

11/3/14

Monday Q&A Class

• Finish from 10/29/14 notes

- ⑦ correspondence of model and full field theory (see Appendix A)
- ⑧ Green's function
- ⑨ finite range potential

• Nonperturbative expansion

- Recap ⑥ from 10/29/14
- ⑩-⑪ (11/3/14)
- brief: auxiliary fields ⑤

• Euclidean vs. Minkowski space ⑥, ⑦

• Self-consistent Hartree-Fock at $T=0$ with real time propagators.

- ⑧-⑪
- Details via Piazza

11/3/14

Nonperturbative approximations

- We have diagrammatic expansions to calculate the ground state energy or finite temperature thermodynamic functions in perturbation theory
- What if pert. theory is inadequate?

Here: Nonperturbative "conserving approximation" that sum infinite classes of diagrams. [Also called "phi-derivable"]

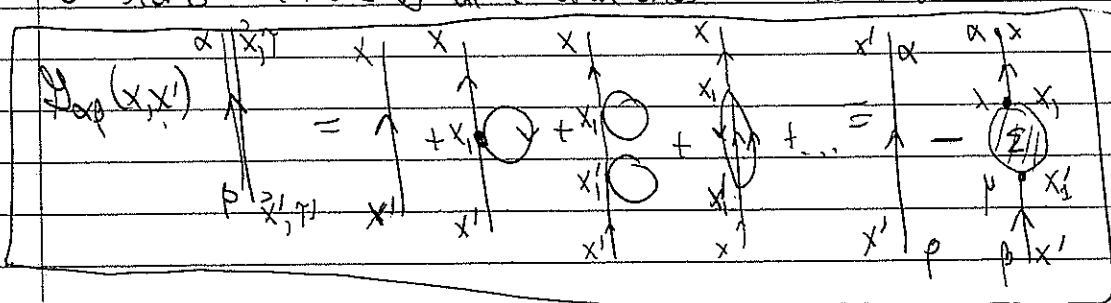
First, go back to one body Green's function:

$$G_{\text{op}}(\vec{x}, \tau; \vec{x}', \tau') = \frac{\text{Tr}[\hat{T} \psi(\vec{x}, \tau) \psi^\dagger(\vec{x}', \tau') e^{-S\psi}]}{\text{Tr}[\hat{T} e^{-S\psi}}$$

$$= \text{Tr} [e^{-\beta(\hat{H} - \mu\hat{N})} \hat{T} [\psi(\vec{x}, \tau) \psi^\dagger(\vec{x}', \tau')]] / \text{Tr} [e^{-\beta(\hat{H} - \mu\hat{N})}]$$

see Appendix A for more details on field operators.

In the diagrammatic expansion, each term is either $G_{\text{op}}^0(x, x')$ with $x = (\vec{x}, \tau)$ or starts with one G^0 at x' and ends with another at x . So



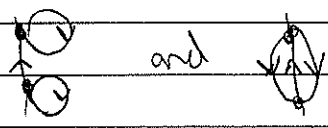
is the general structure, where the "improper self-energy" Σ stands for all possible (connected) diagram insertions.

In equation form, this is an integral equation (defining the self energy)

$$G_{\text{op}}(x, x') = G_{\text{op}}^0(x, x') - \int \int d^2x_1 \int d\tau_1 \int d^2x_2 \int d\tau_2 G_{\text{op}}^0(x, x_1) \Sigma(x_1, x_2) G_{\text{op}}(x_2, x')$$

11/3/14

But we can go further. Compare two 2nd order contributions to Σ :

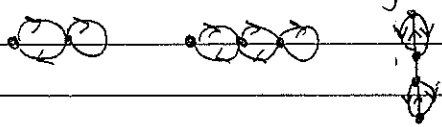





\Rightarrow each part of the left diagram looks like a self-energy piece, because there is a single line joining them.

\Rightarrow call it 1PR "one-particle reducible"

The right diagram is 1PI "one-particle irreducible" because it does not fall into two pieces when a single line is cut (unlike the left diagram).

Check: are the following 1PR or 1PI?



The diagrams in Σ that are 1PI are called the "proper" self-energy and designated Σ^* =  +  +  + ...

Diagrammatically, Σ and Σ^* are related by

$$\Sigma = \Sigma^* - \begin{matrix} \text{circle} \\ \uparrow \\ \text{circle} \\ \uparrow \\ \text{circle} \end{matrix} + \begin{matrix} \text{circle} \\ \uparrow \\ \text{circle} \\ \uparrow \\ \text{circle} \end{matrix} + \dots$$

or, in equations (suppressing spin indices)

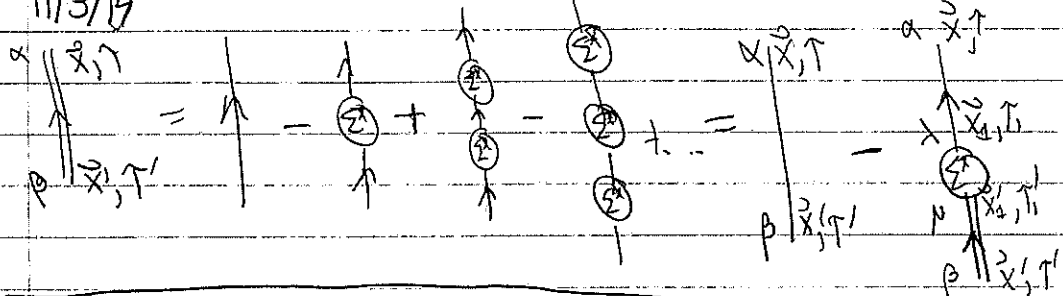
$$\Sigma(x_1, x_1') = \Sigma^*(x_1, x_1') - \int d^4x_2 d^4x_2' \Sigma^*(x_1, x_2) G^0(x_2, x_2') \Sigma^*(x_2', x_1') + \dots$$

where $\int d^4x_2$ means $\int d^3x_2 dt_2$, etc.

Now we can insert this equation back into our original equation for G to derive another integral equation:

11/3/14

(4)



$$\Rightarrow \underline{D}_{\text{op}}(x, x') = \underline{D}_{\text{op}}^0(x, x') - \int_{\mathcal{T}} \int_{\mathcal{X}} \int_{\mathcal{T}'} \int_{\mathcal{X}'} \underline{D}_{\text{op}}^0(x, x_1) \underline{\Sigma}(x_1, x_1') \underline{D}_{\text{op}}(x_1', x')$$

- This is "Dyson's equation" for the propagator ("2-point function")
- Note that it is like the previous equation for \underline{D} except in the integral $\underline{\Sigma} \Rightarrow \underline{\Sigma}^*$ and $\underline{D}^0 \Rightarrow \underline{D}$.

If we approximate $\underline{\Sigma}^*$ to some order in perturbation theory, we get an all orders approximation to \underline{D} !

Think of the Dyson's equation as a matrix equation (in spin and space time indices)

$$\underline{D} = \underline{D}^0 + \underline{D}^0 \underline{\Sigma}^* \underline{D} \Rightarrow \underline{D} - \underline{D}^0 \underline{\Sigma}^* \underline{D} = \underline{D}^0$$

$$\Rightarrow (\underline{1} - \underline{D}^0 \underline{\Sigma}^*) \underline{D} = \underline{D}^0 \Rightarrow \underline{D} = (\underline{1} - \underline{D}^0 \underline{\Sigma}^*)^{-1} \underline{D}^0 \Rightarrow \underline{D} = \underline{D}^0 (\underline{1} - \underline{D}^0 \underline{\Sigma}^*)^{-1} = \underline{D}^0 \underline{\Sigma}^*$$

Recall that $\underline{D}^0 = \frac{d}{dt} + \frac{v^2}{2m}$, so $\underline{\Sigma}^*$ is like an external potential.

- How do we get the energy? We might think $\int \underline{D} \underline{\Sigma} \Rightarrow \textcircled{\underline{\Sigma}^*}$ because, if expanded, it has all the diagrams for $\ln Z$.
- But the factors are incorrect!

\Rightarrow we'll come back and see how to do it correctly!

11/3/14

Partition Function for $V(\vec{x}-\vec{x}')$:

$$Z = \text{Tr} e^{-\beta(\hat{H}-\mu\hat{N})}$$

$$= \int \mathcal{D}[\psi(x,\tau)] \mathcal{D}[\bar{\psi}(x,\tau)] e^{-\int_0^\beta d\tau \int d^3x \bar{\psi}(x,\tau) \left(\frac{d}{d\tau} - \frac{\nabla^2}{2m} - \mu \right) \psi(x,\tau)}$$

$$\psi(\vec{x},\beta) = -\psi(\vec{x},0) \quad \times e^{-\frac{1}{2} \int_0^\beta d\tau \int d^3x \bar{\psi}(\vec{x},\tau) \left[\frac{d}{d\tau} - \mu - \frac{\nabla^2}{2m} + \frac{1}{2} \int d^3y V(\vec{x}-\vec{y}) \right] \psi(\vec{x},\tau)}$$

$V = \lambda \delta(\vec{x}-\vec{x}')$

$$\int \mathcal{D}[\psi(x,\tau)] \mathcal{D}[\bar{\psi}(x,\tau)] e^{-\int_0^\beta d\tau \int d^3x \bar{\psi}(x,\tau) \left(\frac{d}{d\tau} - \frac{\nabla^2}{2m} - \mu \right) \psi(x,\tau)}$$

$$\psi(\vec{x},\beta) = -\psi(\vec{x},0) \quad \times e^{-\frac{\lambda}{2} \int_0^\beta d\tau \int d^3x \bar{\psi}(\vec{x},\tau) \psi(\vec{x},\tau) \psi(\vec{x},\tau) \psi(\vec{x},\tau)}$$

As we've seen, this is suitable for perturbation theory, but what about stochastic simulation?

The fermion form with Grassman fields is not suitable
 \Rightarrow introduce an auxiliary field

analogy $1 = \frac{\int_{-\infty}^{\infty} d\sigma e^{-(\sigma - \sqrt{S^2})^2}}{\int_{-\infty}^{\infty} d\sigma e^{-\sigma^2}} \Rightarrow \int_{-\infty}^{\infty} d\sigma e^{-\sigma^2 + \sigma^4} = N \int_{-\infty}^{\infty} d\sigma \int_{-\infty}^{\infty} d\tau e^{-\sigma^2 + \sigma^4} e^{-\sigma^2 + 2\sigma \tau S^2} e^{-\tau^2 S^4}$

\Rightarrow we can do the Gaussian τ integral leaving one over σ .

In path integral Z with full $V(\vec{x}-\vec{x}')$:

$$Z = \int \mathcal{D}[\psi] \mathcal{D}[\bar{\psi}] \mathcal{D}[\sigma(x)] e^{-\int d\tau d\tau' \int d^3x d^3x' \bar{\psi}(x,\tau) \left[\frac{d}{d\tau} - \mu - \frac{\nabla^2}{2m} + \frac{1}{2} \int d^3y V(\vec{x}-\vec{y}) \sigma(y,\tau) \right] \psi(x',\tau)}$$

$$\int \mathcal{D}[\sigma(x)] [\det M(\sigma)]^2 e^{-\frac{1}{2} \int d\tau \int d^3x d^3x' \sigma(\vec{x},\tau) V(\vec{x}-\vec{x}') \sigma(\vec{x}',\tau)}$$

where $M_{\vec{x}\tau; \vec{x}'\tau'} = \left[\frac{d}{d\tau} - \mu - \frac{\nabla^2}{2m} + \frac{1}{2} \int d^3y V(\vec{x}-\vec{y}) \sigma(\vec{y},\tau) \right]_{\vec{x}\tau; \vec{x}'\tau'}$

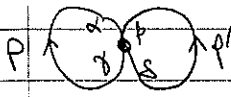
11/3/14

(6)

Previously we've evaluated the "bubble" diagram using Feynman rules with real time (Minkowski) propagators.

- Those ones have θ functions for momenta compared to the Fermi momentum with poles just below or above the real p_0 axis.
- The imaginary time ("Euclidean") propagators have poles for imaginary p_0 and include the chemical potential μ_0 .

• Bubble diagram in momentum space for energy density



• Same rules for spin and symmetry factor parts and vertex \Rightarrow yield $\frac{1}{2}(\delta_{\alpha\beta}\delta_{\gamma\delta} + \delta_{\alpha\delta}\delta_{\beta\gamma})\delta_{\alpha\beta\gamma\delta}$

• The lines become momentum space Green's functions: (F_0)

$$G_{\alpha\beta}^0(p) = \frac{-\delta_{\alpha\beta}}{ip_0 - (\epsilon_p^0 - \mu_0)} \quad \text{with} \quad \epsilon_p^0 = \frac{\hbar^2 p^2}{2m} \quad (\text{but not relativistic!})$$

instead of $G_{\alpha\beta}^0(x, x')$ from before. Each line gets a four-momentum, which is conserved (no constraint here because p and p' both go in and out of the vertex).

• Integrate $\int \frac{d^4p}{(2\pi)^4} e^{ip_0\eta}$ over energy four-momentum. $e^{ip_0\eta}$ is a convergence factor that tells us how to close contours. $\eta \rightarrow 0^+$ at end.

• At $T \neq 0$, $p_0 \Rightarrow \omega_n$ a discrete frequency and we have "Matsubara sums".

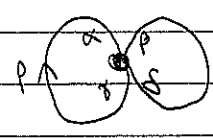
$$\begin{aligned} \Rightarrow \text{Bubble} &= \frac{1}{2}(\delta_{\alpha\beta}\delta_{\gamma\delta} + \delta_{\alpha\delta}\delta_{\beta\gamma})\delta_{\alpha\beta\gamma\delta} \int \frac{d^4p}{(2\pi)^4} \frac{-e^{ip_0\eta}}{ip_0 - (\epsilon_p^0 - \mu_0)} \int \frac{d^4p'}{(2\pi)^4} \frac{-e^{ip'_0\eta}}{ip'_0 - (\epsilon_{p'}^0 - \mu_0)} \\ &= \frac{1}{2}V(2\pi)^3 \left[-\int \frac{d^3p}{(2\pi)^3} \theta(\mu_0 - \epsilon_p^0) \right] \left[-\int \frac{d^3p'}{(2\pi)^3} \theta(\mu_0 - \epsilon_{p'}^0) \right] \\ &= \frac{1}{2} \left(1 - \frac{n}{V}\right)^2 \bar{\rho}^2 \quad \text{which is what we get before for the energy density} \end{aligned}$$

• note that we close in the upper half plane so $e^{(k_0 + i\text{Im} p_0)\eta} \propto e^{-\text{Im} p_0 \eta}$ makes the integral converge. We pick up a pole if $\mu_0 > \epsilon_p^0$, otherwise 0.

Euclidean \leftrightarrow Minkowski
 $t \leftrightarrow -i\tau$ so $\mathcal{L}_E(x, \tau) = -\mathcal{L}(x, -i\tau) \quad Z = \int e^{iS}$
 or $\tau \leftrightarrow it$ or $\mathcal{L}(x, t) = -\mathcal{L}_E(x, it) \quad Z = \int e^{-S_E}$

11/3/14

Check Minkowski vs. Euclidean versions.



← contribution to energy density $\frac{E}{V}$ at $T \rightarrow 0$

symmetry factor has

Minkowski

same spin factors

Euclidean

$$\frac{1}{2} (\delta_{\alpha\beta} \delta_{\gamma\delta} + \delta_{\alpha\gamma} \delta_{\beta\delta}) \delta_{\alpha\gamma} \delta_{\beta\delta}, \quad \frac{1}{2} (\delta_{\alpha\beta} \delta_{\gamma\delta} + \delta_{\alpha\gamma} \delta_{\beta\delta}) \delta_{\alpha\gamma} \delta_{\beta\delta}$$

$$\times (-i C_0)$$

$$\times (C_0)$$

(vertex factor of i or $-i$) e^{iS} vs. e^{-SE}

$$\times \int \frac{d^4 p}{(2\pi)^4} i \left[\frac{\theta(p^0 - k_f)}{p^0 - \epsilon_f + i\epsilon} + \frac{\theta(k_f - p^0)}{p^0 - \epsilon_f - i\epsilon} \right] e^{i p_0 \eta}, \quad \times \int \frac{d^4 p}{(2\pi)^4} \left[\frac{-1}{i p^0 - (\epsilon_f - \mu_0)} \right] e^{i p_0 \eta}$$

$$\times \int \frac{d^4 p}{(2\pi)^4} i \left[\frac{\theta(p^0 - k_f)}{p^0 - \epsilon_f + i\epsilon} + \frac{\theta(k_f - p^0)}{p^0 - \epsilon_f - i\epsilon} \right] e^{i p_0 \eta}, \quad \times \int \frac{d^4 p}{(2\pi)^4} \left[\frac{-1}{i p^0 - (\epsilon_f - \mu_0)} \right] e^{i p_0 \eta}$$

$$= \frac{C_0}{2} \nu(\nu-1) \left[\int \frac{d^4 p}{(2\pi)^4} \theta(k_f - p^0) \right] \left[\int \frac{d^4 p}{(2\pi)^4} \theta(\mu_0 - \epsilon_f) \right]$$

these are the same; use either!

$$= \frac{C_0}{2} \left(1 - \frac{4}{\nu}\right) g^2 \quad \text{where } g = \frac{\nu k_f^3}{6\pi^3}, \quad \mu_0 = \frac{k_f^2}{2m}$$

Note that we are just taking $p_0 \rightarrow i p_0$, as implied by the Wick rotation.

11/3/14

Goal: Recover HF result from Feynman rules resummed

$$\hat{H} = \sum_{i=1}^A \hat{T}(\vec{x}_i) + \frac{1}{2} \sum_{i,j=1}^A \hat{V}(\vec{x}_i, \vec{x}_j)$$

$$\left[-\frac{\nabla^2}{2m} + \Gamma_H(\vec{x}) \right] \phi_i(\vec{x}) + \int d^3y \Gamma_F(\vec{x}, \vec{y}) \phi_i(\vec{y}) = \epsilon_i \phi_i(\vec{x})$$

$$\Gamma_H(\vec{x}) = \int d^3y V(\vec{x}, \vec{y}) \sum_{j=1}^A |\phi_j(\vec{y})|^2 = \int d^3y V(\vec{x}, \vec{y}) \rho(\vec{y}, \vec{y})$$

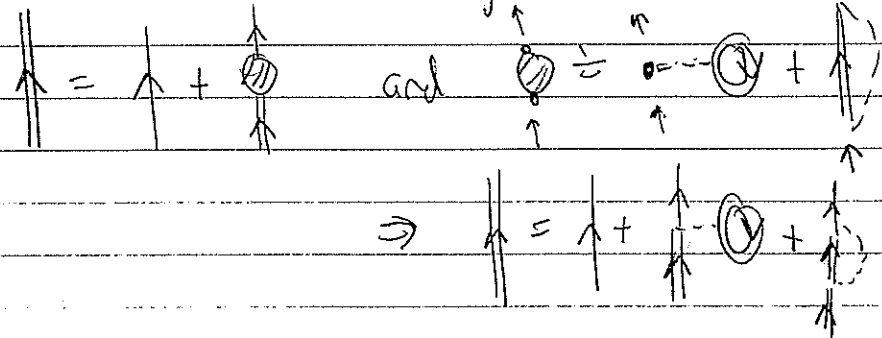
$$\Gamma_F(\vec{x}, \vec{y}) = -V(\vec{x}, \vec{y}) \sum_{j=1}^A \phi_j^+(\vec{y}) \phi_j(\vec{x}) = -V(\vec{x}, \vec{y}) \rho(\vec{y}, \vec{x})$$

$$E_{HF} = \sum_{i=1}^A \frac{1}{2m} \int d^3x \nabla \phi_i^+ \nabla \phi_i$$

$$+ \frac{1}{2} \int d^3x \int d^3y \sum_i |\phi_i(\vec{x})|^2 V(\vec{x}, \vec{y}) \sum_j |\phi_j(\vec{y})|^2 \quad \left(\begin{matrix} x & y \\ \uparrow & \downarrow \end{matrix} \right)$$

$$- \frac{1}{2} \int d^3x \int d^3y \sum_{ij} \phi_i^+(\vec{x}) \phi_i(\vec{x}) V(\vec{x}, \vec{y}) \phi_j^+(\vec{y}) \phi_j(\vec{x}) \quad \left(\begin{matrix} \dots & y \\ \dots & \dots \end{matrix} \right)$$

Self-consistent Hartree-Fock in diagrams



• We'll follow Fetter and Walecka section 10, which uses a real-time $T=0$ formulation. A finite temperature, Euclidean version of the same discussion is in F&W section 27.

(9)

11/3/14

Let's take $\hat{H} = \hat{H}_0 + \hat{H}_2$ where \hat{H}_0 includes a single-particle potential $U(\vec{x})$, which then defines a single-particle complete set of states: (here in first quantization)

$$H_0 \psi_j^0(\vec{x}) = \left[-\frac{\nabla^2}{2m} + U(\vec{x}) \right] \psi_j^0(\vec{x}) = \epsilon_j^0 \psi_j^0(\vec{x})$$

• Even if $\hat{H} = \hat{T} + \hat{V}$, we can always write $\hat{H} = (\hat{T} + \hat{U}) + (\hat{V} - \hat{U}) \equiv \hat{H}_0 + \hat{H}_2$.

In 2nd quantized form

$$\hat{H}_0 = \int d^3x \hat{\psi}_\alpha^\dagger(\vec{x}) \left[-\frac{\nabla^2}{2m} + U(\vec{x}) \right] \hat{\psi}_\alpha(\vec{x})$$

$$\hat{H}_2 = \frac{1}{2} \int d^3x d^3x' \hat{\psi}_\alpha^\dagger(\vec{x}) \hat{\psi}_\beta^\dagger(\vec{x}') V(\vec{x} - \vec{x}') \hat{\psi}_\beta(\vec{x}') \hat{\psi}_\alpha(\vec{x})$$

where we can choose the ψ_j^0 to be the basis for the field operators $\hat{\psi}_\alpha(\vec{x})$:

$$\hat{\psi}_\alpha(\vec{x}) = \sum_j \psi_j^0(\vec{x}) a_j; \quad \hat{\psi}_\alpha^\dagger(\vec{x}) = \sum_j \psi_j^{0*}(\vec{x}) a_j^\dagger$$

• Dyson's equation for Hartree-Fock is then (suppressing spin indices)

$$G(x, y) = G^0(x, y) + \int d^4x_1 d^4x_1' G^0(x, x_1) \Sigma^*(x_1, x_1') G(x_1', y)$$

where the proper self-energy is

$$\Sigma^*(x_1, x_1') = -i \int dt_2 dt_2' \left[\int d^3x_2 \int d^3x_2' G(\vec{x}_2, t_2; \vec{x}_2, t_2') V(\vec{x}_1 - \vec{x}_2) - V(\vec{x}_1 - \vec{x}_1') G(\vec{x}_1, t_1; \vec{x}_1, t_1') \right]$$

• \hat{H}_0 and \hat{H}_2 are both time independent, so switching to a Fourier representation in frequency makes sense.

$$G(\vec{x}t, \vec{x}'t') = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} e^{i\omega(t-t')} G(\vec{x}, \vec{x}'; \omega) \quad (\text{similarly with } G^0)$$

$$\Sigma^*(\vec{x}t, \vec{x}'t') = \Sigma^*(\vec{x}, \vec{x}') \delta(t-t') = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} e^{i\omega(t-t')} \Sigma^*(\vec{x}, \vec{x}')$$

11/3/14

Let's first consider G^0 in terms of field operators:

$$iG_{\text{op}}^0(\vec{x}t, \vec{x}'t') = \frac{\langle \Phi_0 | T [\hat{\psi}_{H_0}(\vec{x}t) \hat{\psi}_{H_0}^\dagger(\vec{x}'t')] | \Phi_0 \rangle}{\langle \Phi_0 | \Phi_0 \rangle}$$

↑ Ground state of H_0

- The time-ordered product means $\hat{\psi}\hat{\psi}^\dagger$ if $t > t'$ and $-\hat{\psi}^\dagger\hat{\psi}$ if $t < t'$.
- The Heisenberg field operators for \hat{H}_0 (in interaction picture) are

$$\hat{\psi}_{H_0}(\vec{x}t) = e^{i\hat{H}_0 t} \hat{\psi}(\vec{x}) e^{-i\hat{H}_0 t}$$

$$\Rightarrow iG_{\text{op}}^0(\vec{x}t, \vec{x}'t') = \sum_j \psi_j^0(\vec{x}) \psi_j^0(\vec{x}') e^{-i\epsilon_j^0(t-t')} \times \left[\theta(t-t') \langle \Phi_0 | a_j^\dagger a_j | \Phi_0 \rangle - \theta(t-t') \langle \Phi_0 | a_j a_j^\dagger | \Phi_0 \rangle \right]$$

$$= \sum_j \psi_j^0(\vec{x}) \psi_j^0(\vec{x}') e^{-i\epsilon_j^0(t-t')} \times \left[\theta(t-t') \delta(\epsilon_j^0 - \epsilon_j^0) - \theta(t-t') \delta(\epsilon_j^0 - \epsilon_j^0) \right]$$

where ϵ_j^0 is the energy of the last filled state

• Check that $[i\frac{\partial}{\partial t} + \frac{\nabla^2}{2m} - U(\vec{x})] G_{\text{op}}^0(\vec{x}t, \vec{x}'t') = \delta^3(\vec{x}-\vec{x}') \delta(t-t') \delta_{\text{op}}$

• If we calculate the Fourier transform: (Piazza!)

$$G^0(\vec{x}, \vec{x}', \omega) = \int_{-\infty}^{\infty} dt (t-t') e^{-i\omega(t-t')} G^0(\vec{x}t, \vec{x}'t')$$

$$= \sum_j \psi_j^0(\vec{x}) \psi_j^0(\vec{x}')^* \left[\frac{\theta(\epsilon_j^0 - \epsilon_F^0)}{\omega - \epsilon_j^0 + i\eta} + \frac{\theta(\epsilon_F^0 - \epsilon_j^0)}{\omega - \epsilon_j^0 - i\eta} \right]$$

- If we took $U(\vec{x}) \equiv 0$, then the $\psi_j^0(\vec{x}) \rightarrow \frac{1}{\sqrt{V}} e^{i\vec{k}_j \cdot \vec{x}}$ and $\epsilon_j^0 \rightarrow \frac{\hbar^2 k_j^2}{2m}$, recovering the result for G^0 from the non-interacting case.
- The analytic structure is simple poles at the non-interacting energies with residues given by the single-particle wave functions.
- This is an example of a spectral representation.

• Defining the differential operator $L_1 = \omega - H_0 = \omega + \frac{\nabla^2}{2m} - U(\vec{x}_1)$

as expected \Rightarrow from a Green's function

$$L_1 G^0(\vec{x}_1, \vec{x}_2, \omega) = \sum_j (\omega - \epsilon_j^0) \psi_j^0(\vec{x}_1) \psi_j^0(\vec{x}_2)^* \left[\frac{\theta(\epsilon_j^0 - \epsilon_F^0)}{\omega - \epsilon_j^0 + i\eta} + \frac{\theta(\epsilon_F^0 - \epsilon_j^0)}{\omega - \epsilon_j^0 - i\eta} \right]$$

$$\Rightarrow \sum_j \psi_j^0(\vec{x}_1) \psi_j^0(\vec{x}_2)^* = \delta^3(\vec{x}_1 - \vec{x}_2) \text{ by completeness,}$$

(1)

11/3/14

For Hartree-Fock, the full propagator G will also depend only on $t-t'$ and with Σ^* having a $\delta(t-t')$ dependence, the frequency Fourier transformation of the Dyson's equation is

$$G(\vec{x}_1, \vec{y}, \omega) = G^0(\vec{x}_1, \vec{y}, \omega) + \int d^3x_2 \int d^3x_2' G^0(\vec{x}_1, \vec{x}_2, \omega) \Sigma^*(\vec{x}_2, \vec{x}_2') G(\vec{x}_2', \vec{y}, \omega)$$

with

$$\Sigma^*(\vec{x}_2, \vec{x}_2') = -i\nu \delta(\vec{x}_2, \vec{x}_2') \int d^3x_3 V(\vec{x}_2, \vec{x}_3) \int \frac{d\omega'}{2\pi} e^{i\omega'\tau} G(\vec{x}_3, \vec{x}_3, \omega')$$

$$+ iV(\vec{x}_2, \vec{x}_2') \int \frac{d\omega'}{2\pi} e^{i\omega'\tau} G(\vec{x}_2, \vec{x}_2', \omega')$$

To solve this, we realize that because Σ^* (for HF) is independent of frequency, G will have the same structure as G^0 :

$$G(\vec{x}, \vec{x}', \omega) = \sum_j \psi_j(\vec{x}) \psi_j^*(\vec{x}') \left[\frac{G(\epsilon_j - \epsilon)}{\omega - \epsilon_j + i\eta} + \frac{\theta(\epsilon_f - \epsilon_j)}{\omega - \epsilon_j - i\eta} \right]$$

where the single-particle wave functions $\{\psi_j(\vec{x})\}$ have energy ϵ_j .
How do we determine them?

First we note that evaluating Σ^* looks like the HF potential from before:

$$\Sigma^*(\vec{x}_2, \vec{x}_2') = \nu \int d^3x_3 \delta(\vec{x}_2, \vec{x}_2') \int d^3x_3 V(\vec{x}_2, \vec{x}_3) \sum_j |\psi_j(\vec{x}_3)|^2 \theta(\epsilon_f - \epsilon_j)$$

$$= \int d^3x_3 \delta(\vec{x}_2, \vec{x}_2') \int d^3x_3 V(\vec{x}_2, \vec{x}_3) \rho(\vec{x}_3)$$

$$- V(\vec{x}_2, \vec{x}_2') \sum_j \psi_j(\vec{x}_2) \psi_j^*(\vec{x}_2') \theta(\epsilon_f - \epsilon_j)$$

Now apply $L_2 = G_0^{-1}$ to the Dyson equation for G :

$$L_2 G(\vec{x}_1, \vec{x}_1', \omega) = \delta(\vec{x}_1, \vec{x}_1') + \int d^3x_2 \Sigma^*(\vec{x}_1, \vec{x}_2) G(\vec{x}_2, \vec{x}_1', \omega)$$

We'll leave it as a Piazza exercise to show that this implies the HF equation:

$$\left[\frac{-\nabla^2}{2m} + U(\vec{x}_1) \right] \psi_j(\vec{x}_1) + \int d^3x_2 \Sigma^*(\vec{x}_1, \vec{x}_2) \psi_j(\vec{x}_2) = \epsilon_j \psi_j(\vec{x}_1)$$