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The S-Eqn in First and Second Quantization

- see FTW, sect. 1

- The time-dependent Schrodinger equation is

$$i\hbar \frac{\partial}{\partial t} \Psi(x_1, \dots, x_N, t) = H \Psi(x_1, \dots, x_N, t)$$

where $x_k \equiv \{x_k, s_{z_k}, \dots\}$ specifies fully the k^{th} particle.
With an appropriate set of boundary conditions and

$$H = \sum_{k=1}^N T(x_k) + \frac{1}{2} \sum_{k \neq l=1}^N V(x_k, x_l)$$

we have our problem to solve.

- We start by expanding Ψ in a complete set of time-independent, single-particle wave functions that incorporate the boundary conditions.

- eg. large, uniform system, expand in plane waves in a large box with periodic boundary conditions
- in harmonic trap, expand in harmonic oscillator wf's
- in atoms, expand in complete set of Coulomb wf's

- Let $\psi_{E_k}(x_k)$ be the single particle wave function, where E_k is a complete set of single-particle quantum numbers
 - spinless bosons in a box: $E_k \rightarrow \vec{p}$
 - " " in central field: $E_k \rightarrow E, J, m$
 - fermions in a box: $E_k \rightarrow \vec{p}, s_z$

- Imagine that this infinite set E_k is ordered:
(1, 2, 3, ..., $\vec{p}, s, t, \dots, \infty$)

- ASSUME THE $\psi_{E_k}(x_k)$ ARE ORTHONORMAL.

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Consider 1 particle P1st, with $U(x_2)$ instead of $V(x_1, x_2)$

$$\Psi(x_2, t) = \sum_{E_2} C(E_2, t) \psi_{E_2}(x_2)$$

• eg. expanding in harmonic oscillator wf.'s, so E_2 runs over $n=0, l=0, n=1, l=0, n=2, l=0, n=0, l=1$, and so on.

• Let's find an equation for the C's from the S-eqn. Plug in:

$$i\hbar \sum_{E_2} \left[\frac{\partial}{\partial t} C(E_2, t) \right] \psi_{E_2}(x_2) = \sum_{E_2} C(E_2, t) H \psi_{E_2}(x_2)$$

• Use orthonormality of ψ_{E_2} to project out $C(E_2, t)$ out of all the E_2 terms. Do this by multiplying by $\psi_{E_2}^*(x_2)$ and integrating over x_2 (which may mean a sum over spin) (so that spinor is correct)

$$\int dx_2 \psi_{E_2}^*(x_2) \psi_{E_1}(x_2) = \sum_{E_2} \delta_{E_2 E_1} = \begin{cases} 0 & \text{if } E_2 \neq E_1 \\ 1 & \text{if } E_2 = E_1 \end{cases}$$

$$\Rightarrow i\hbar \frac{\partial}{\partial t} C(E_2, t) = \sum_{E_1} \underbrace{\int dx_2 \psi_{E_2}^*(x_2) [T(x_2) + U(x_2)] \psi_{E_1}(x_2)}_{\text{Hamiltonian matrix element} \rightarrow \text{number}} C(E_1, t)$$

• We get an equation for each $E_2 \Rightarrow$ infinite set of coupled differential equations for the C's.

• IF we want a stationary state, then time dependence of all C's is proportional to $e^{-iEt/\hbar}$, so we get a matrix eigenvalue problem

• in practice, truncate basis to get finite matrix to diagonalize

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Now suppose we have two (identical) particles, with $V(x_1, x_2)$ again.

Expand $\Psi(x_1, x_2, t) = \sum_{E_1', E_2'} C(E_1', E_2', t) \psi_{E_1'}(x_1) \psi_{E_2'}(x_2)$

\Rightarrow complete basis for Ψ .

Plug into S-eqn again, multiply by $\psi_{E_1'}^+(x_1) \psi_{E_2'}^+(x_2)$ and integrate over x_1 and x_2 .

$$i\hbar \frac{d}{dt} C(E_1', E_2', t) = \sum_{E_1'', E_2''} \int dx_1 dx_2 \psi_{E_1''}^+(x_1) \psi_{E_2''}^+(x_2) \left[T(x_1) + T(x_2) + \frac{1}{2} V(x_1, x_2) + \frac{1}{2} V(x_2, x_1) \right] \psi_{E_1'}(x_1) \psi_{E_2'}(x_2) \times C(E_1'', E_2'', t)$$

\Rightarrow Kinetic energy T involves only one E_k at a time, while V involves two E_k 's (since two-body operator).

If a coordinate x_k is not "mentioned" by T or V in a given term, then orthonormality just sets $E_k' = E_k$

$$\Rightarrow i\hbar \frac{d}{dt} C(E_1', E_2', t) = \int dx_1 \psi_{E_1'}^+(x_1) T(x_1) \psi_{E_1'}(x_1) C(E_1', E_2', t) + \int dx_2 \psi_{E_2'}^+(x_2) T(x_2) \psi_{E_2'}(x_2) C(E_1', E_2', t) + \frac{1}{2} \int dx_1 dx_2 \psi_{E_1'}^+(x_1) \psi_{E_2'}^+(x_2) V(x_1, x_2) \psi_{E_1'}(x_1) \psi_{E_2'}(x_2) \times C(E_1', E_2', t) + (1 \leftrightarrow 2)$$

So, again, we have an infinite set of coupled differential equations. Stationary states \Rightarrow just a bigger matrix problem.

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Now N-body: same thing:

$$\Psi(x_1, \dots, x_N, t) = \sum_{E_1, \dots, E_N} C(E_1 \dots E_N, t) \psi_{E_1}(x_1) \dots \psi_{E_N}(x_N)$$

• Multiplying by $\psi_{E_1}^+(x_1) \dots \psi_{E_N}^+(x_N)$ and integrating, orthonormality eliminates all but one integration in each term in the sum of kinetic energy terms and all but 2 integrations in the potential energy sum.

• We say that only one sp. wavefunction is "changed" in the first case and only 2 in the second case:
↑ "single particle"

$$i\hbar \frac{\partial}{\partial t} C(E_1, \dots, E_N, t) = \sum_{k=1}^N \sum_{W'} \int dx_k \psi_{E_k}^+(x_k) T(x_k) \psi_{W'}(x_k) C(E_1, \dots, E_{k-1}, E_{k+1}, \dots, E_N, t) + \frac{1}{2} \sum_{k \neq l=1}^N \sum_{W, W'} \iint dx_k dx_l \psi_{E_k}^+(x_k) \psi_{E_l}^+(x_l) V(x_k, x_l) \psi_{W'}(x_k) \psi_{W''}(x_l) C(E_1, \dots, E_{k-1}, E_{k+1}, \dots, E_{l-1}, E_{l+1}, \dots, E_N, t)$$

• So this "works", but you can see the practical difficulties if we have 100 or 10^{23} particles!
 • Big waste of effort, since only 1 or 2 change

• What about the fact that the particles are identical?
 • it means that lots of the C's must be the same!

$$\Psi(\dots x_i \dots x_j \dots, t) = \pm \Psi(\dots x_j \dots x_i \dots, t)$$

for bosons/fermions.

This implies

$$C(\dots E_i \dots E_j \dots, t) = \pm C(\dots E_j \dots E_i \dots, t)$$

So with two particles, $C(E_2, E_2) = +C(E_2, E_1)$ for bosons
 $C(E_1, E_2) = -C(E_1, E_2)$ for fermions (=0 if $E_1 = E_2$)

So lots of redundancy. ↑ Pauli
 means basis states are Slater determinants

• so "real" basis for bosons is $\psi_1(x_1)\psi_2(x_2), \psi_1(x_1)\psi_1(x_2) + \psi_2(x_1)\psi_2(x_2), \dots$
 for fermions $\psi_1(x_1)\psi_2(x_2) - \psi_2(x_1)\psi_1(x_2), \dots$
 \Rightarrow Slater determinants

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• We can regroup the equations and sums by just counting how many times E_1 appears, E_2 appears, and so on.

• Since identical, doesn't matter if particle 1 or particle 2 or whatever

\Rightarrow Keep track of "occupation number" n_k of E_k

• In gory detail one can show how one can rewrite the results in terms of occupation number basis states. We refer the reader to F+W and just jump to the answer without proof.

• Introduce time-independent abstract state vectors:

$$|n_1 n_2 \dots n_{\infty}\rangle$$

where n_1 particles in single-particle eigenstate 1, n_2 in eigenstate 2 and so on.

• Make it complete and orthonormal

$$\langle n'_1 n'_2 \dots n'_{\infty} | n_1 n_2 \dots n_{\infty} \rangle = \delta_{n_1 n'_1} \dots \delta_{n_{\infty} n'_{\infty}}$$

$$\sum_{n_1, n_2, \dots, n_{\infty}} |n_1 n_2 \dots n_{\infty}\rangle \langle n_1 n_2 \dots n_{\infty}| = 1$$

- for bosons n_k goes from 0 to ∞ in the complete space.
- for fermions n_k is either 0 or 1

• In fact, $|n_1 n_2 \dots n_{\infty}\rangle = |n_1\rangle |n_2\rangle \dots |n_{\infty}\rangle$

since (as seen below) the eigenstates can be simultaneous eigenstates of each of the number operators for individual modes.

• for each mode n_i , think about harmonic oscillator.

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To work with this basis, we recall the raising and lowering operators of the harmonic oscillator, and for bosons introduce b_k, b_k^\dagger that satisfy

$$\boxed{[b_k, b_{k'}^\dagger] = \delta_{kk'} \quad \text{bosons}}$$

$$\boxed{[b_k, b_{k'}] = [b_k^\dagger, b_{k'}^\dagger] = 0}$$

As in HO case, the properties of these operators follow from the commutation relations:

$$\boxed{b_k^\dagger b_k |n_k\rangle = n_k |n_k\rangle} \quad n_k = 0, 1, 2, \dots, \infty \quad \text{"number operator"}$$

$$\boxed{b_k |n_k\rangle = (n_k)^{1/2} |n_k - 1\rangle} \quad \text{"destruction operator"}$$

$$\boxed{b_k^\dagger |n_k\rangle = (n_k + 1)^{1/2} |n_k + 1\rangle} \quad \text{"creation operator"}$$

Proofs:

← hermitian \Rightarrow real eigenvalues

Since $n = \langle n | b^\dagger b | n \rangle = \sum_m \langle n | b^\dagger | m \rangle \langle m | b | n \rangle = \sum_m | \langle m | b | n \rangle |^2 > 0$, we can't have a negative eigenvalue of the number operator.

So use b to "count down" from $|n\rangle$: ← from commutation relations

$$b^\dagger b (b | n \rangle) = b (b^\dagger b | n \rangle) + [b^\dagger b, b] | n \rangle = (n-1) (b | n \rangle)$$

which must terminate in zero eventually if repeated to avoid going negative. The same argument applying b^\dagger establishes the first and we get the normalization (up to a phase chosen to be 1) by combining the last two equations.

The extension to many modes is immediate since $[b_k^\dagger b_k, b_{k'}^\dagger b_{k'}] = 0$ for $k \neq k'$ \Rightarrow eigenstates of total system can be simultaneous number operator eigenstates.

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Now here's the big claim: We can exactly rewrite the many-body S-equation as

$$i\hbar \frac{d}{dt} |\Psi(t)\rangle = \hat{H} |\Psi(t)\rangle$$

with

$$\hat{H} = \sum_{ij} b_i^\dagger \langle i|T|j\rangle b_j + \frac{1}{2} \sum_{ijkl} b_i^\dagger b_j^\dagger \langle ij|V|kl\rangle b_l b_k$$

- The b_i 's and b_i^\dagger 's take care of the symmetrization for us
- kinetic term \Rightarrow destroy in state j , create in state i (\Rightarrow "change" j to i as before)
- potential term \Rightarrow 2 states changed by 1.

- Could expand $|\Psi(t)\rangle$ in coefficients $f(n_1, n_2, \dots, n_{\infty}, t) |n_1 \dots n_{\infty}\rangle$ and show we get the same answer as before. We'll defer that!
 \Rightarrow See Fetter and Walecka, section 4.

- Note that the matrix elements $\langle i|T|j\rangle$ and $\langle ij|V|kl\rangle$ are exactly the same sort we had in the equations for the C 's.
 • The b 's just do the accounting for us now!

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For fermions we need to use anticommutation relations

$$\{a_r, a_s^\dagger\} = \delta_{rs} \text{ fermions}$$

$$\{a_r, a_s\} = \{a_r^\dagger, a_s^\dagger\} = 0$$

$$\{A, B\} \equiv [A, B]_{\pm} \equiv AB \mp BA$$

1. $a_s^2 = a_s^{\dagger 2} = 0$ so $a_s^\dagger a_s^\dagger |0\rangle = 0 \Rightarrow$ no two particles in same state

2. $\hat{n}_s = a_s^\dagger a_s$ is the number operator

3. $a^\dagger |0\rangle = |1\rangle$ $a |1\rangle = |0\rangle$ (overall choice of phase has been made)
 $a^\dagger |1\rangle = 0$ $a |0\rangle = 0$

4. $|n_1 \dots n_\infty\rangle \equiv (a_1^\dagger)^{n_1} \dots (a_\infty^\dagger)^{n_\infty} |0\rangle$
 implies signs. Define $S_s = n_1 + n_2 + \dots + n_{s-1}$

← just counts how many a^\dagger 's a must be commuted with to get to its proper place.

$$a_s | \dots n_s \dots \rangle = \begin{cases} 0 & \text{if } n_s = 0 \\ (-1)^{S_s} | \dots n_s - 1 \dots \rangle & \text{if } n_s = 1 \end{cases}$$

$$a_s^\dagger | \dots n_s \dots \rangle = \begin{cases} (-1)^{S_s} (n_s + 1)^{1/2} | \dots n_s + 1 \dots \rangle & \text{if } n_s = 0 \\ 0 & \text{if } n_s = 1 \end{cases}$$

$$* a_s^\dagger a_s | \dots n_s \dots \rangle = n_s | \dots n_s \dots \rangle \quad n_s = 0, 1$$

$$5. \quad i\hbar \frac{\partial}{\partial t} |\Psi(t)\rangle = \hat{H} |\Psi(t)\rangle$$

$$\hat{H} = \sum_{rs} a_r^\dagger \langle r|T|s\rangle a_s + \frac{1}{2} \sum_{rstu} a_r^\dagger a_s^\dagger \langle rs|V|tu\rangle a_u a_t$$

↑ note ordering

Alternative example: degenerate electron gas (see F1W)
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Example: Dilute Fermi Gas with Short-Range Interaction

- As a prototype system we can consider a set of atoms confined magnetically or optically (the difference is important; see below) where we can consider the trap potential constant (therefore constant energy contribution). Very big trap, so treat as effectively infinite.

$$\Rightarrow V(\vec{x}-\vec{x}') = \lambda \delta^3(\vec{x}-\vec{x}')$$

(recall $\int d^3x' \delta^3(\vec{x}-\vec{x}') f(\vec{x}') = f(\vec{x})$)
so integrals are easy

- spin-independent (same matrix element between any combination of spin up and spin down)
- if $\lambda > 0 \Rightarrow$ repulsive, $\lambda < 0 \Rightarrow$ attractive
- $g \equiv 2S+1$ spin states in general

Goal: Find ground state energy per particle as a function of density. Also find the pressure and consider stability.

Plan: Do perturbation theory about non-interacting system

To treat a bulk medium, put the system in a large box of side L and take $L \rightarrow \infty$ at the end.

Uniform, infinite \Rightarrow translationally invariant so apply periodic boundary conditions on the single-particle

wave functions \Rightarrow plane waves (momentum eigenstates)

$$\psi_{\vec{k}\alpha}(\vec{x}) = \frac{1}{\sqrt{V}} e^{i\vec{k}\cdot\vec{x}} \eta_{\alpha}$$

where $V \equiv L^3$ is the volume of the box and η_{α} is the spin function. For spin-1/2 along a given z -axis,

$$\eta_{\uparrow} = \begin{bmatrix} 1 \\ 0 \end{bmatrix} \quad \text{and} \quad \eta_{\downarrow} = \begin{bmatrix} 0 \\ 1 \end{bmatrix} \quad \Rightarrow \quad \eta_{\alpha}^{\dagger} \eta_{\beta} = \delta_{\alpha\beta}$$

Periodic boundary conditions imply

$$k_i = \frac{2\pi n_i}{L} \quad i=x,y,z \quad n_i = 0, \pm 1, \pm 2$$

$\alpha = 0$ or 1
and $\beta = 0$ or 1

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Note that $\int \psi_{\vec{k}_\alpha}^\dagger(\vec{x}) \psi_{\vec{k}_\alpha}(\vec{x}) d^3x = \frac{1}{\Omega} \int e^{-i(\vec{k}-\vec{k})\cdot\vec{x}} \underbrace{\eta_{\vec{k}_\alpha}^\dagger \eta_{\vec{k}_\alpha}}_{=1} d^3x = \frac{1}{\Omega} \int d^3x = 1 \Rightarrow \text{normalized}$

Hamiltonian is (with $\hat{p}_i \rightarrow \hbar/i \nabla$)

$$H = \sum_{i=1}^N \frac{\hat{p}_i^2}{2m} + \frac{1}{\Omega} \sum_{i \neq j}^N V(\vec{x}_i - \vec{x}_j)$$

[Warning: As we'll note later this is ill-defined without a regularization/renormalization scheme]

Write this in 2nd quantization

$$\begin{aligned} \langle \vec{k}_1 \alpha_1 | \frac{\hat{p}_1^2}{2m} | \vec{k}_2 \alpha_2 \rangle &= \frac{1}{\Omega m^2} \int d^3x e^{-i\vec{k}_1 \cdot \vec{x}} \eta_{\alpha_1}^\dagger (-\hbar^2 \nabla^2) e^{i\vec{k}_2 \cdot \vec{x}} \eta_{\alpha_2} \\ &= \frac{\hbar^2 k_1^2}{2m \Omega} \delta_{\alpha_1 \alpha_2} \int d^3x e^{i(\vec{k}_2 - \vec{k}_1) \cdot \vec{x}} \underbrace{\int d^3k \delta(\vec{k} - \vec{k}_2)}_{\left(\text{to see this, integrate in Cartesian coordinates from } -\frac{1}{2} \leq x \leq \frac{1}{2} \text{ or } 0 \leq x \leq L \right)} \\ &= \frac{\hbar^2 k_1^2}{2m} \delta_{\alpha_1 \alpha_2} \delta_{\vec{k}_1 \vec{k}_2} \end{aligned}$$

kinetic energy \Rightarrow

$$\hat{T} = \sum_{\vec{k}\alpha} \frac{\hbar^2 k^2}{2m} a_{\vec{k}\alpha}^\dagger a_{\vec{k}\alpha}$$

← same for any V

- add up the kinetic energy of each occupied mode, as checked by the number operator $a_{\vec{k}\alpha}^\dagger a_{\vec{k}\alpha}$
- simple since diagonal in this basis

The potential energy is

← from WF's

$$\begin{aligned} \langle \vec{k}_1 \alpha_1, \vec{k}_2 \alpha_2 | V | \vec{k}_3 \alpha_3, \vec{k}_4 \alpha_4 \rangle &= \frac{1}{\Omega^2} \iint d^3x_1 d^3x_2 e^{-i\vec{k}_1 \cdot \vec{x}_1} \eta_{\alpha_1}(\vec{x}_1) e^{-i\vec{k}_2 \cdot \vec{x}_2} \eta_{\alpha_2}(\vec{x}_2) \\ &\quad \times V(\vec{x}_1 - \vec{x}_2) e^{i\vec{k}_3 \cdot \vec{x}_1} \eta_{\alpha_3}(\vec{x}_1) e^{i\vec{k}_4 \cdot \vec{x}_2} \eta_{\alpha_4}(\vec{x}_2) \\ &= \frac{1}{\Omega^2} \int d^3x_1 e^{-i(\vec{k}_1 + \vec{k}_3 - \vec{k}_2 - \vec{k}_4) \cdot \vec{x}_1} \delta_{\alpha_1 \alpha_3} \delta_{\alpha_2 \alpha_4} \\ &= \frac{1}{\Omega} \delta_{\alpha_1 \alpha_3} \delta_{\alpha_2 \alpha_4} \int d^3k \delta(\vec{k}_1 + \vec{k}_3 - \vec{k}_2 - \vec{k}_4) \end{aligned}$$

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So now the 2nd quantized Hamiltonian \hat{H} is

$$\hat{H} = \sum_{\vec{k}, \alpha} \frac{\hbar^2 k^2}{2m} a_{\vec{k}, \alpha}^\dagger a_{\vec{k}, \alpha} + \frac{\lambda}{2\Omega} \sum_{\vec{k}, \alpha_1} \sum_{\vec{k}', \alpha_2} \sum_{\vec{k}_3, \alpha_3} \sum_{\vec{k}_4, \alpha_4} a_{\vec{k}, \alpha_1}^\dagger a_{\vec{k}', \alpha_2}^\dagger a_{\vec{k}_3, \alpha_3} a_{\vec{k}_4, \alpha_4} \times \delta_{\alpha_1, \alpha_3} \delta_{\alpha_2, \alpha_4} \delta_{\vec{k} + \vec{k}', \vec{k}_3 + \vec{k}_4}$$

- The Kronecker delta's let us do two spin summations and one of the momentum summations (momentum conservation)
- We have many possibilities for choosing three independent momenta — until you get some experience you have no way of knowing what choice is best, but you will find in most cases that a particular choice can simplify your life. (this \vec{k}, \vec{p} in, exchange \vec{k} , so $\vec{k} + \vec{q}, \vec{p} - \vec{q}$ out.)
- Here we'll just follow Fetter+Walecka's (F+W's) lead and change to

$$\begin{aligned} \vec{k}_1 &= \vec{k} + \vec{q} & \vec{k}_3 &= \vec{k} \\ \vec{k}_2 &= \vec{p} - \vec{q} & \vec{k}_4 &= \vec{p} \end{aligned}$$

so \vec{q} is a momentum transfer.

$$\Rightarrow \hat{H} = \sum_{\vec{k}, \alpha} \frac{\hbar^2 k^2}{2m} a_{\vec{k}, \alpha}^\dagger a_{\vec{k}, \alpha} + \frac{\lambda}{2\Omega} \sum_{\vec{k}, \vec{q}} \sum_{\alpha_1, \alpha_2} a_{\vec{k} + \vec{q}, \alpha_1}^\dagger a_{\vec{p} - \vec{q}, \alpha_2}^\dagger a_{\vec{p}, \alpha_2} a_{\vec{k}, \alpha_1}$$

- In principle, the a 's and a^\dagger 's should have \wedge 's, since they are operators, but that gets too messy!
- It will be assumed that we have N particles in the volume Ω , so the density is $\rho \equiv N/\Omega$. (F+W: $n \equiv N/V$) We keep ρ fixed in the implied limit $N \rightarrow \infty, \Omega \rightarrow \infty$.

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We can see what to expect about the relative contributions of the two terms in \hat{H} by introducing dimensionless variables. This is a good way to analyze the expected size of different parts of the energy.

One length scale comes from the volume per particle, which we use to define the radius r_0 of a sphere equal to that volume:

$$\boxed{\frac{\text{volume}}{\text{particle}} = \frac{\Omega}{N} \equiv \frac{4}{3}\pi r_0^3} \Rightarrow r_0 \text{ is more-or-less the interparticle spacing,}$$

Another length scale for comparison comes from the interaction strength λ . We'll find out later that

$$\boxed{\lambda = \frac{4\pi\hbar^2 a_s}{m}}$$

$k \cot \delta_0 = -\frac{1}{a_s} + O(k^2)$
 \swarrow for small k where δ_0 is the s-wave phase shift.

where a_s is the s-wave "scattering length." (We'll come back to it!)

So the dimensionless ratio $\tilde{r}_s \equiv \frac{r_0}{a_s}$ (scaled particle separation) characterizes the density of the system:

- $\tilde{r}_s \rightarrow 0 \Rightarrow$ high density ($r_0 \rightarrow 0$)
- $\tilde{r}_s \rightarrow \infty \Rightarrow$ low density

If we rescale the volume and momenta by r_0 :

$$\boxed{\Omega = \Omega/r_0^3 \quad \vec{K} = r_0 \vec{K} \quad \vec{p} = r_0 \vec{p} \quad \vec{q} = r_0 \vec{q}}$$

Then we can rewrite \hat{H} as

$$\boxed{\hat{H} = \frac{\hbar^2}{m a_s^2} \left(\sum_{\vec{K}} \frac{1}{2} \vec{K}^2 a_{\vec{K}\downarrow}^\dagger a_{\vec{K}\downarrow} + \frac{a_s}{r_0} \frac{4\pi}{\Omega} \sum_{\vec{K}, \vec{q}} \sum_{\alpha, \alpha_2} a_{\vec{K}+\vec{q}, \alpha}^\dagger a_{\vec{p}=\vec{q}, \alpha_2}^\dagger a_{\vec{p}\alpha_2} a_{\vec{K}\alpha} \right)}$$

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 $(\tilde{r}_s \rightarrow \infty)$

- So at low density, the potential energy becomes a perturbation to the kinetic energy

\Rightarrow if $a_s \ll r_0$, apply perturbation theory to find the energy

- You might expect that the ground state energy has a power-series expansion in a_s , but we'll find, in fact, that it doesn't (stay tuned!)

- So let us assume from now on that we are in the dilute (low-density) limit

Separate $\hat{H} = \hat{H}_0 + \hat{H}_1$ where

$\hat{H}_0 = \frac{\hbar^2 k^2}{2m} \sum_{\vec{k}\alpha} a_{\vec{k}\alpha}^\dagger a_{\vec{k}\alpha}$ is the unperturbed Hamiltonian, which represents a noninteracting Fermi system

$$\hat{H}_1 = \frac{\lambda}{2L} \sum_{\vec{p}\vec{q}} \sum_{\alpha\alpha_2} a_{\vec{p}\vec{q}\alpha_1}^\dagger a_{\vec{p}-\vec{q}\alpha_2}^\dagger a_{\vec{p}\alpha_2} a_{\vec{q}\alpha_1}$$

is the small perturbation.

- Expand the ground-state energy E as $E^{(0)} + E^{(1)} + \dots$ where $E^{(0)}$ is the ground-state energy of a free Fermi gas while $E^{(1)}$ is the first-order energy shift

- In carrying out the evaluation of $E^{(0)}$ and $E^{(1)}$ we'll evaluate sums by converting to integrals (in the $L \rightarrow \infty$ limit)

$$\int F(k) dk \doteq \sum_i F(k_i) \Delta k_i = \sum_i F(k_i) \frac{2\pi}{L} \Delta i = \frac{2\pi}{L} \sum_i F(k_i)$$

$$\Rightarrow \sum_{\vec{k}\alpha} F_\alpha(\vec{k}) = \int \sum_\alpha \int \frac{d^3k}{(2\pi)^3} F_\alpha(\vec{k}) \quad [dk \text{ always comes with } 2\pi\text{'s}]$$

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- We denote the ground state of the non-interacting Fermi gas by $|F\rangle$.
- If spin $1/2$, then Pauli exclusion principle allows 2 fermions per momentum state (\uparrow and \downarrow)
- If spin S , then $g = 2S+1$ per momentum state
- Lowest energy by minimizing kinetic energy \Rightarrow fill lowest momentum states
- Call the highest filled momentum the Fermi momentum $p_F = \hbar k_F$

• Find k_F in terms of N, Ω , and $g = N/\Omega$ by taking the expectation value of the number operator $\leftarrow \theta(x) = \begin{cases} 1 & \text{if } x \geq 0 \\ 0 & \text{if } x < 0 \end{cases}$

$$N = \langle F | \hat{N} | F \rangle = \sum_{\mathbf{k}} \langle F | \hat{n}_{\mathbf{k}} | F \rangle = \sum_{\mathbf{k}} \theta(k_F - k)$$

$$= \frac{\Omega}{(2\pi)^3} \int d^3k \theta(k_F - k) = \frac{\Omega}{(2\pi)^3} g \times (\text{volume of sphere of radius } k_F)$$

$$= \frac{\Omega}{(2\pi)^3} g \frac{4}{3} \pi k_F^3 = \Omega \left(\frac{g k_F^3}{6\pi^2} \right)$$

$$\Rightarrow \boxed{g = \frac{9k_F^3}{6\pi^2}} \quad \text{and} \quad \boxed{k_F = \left(\frac{6\pi^2 g}{9} \right)^{1/3} = \left(\frac{18\pi^2}{4g} \right)^{1/3} \frac{1}{r_0} \rightarrow \frac{1.92}{r_0}}$$

so k_F^{-1} is comparable to the interparticle spacing.

- Also, our small parameter $\frac{a_0}{r_0}$ can be replaced by an alternative small parameter: $k_F a_0$.
- This is the usual expansion parameter for a dilute Fermi system.

1/8/03

So now evaluate $E^{(0)} \rightarrow$ expectation value of \hat{H}_0 in $|F\rangle$

$$\begin{aligned} E^{(0)} &= \langle F | \hat{H}_0 | F \rangle = \frac{\hbar^2}{2m} \sum_{\mathbf{k}\alpha} k^2 \langle F | \hat{n}_{\mathbf{k}\alpha} | F \rangle \\ &= \frac{\hbar^2}{2m} \sum_{\mathbf{k}\alpha} k^2 \theta(k_f - k) \quad \leftarrow \hat{n}_{\mathbf{k}\alpha} | F \rangle = \begin{cases} 0 & \text{if } |\mathbf{k}| > k_f \\ |F\rangle & \text{if } |\mathbf{k}| < k_f \end{cases} \\ &= \frac{\hbar^2}{2m} \sum_{\alpha} \frac{\Omega}{(2\pi)^3} \int d^3k k^2 \theta(k_f - k) \\ &= \frac{\hbar^2}{2m} \frac{9\sqrt{2}}{9\pi^2} \frac{k_f^5}{5} = \frac{\hbar^2 k_f^5}{2m} \frac{3}{5} \frac{9k_f^3}{6\pi^2} \Omega \end{aligned}$$

$$\begin{aligned} \frac{E^{(0)}}{N} &= \frac{3}{5} \frac{\hbar^2 k_f^5}{2m} \frac{1}{N} \\ &= \frac{2}{5} E_F^0 \end{aligned}$$

$= \frac{3}{5} \frac{\hbar^2 k_f^5}{2m} \frac{1}{N}$
 Fermi energy $E_F^0 = \frac{\hbar^2 k_f^2}{2m}$ ← energy of last atom at top of Fermi sea

• So in a free Fermi gas, the kinetic energy per particle is $\frac{2}{5}$ the kinetic energy of the last particle.

• The first-order energy shift is

$$E^{(1)} = \langle F | \hat{H}_1 | F \rangle = \frac{\lambda}{\Omega} \sum_{\mathbf{k}, \mathbf{q}, \alpha} \sum_{\alpha_1, \alpha_2} \langle F | a_{\mathbf{k}+\mathbf{q}, \alpha_1}^\dagger a_{\mathbf{q}, \alpha_2}^\dagger a_{\mathbf{k}, \alpha} a_{\mathbf{k}, \alpha} | F \rangle$$

- Ok, we'll analyze it by hand
- $a_{\mathbf{k}, \alpha_1}$ and $a_{\mathbf{q}, \alpha_2}$ annihilate particles in $|F\rangle$. If there is no particle to destroy, the result is zero. $\Rightarrow |\mathbf{k}|, |\mathbf{q}| < k_f$ or zero
- acting to the left, we similarly conclude $|\mathbf{k}+\mathbf{q}|, |\mathbf{k}| < k_f$
- The states must match up, or we'll get zero by orthogonality.

$$\Rightarrow \mathbf{k}+\mathbf{q}, \alpha_1 = \mathbf{k}, \alpha_2 \quad \text{and} \quad \mathbf{q}, \alpha_2 = \mathbf{k}, \alpha_1$$

$$\mathbf{k}+\mathbf{q}, \alpha_1 = \mathbf{q}, \alpha_2 \quad \text{and} \quad \mathbf{q}, \alpha_2 = \mathbf{k}, \alpha_1$$

} two terms call them "direct" and "exchange"

1/8/03

direct: The first pairing forces $\vec{q} = 0$ but doesn't constrain \vec{k} or \vec{p} (except we know they are less than k_F in magnitude) or either spin sum. (So the energy will be proportional to g^2 .)

exchange: The second pairing determines \vec{p} in terms of \vec{k} and \vec{q} (you can pick any one of the three momenta), eliminating the \vec{p} sum. Also, we must have $\alpha = \alpha_2$, eliminating one spin sum (so this part of the energy will be proportional to g .)

• Work out each one in turn, using $\hat{n}_{\vec{k}\alpha} |F\rangle = \theta(k_F - |\vec{k}|) |F\rangle$ repeatedly,

direct: $\langle F | a^\dagger a^\dagger a a | F \rangle \rightarrow \sum_{\vec{q}, \alpha_1, \alpha_2} \langle F | a^\dagger_{\vec{k}\alpha_1} a^\dagger_{\vec{p}\alpha_2} a_{\vec{p}\alpha_2} a_{\vec{k}\alpha_1} | F \rangle$

• Note that $\{\vec{p}\alpha_2\} \neq \{\vec{k}\alpha_1\}$, since otherwise the product of the a 's (or a^\dagger 's) would be zero.

\Rightarrow we can move $a_{\vec{k}\alpha_1}$ two places to the left without picking up extra terms from the anticommutators:

$$a^\dagger_{\vec{p}\alpha_2} a_{\vec{p}\alpha_2} a_{\vec{k}\alpha_1} = -a^\dagger_{\vec{p}\alpha_2} a_{\vec{k}\alpha_1} a_{\vec{p}\alpha_2} = +a_{\vec{k}\alpha_1} a^\dagger_{\vec{p}\alpha_2} a_{\vec{p}\alpha_2}$$

$\Rightarrow \langle F | a^\dagger a^\dagger a a | F \rangle \rightarrow \sum_{\vec{q}, \alpha_1, \alpha_2} \langle F | \hat{n}_{\vec{k}\alpha_1} \hat{n}_{\vec{p}\alpha_2} | F \rangle = \theta(k_F - |\vec{k}|) \theta(k_F - |\vec{p}|) \sum_{\alpha_1, \alpha_2}$

$$\begin{aligned} \Rightarrow E_{\text{direct}}^{(1)} &= \frac{1}{2\Omega} \sum_{\alpha_1, \alpha_2} \sum_{\vec{k}, \vec{p}} \sum_{\vec{q}, \alpha_1, \alpha_2} \theta(k_F - |\vec{k}|) \theta(k_F - |\vec{p}|) \\ &= \frac{1}{2\Omega} (g^2) \left(\frac{\Omega}{(2\pi)^3} \int d^3p \theta(k_F - |\vec{p}|) \right) \left(\frac{\Omega}{(2\pi)^3} \int d^3k \theta(k_F - |\vec{k}|) \right) \\ &= \frac{1}{2\Omega} N^2 \text{ using the previous result for } N. \end{aligned}$$

The energy per particle (which is well defined in the thermodynamic limit)

β $E_{\text{direct}}^{(1)} / N = \frac{1}{2\Omega} N = \frac{1}{2} \rho$ and the energy density \mathcal{E} (energy/volume)

β $E_{\text{direct}}^{(1)} = \frac{1}{2\Omega} N^2 = \frac{1}{2} \rho^2$

18/03

The other term is similar but messier:

exchange: $\langle F | a^\dagger a^\dagger a a | F \rangle \rightarrow \delta_{\vec{k}+\vec{q}, \vec{p}} \delta_{\alpha_1 \alpha_2} \langle F | a_{\vec{k}+\vec{q}, \alpha_1}^\dagger a_{\vec{k}, \alpha_2}^\dagger a_{\vec{k}, \alpha_1} a_{\vec{k}+\vec{q}, \alpha_2} | F \rangle$

only one interchange \Rightarrow $= -\delta_{\vec{k}+\vec{q}, \vec{p}} \delta_{\alpha_1 \alpha_2} \langle F | \hat{n}_{\vec{k}+\vec{q}, \alpha_1} \hat{n}_{\vec{k}, \alpha_1} | F \rangle$

$= -\delta_{\vec{k}+\vec{q}, \vec{p}} \delta_{\alpha_1 \alpha_2} \theta(k_f - |\vec{k}+\vec{q}|) \theta(k_f - k)$

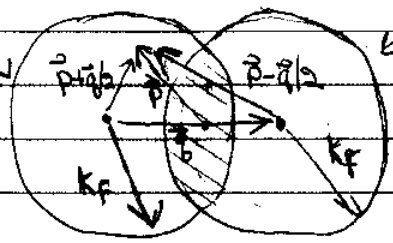
$\Rightarrow E_{\text{exchange}}^{(1)} = \sum_{\vec{k}, \alpha_1, \alpha_2} \sum_{\vec{k}, \vec{q}} (-\delta_{\vec{k}+\vec{q}, \vec{p}}) \delta_{\alpha_1 \alpha_2} \theta(k_f - |\vec{k}+\vec{q}|) \theta(k_f - k)$

$= -\frac{1}{2} g \frac{\Omega}{(2\pi)^6} \int d^3k d^3q \theta(k_f - |\vec{k}+\vec{q}|) \theta(k_f - k)$

It is convenient (and useful!) to make the integration more symmetric by defining $\vec{p} = \vec{k} + \frac{1}{2}\vec{q}$ and integrating over \vec{p} and \vec{q} .

$\Rightarrow E_{\text{exchange}}^{(1)} = -\frac{g\lambda}{2} \frac{\Omega}{(2\pi)^6} \int d^3q \int d^3p \theta(k_f - |\vec{p} + \frac{1}{2}\vec{q}|) \theta(k_f - |\vec{p} - \frac{1}{2}\vec{q}|)$

We can represent the integration region diagrammatically as $|\vec{p} + \frac{1}{2}\vec{q}| = k_f$ and $|\vec{p} - \frac{1}{2}\vec{q}| = k_f$ (2-d slice)



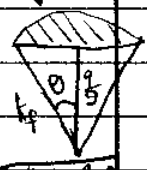
The diagram is for fixed \vec{q} , which is the separation of the two spheres.

- \vec{p} is measured from the center and is integrated over.
- The two sphere surfaces are the boundaries of the θ -functions.
- The shaded region is the "volume" (in p space) we want. $\Rightarrow q < 2k_f$

$\Rightarrow \int d^3p \theta(k_f - |\vec{p} + \frac{1}{2}\vec{q}|) \theta(k_f - |\vec{p} - \frac{1}{2}\vec{q}|) = 2\theta(2k_f - q) \int_{x=0}^1 dx \int_{r=0}^{k_f/x} 4\pi r^2 dr$

$= \frac{4\pi}{3} k_f^3 \theta(1-x) \int_0^1 dx (1 - \frac{x^2}{2}) = \frac{4\pi}{3} k_f^3 \theta(1-x) [1 - \frac{3}{2}x + \frac{1}{2}x^3]$

we want 2x this volume



$\cos \theta_{\text{max}} = \frac{x}{k_f} \Rightarrow x$

$\Rightarrow E_{\text{exchange}}^{(1)} = -\frac{g\lambda}{2} \frac{\Omega}{(2\pi)^6} \frac{4\pi k_f^3}{3} (2k_f)^3 \int_0^1 dx x^2 (1 - \frac{3}{2}x + \frac{1}{2}x^3) \cdot 4\pi = -\frac{1}{2} \frac{1}{g} \left(\frac{gk_f^3}{6\pi^2}\right) \left(\frac{gk_f^3}{6\pi^2}\right) \Omega = -\frac{1}{9} \frac{1}{g} \Omega$

1/8/03

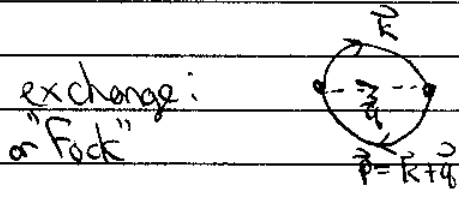
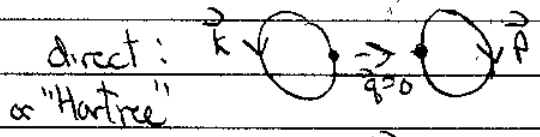
Putting the direct and exchange together:

$$\frac{E^{(4)}}{N} = \frac{\lambda}{2} \rho \left(1 - \frac{1}{g}\right) = \frac{4\pi\hbar^2 a_s}{2m} \frac{g k_F^3}{6\pi^2} \left(1 - \frac{1}{g}\right) = \frac{\hbar^2 k_F^2}{2m} \left(g - 1\right) \frac{2k_F a_s}{3\pi}$$

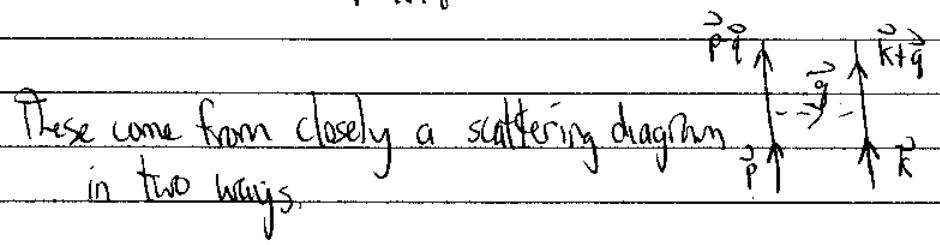
The last form provides a direct comparison with $E^{(0)}$:

$$\frac{E}{N} = \frac{E^{(0)}}{N} + \frac{E^{(4)}}{N} + \dots = \frac{\hbar^2 k_F^2}{2m} \left[\frac{3}{5} + (g-1) \frac{2k_F a_s}{3\pi} + \dots \right]$$

- Why does the correction vanish for $g=1$? What if we trap only one spin state?
- So for $k_F a_s \ll 1$, we have a meaningful expansion \Rightarrow dilute Fermi gas.
- It's actually trickier than this if $a_s < 0$ (attractive), since there is a lower energy ground state than this, (more later!)
- This is well defined for a repulsive $a_s > 0$ (and small).
- How do we find the higher-order terms?
- How do we find the pressure? $\Rightarrow P = -\left(\frac{\partial E}{\partial V}\right)_N = g^2 \frac{\partial (E/N)}{\partial g}$
- How do you expect the correction term $E^{(4)}$ to go
- Quick preview of Feynman diagrams



- --- $\Rightarrow \lambda$
- conserve momentum at \bullet 's
- spin sums by following solid (fermion) lines $\Rightarrow g^2$ vs. g



1/8/03

Often we want to switch to the \vec{x} basis. We can do this by forming the "field operators" $\hat{\psi}(\vec{x})$ and $\hat{\psi}^\dagger(\vec{x})$:

$$\hat{\psi}(\vec{x}) \equiv \sum_{\vec{k}} \hat{\psi}_{\vec{k}}(\vec{x}) c_{\vec{k}} \quad c_{\vec{k}} = a_{\vec{k}} \text{ or } b_{\vec{k}}$$

$$\hat{\psi}^\dagger(\vec{x}) \equiv \sum_{\vec{k}} \hat{\psi}_{\vec{k}}^\dagger(\vec{x}) c_{\vec{k}}^\dagger$$

where $\vec{k} = \{k, s_z\}$ or $\{E, L, J, m\}$ or ... for spin 1/2 fermions

In flat case,
$$\hat{\psi}_{\vec{k}}(\vec{x}) = \begin{bmatrix} \hat{\psi}_{\vec{k}}(\vec{x})_1 \\ \hat{\psi}_{\vec{k}}(\vec{x})_2 \end{bmatrix} \equiv \hat{\psi}_{\vec{k}}(\vec{x})_\alpha$$

Using $c_{\vec{k}}, c_{\vec{k}}^\dagger$ relations,

$$\begin{aligned} [\hat{\psi}_\alpha(\vec{x}), \hat{\psi}_\beta^\dagger(\vec{x}')]_{\pm} &= \sum_{\vec{k}} \hat{\psi}_{\vec{k}}(\vec{x})_\alpha \hat{\psi}_{\vec{k}}^\dagger(\vec{x}')_\beta = \delta_{\alpha\beta} \delta(\vec{x} - \vec{x}') \\ [\hat{\psi}_\alpha(\vec{x}), \hat{\psi}_\beta(\vec{x}')]_{\pm} &= [\hat{\psi}_\alpha^\dagger(\vec{x}), \hat{\psi}_\beta^\dagger(\vec{x}')]_{\pm} = 0 \end{aligned}$$

Operators

$$\hat{H} = \int d^3x \hat{\psi}^\dagger(\vec{x}) \hat{T}(\vec{x}) \hat{\psi}(\vec{x}) + \frac{1}{2} \iint d^3x d^3x' \hat{\psi}^\dagger(\vec{x}) \hat{\psi}^\dagger(\vec{x}') V(\vec{x}, \vec{x}') \hat{\psi}(\vec{x}) \hat{\psi}(\vec{x}')$$

⇒ "second quantization": looks like expectation values wrt operators!
• The integration over \vec{x}, \vec{x}' generate the matrix elements

If $J = \sum_{i=1}^N J(\vec{x}_i)$ is a first-quantized operator ⇒
$$\hat{J} = \sum \langle r | J | s \rangle c_r^\dagger c_s = \int d^3x \sum_{r,s} \hat{\psi}_r^\dagger(\vec{x}) J(\vec{x}) \hat{\psi}_s(\vec{x}) c_r^\dagger c_s = \int d^3x \hat{\psi}^\dagger(\vec{x}) J(\vec{x}) \hat{\psi}(\vec{x})$$

number density $\rho(\vec{x}) = \sum_{i=1}^N \delta(\vec{x} - \vec{x}_i) \Rightarrow \hat{n}(\vec{x}) = \sum_{\alpha} \hat{\psi}_\alpha^\dagger(\vec{x}) \hat{\psi}_\alpha(\vec{x}) c_r^\dagger c_s = \hat{\psi}^\dagger(\vec{x}) \hat{\psi}(\vec{x})$

and total number $\hat{N} = \int d^3x \hat{n}(\vec{x}) = \sum_r c_r^\dagger c_r = \sum_r \hat{n}_r = \int d^3x \hat{\psi}^\dagger(\vec{x}) \hat{\psi}(\vec{x})$