H. He, Dinh-V-Trung, S. Kwok, H. S. P. Müller, Y. Zhang, T. Hasegawa, T. C. Peng, and Y. C. Huang, *Astrophys. Journ. Suppl. Ser.*, **177**, 275 (2008).

^cN. A. Patel, K. H. Young, S. Brünken, R. W. Wilson, P. Thaddeus, K. M. Menten, M. Reid, M. C. McCarthy, Dinh-V-Trung, C. A. Gottlieb, and A. Hedden, *Astrophys. Journ.*, in press (2009).