

# Chapter 6

## Electronic Structure of Atoms

The number & arrangement of  $e^-$  in an atom is responsible for its chemical behavior

### I) The Wave Nature of Light

#### A) Electromagnetic Radiation

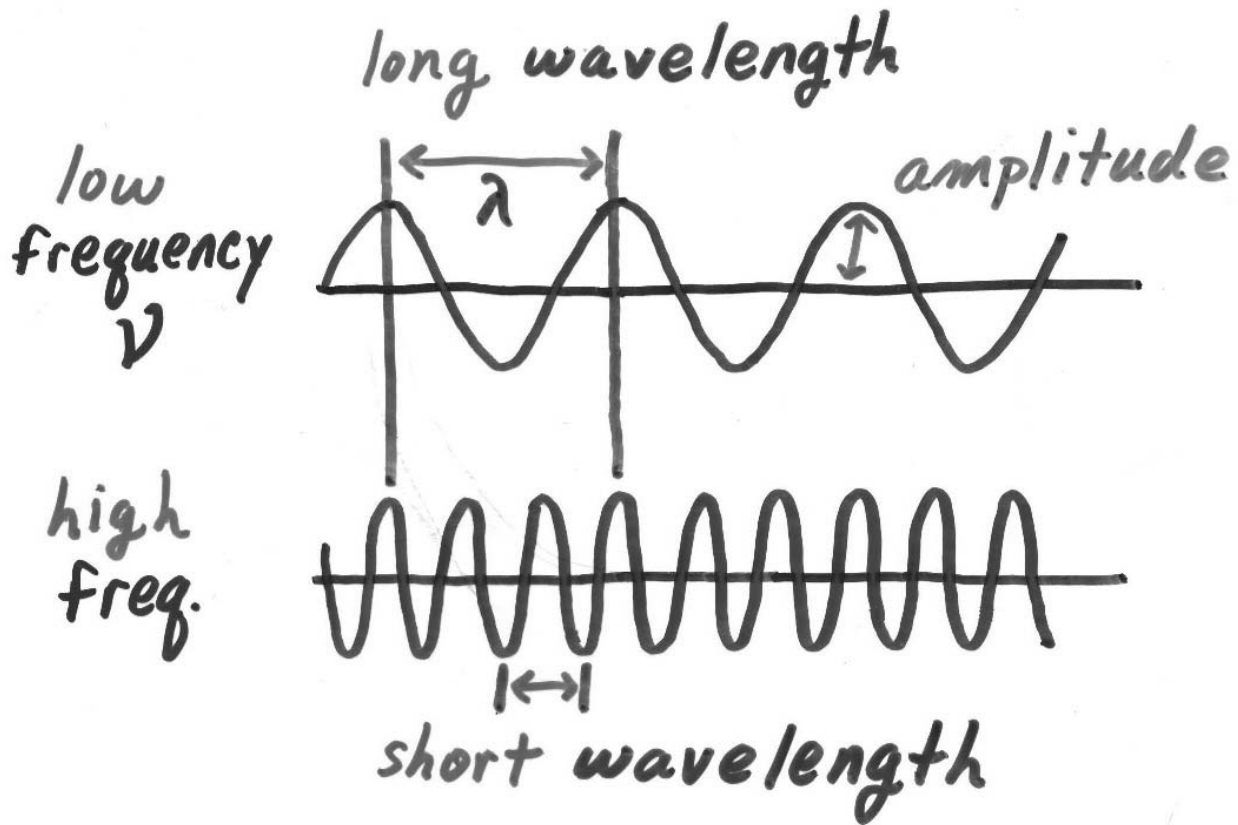
##### Radiant Energy

light, X-rays, UV, microwaves, etc.

All move at the speed of light,

$$c = 2.99792 \times 10^8 \text{ m/s}$$

have wavelike characteristics

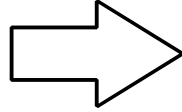


$\lambda$ , wavelength      distance between successive peaks

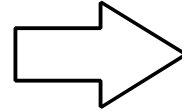
$\nu$ , frequency      number of complete wavelengths or cycles which pass a given point per second

amplitude      height of peak - related

to intensity of radiation



long  
wavelength



low  
frequency

short  
wavelength

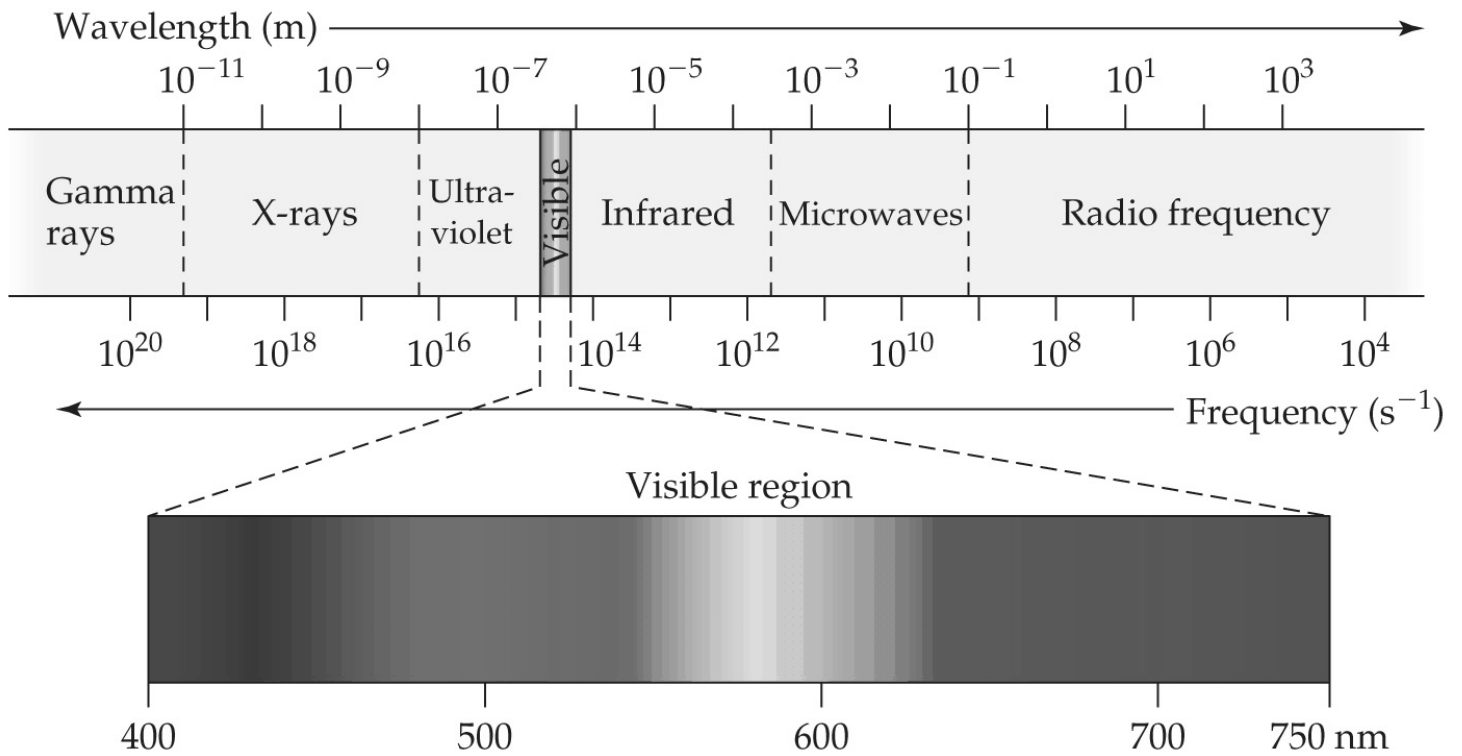
high  
frequency

$$\nu \propto \frac{1}{\lambda}$$

$$\nu = \frac{c}{\lambda} \quad \text{or} \quad \nu \cdot \lambda = c$$

units for  $\nu$

$s^{-1}$  ; cycles/s ; hertz, Hz



© 2012 Pearson Education, Inc.

X- rays    visible    IR    microwave    radio

$\lambda$  (m)    10<sup>-9</sup>    10<sup>-7</sup>    10<sup>-5</sup>    10<sup>-2</sup>    10<sup>2</sup>

$\nu$  (s<sup>-1</sup>)    10<sup>17</sup>    10<sup>15</sup>    10<sup>13</sup>    10<sup>10</sup>    10<sup>6</sup>

## II) Quantized Energy and Photons

### A) Plank's Theory

Energy changes are quantized

- discrete energy changes

$$\Delta E = n h \nu \quad n = 1, 2, 3, 4, \dots$$

Planck's constant

$$h = 6.63 \times 10^{-34} \text{ J}\cdot\text{s},$$

Smallest increment of energy, at a given frequency, is termed a quantum of energy

## B) Photoelectric Effect

A minimum freq. of light shining on a metal surface causes it to emit  $e^-$

Einstein: energy is a stream of particle like energy packets called photons

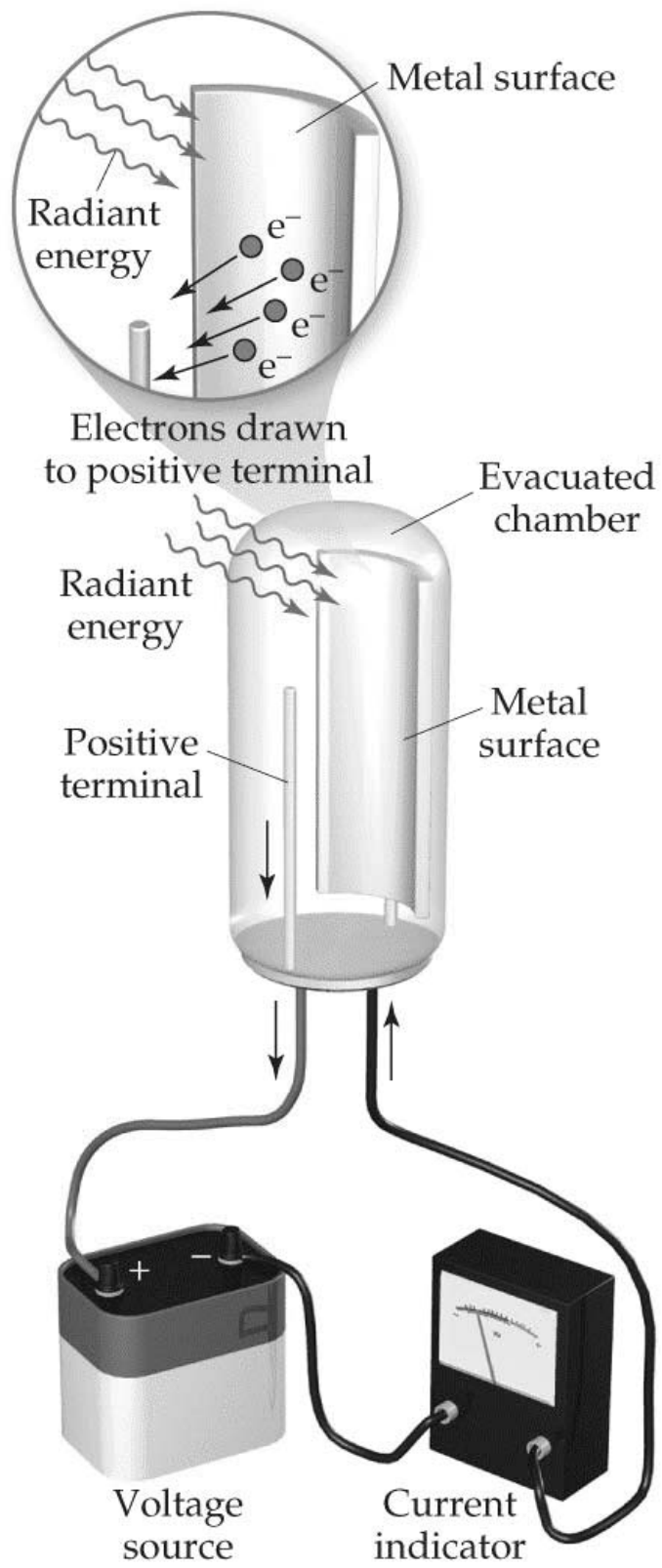
- radiant energy is quantized

$$E_{\text{photon}} = h \nu = \frac{h c}{\lambda}$$

high  $\nu$  (low  $\lambda$ )  $\Rightarrow$  high E

low  $\nu$  (high  $\lambda$ )  $\Rightarrow$  low E

Note : duality of light - behaves both as a wave and particle



© 2012 Pearson Education, Inc.

1) Ex: A laser emits a signal with a wavelength of 351 nm. Calculate the energy of a photon of this radiation.

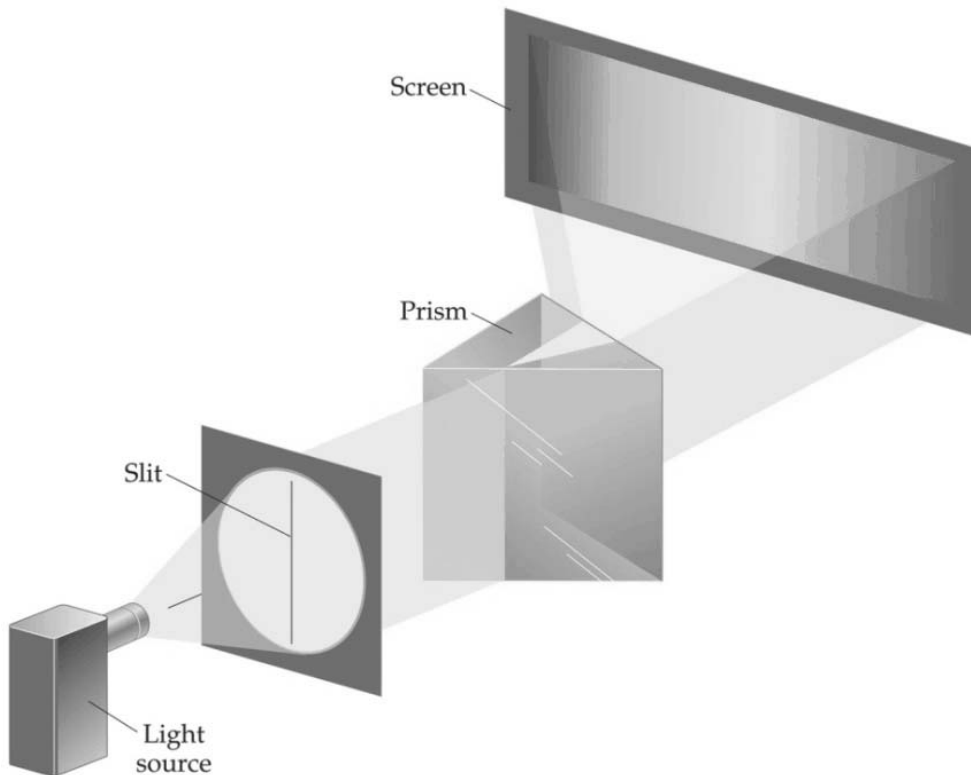
$$\begin{aligned} E &= h \nu = \frac{h c}{\lambda} \\ &= \frac{(6.63 \times 10^{-34} \text{ Js})(3.00 \times 10^8 \text{ m/s})}{3.51 \times 10^{-7} \text{ m}} \\ &= 5.67 \times 10^{-19} \text{ J} \end{aligned}$$



# III) Line Spectra and the Bohr Model

## A) Line Spectra

- 1) White light passing through a prism results in band called a continuous spectrum (rainbow)



Copyright © 2006 Pearson Prentice Hall, Inc.

## 2) monochromatic light

Light with a single wavelength

- lasers

## 3) Line Spectra

discharge tube - atom absorbs energy  
& it can later emit it as light

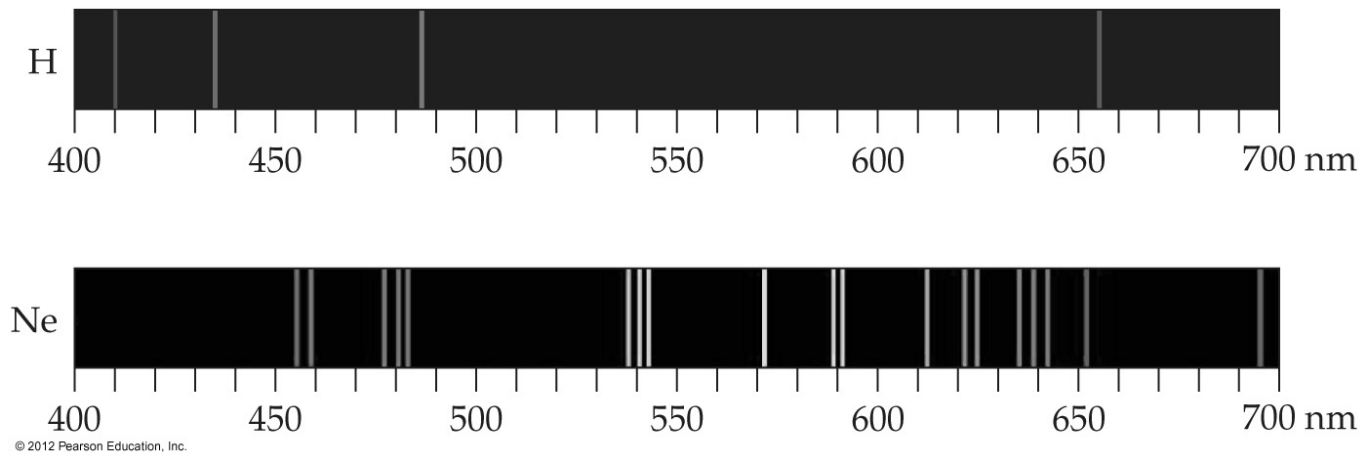
Passed through a prism see a  
series of narrow colored lines  
(specific  $\lambda$ 's)

### Line Spectrum

Each line associated with a  
particular energy and color

# Different elements give different & distinctive line spectra

- characteristic of a particular element
- use to identify elements



## B) Rydberg Equation

Wavelengths of lines in hydrogen spectrum given by,

$$\frac{1}{\lambda} = R_H \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

$$n_2 > n_1$$

### Rydberg Constant

$$R_H = 1.097 \times 10^7 \text{ m}^{-1}$$

## B) Bohr Model

### 1) Energy Levels & Orbits

$e^-$  is restricted to certain energy levels corresponding to spherical orbits, w. certain radii, about the nucleus

$$r = n^2 a_0$$

$$E_n = - h c \cdot R_H \left( \frac{1}{n^2} \right)$$

$n$  = principle quantum number

$$n = 1, 2, 3, \dots, \infty$$

Bohr radius:

$$a_0 = 5.292 \times 10^{-11} \text{ m} = 0.5292 \text{ \AA}$$

$$h c \cdot R_H = 2.180 \times 10^{-18} \text{ J}$$

a) Ground State

$e^-$  in  $n = 1$  orbit  
closest to nucleus

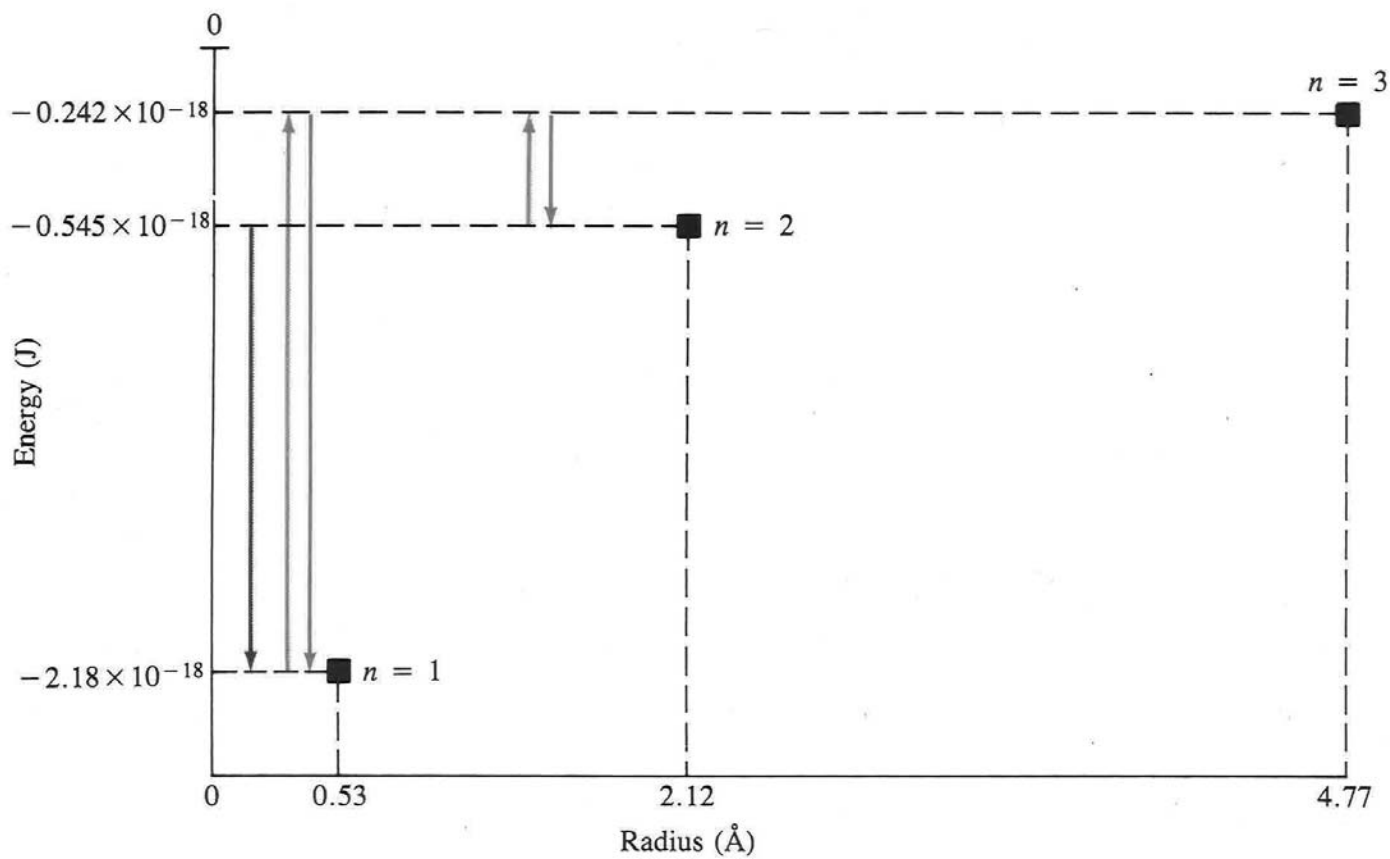
largest value of  $1/n^2$

most negative  $E$

\* Lowest energy level

Note: most neg.  $E$  represents  
most stable state

# Radii and energies of Bohr orbits 1-3



CHEMISTRY: THE CENTRAL SCIENCE  
by Brown/Le May/Bursten

© by Prentice Hall  
A Division of Simon & Schuster  
Englewood Cliffs, New Jersey 07632

## b) Excited States

$$n > 1$$

higher energy

less neg. E, less stable

inc. distance from nucleus

$$r \propto n^2$$

## c) Zero-Point of Energy

$$n = \infty$$

$e^-$  completely separated  
from nucleus

$$E_{\infty} = -h c \cdot R_H \left( \frac{1}{\infty} \right) = 0$$

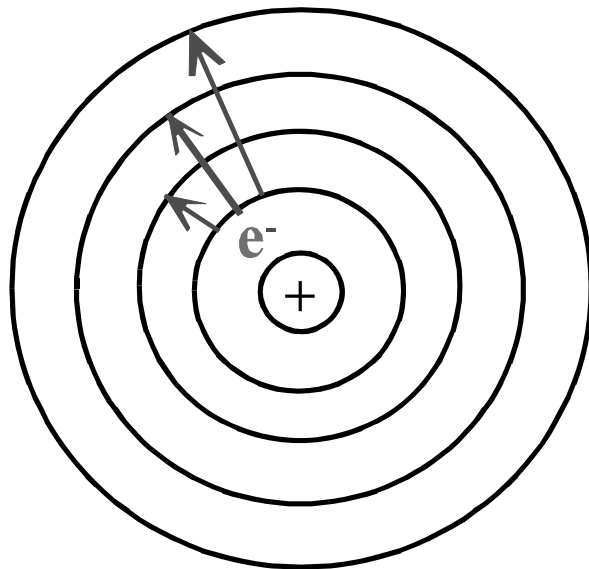


## 2) Energy Transitions

### a) Absorption of Energy

$e^-$  absorbs energy

- jumps to higher energy levels, farther from nucleus
- Excited State

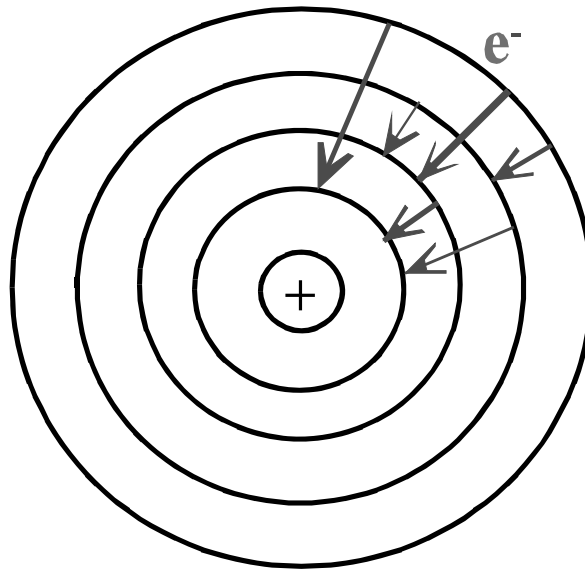


## b) Emission of Energy - light

$e^-$  “falls” to lower level

- emits the energy diff. as  
a quantum of light,

a photon



$$E_{\text{photon}} = - \Delta E_{\text{emission}} = h \nu = \frac{h c}{\lambda}$$

### c) Energy Changes

Energy diff. between orbits

$$\Delta E = E_f - E_i = \frac{-h c \cdot R_H}{n_f^2} - \frac{-h c \cdot R_H}{n_i^2}$$

$$\Delta E = -h c \cdot R_H \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

$$\Delta E = -2.180 \times 10^{-18} \text{ J} \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

1)  $n_f > n_i$

$$\Delta E > 0, \quad E \text{ inc.}$$

Absorption

$$2) \quad \underline{n_f < n_i}$$

$$\Delta E < 0, \quad E \text{ dec.}$$

Emission

$$3) \quad \underline{n_f = \infty}$$

complete removal of  $e^-$

Ionization



$$n_i = 1 \quad n_f = \infty$$

$$\Delta E = h \mathbf{c} \cdot R_H \left( \frac{1}{1^2} \right) = 2.180 \times 10^{-18} \text{ J}$$

## d) Energy of a Photon

Energy of a photon emitted when  $e^-$  “drops” to a lower energy level is related to freq. (wavelength) of radiation

$$E_{\text{photon}} = -\Delta E_{\text{em}} = h \nu = \frac{h c}{\lambda}$$

$$\nu = c \cdot R_H \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

or

$$\frac{1}{\lambda} = R_H \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

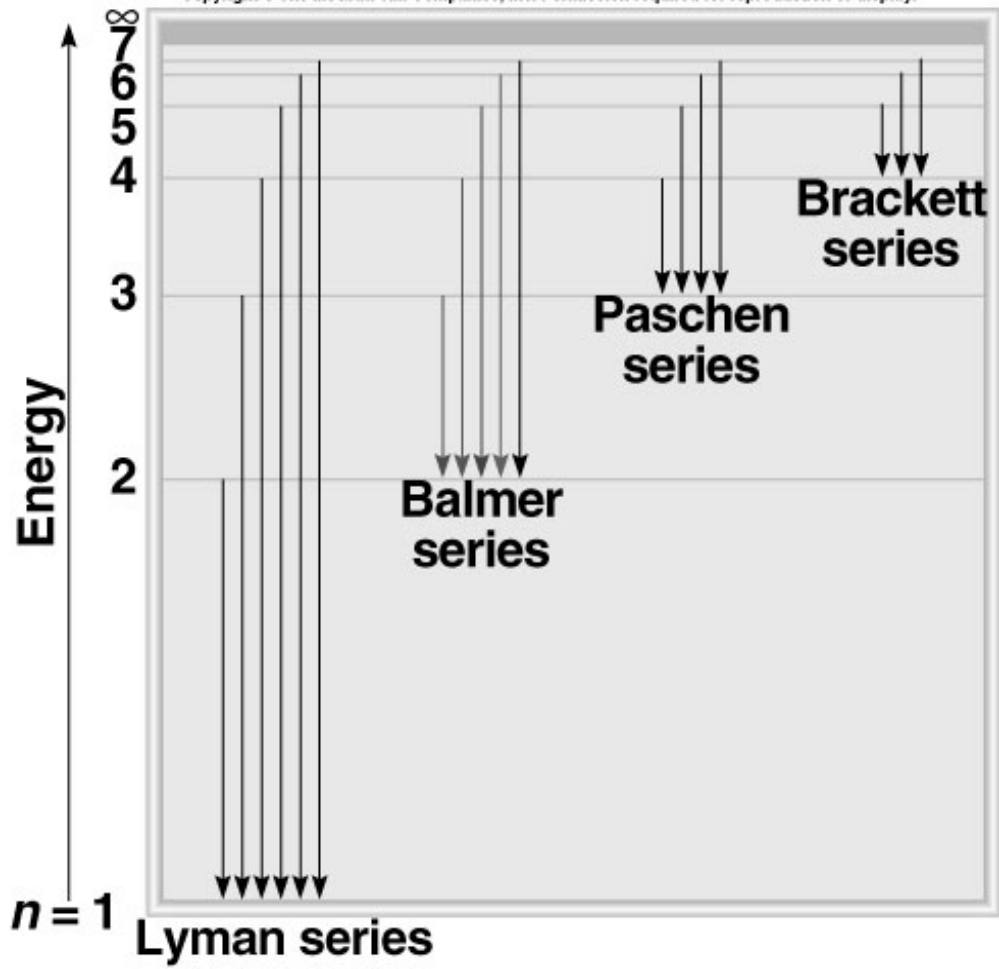
e) Ex : Calc. the wavelength of a line in the visible spectrum for which  $n_i = 3$ .

$$\frac{1}{\lambda} = R_H \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

Balmer Series (visible):

$$n_f = 2$$

Copyright © The McGraw-Hill Companies, Inc. Permission required for reproduction or display.



## IV) Wave Behavior of Matter

### A) de Broglie

Matter should have wave prop.

For photons:

$$E_{\text{photon}} = h \nu = \frac{h c}{\lambda}$$

From Einstein:

$$E = m c^2$$

$$\lambda = \frac{h}{m c}$$

wavelength for photon traveling at  $c$   
with an effective mass,  $m$



## B) de Broglie Wavelength for Particles

$$\lambda = \frac{h}{m v}$$

$v$  = velocity of the particle

$h$  ( $6.63 \times 10^{-34}$  J•s) is extremely small so  $\lambda$  is too small for macroscopic particles.

$\lambda$  can only be detected for particles w. very small mass,

i.e.  $e^-$  ( $m = 9.11 \times 10^{-28}$  g)

1) Ex 1: Calculate the de Broglie wavelength for a 907.2 kg car moving at a speed of 96.6 km/hr.

$$\lambda = \frac{(6.63 \times 10^{-34} \text{ J}\cdot\text{s})}{(907.2 \text{ kg}) (26.83 \text{ m/s})}$$
$$= 2.72 \times 10^{-38} \text{ m}$$

2) Ex 2: Calculate the de Broglie wavelength for an electron moving at a speed of  $3 \times 10^6$  m/s.

$$\lambda = \frac{(6.63 \times 10^{-34} \text{ J}\cdot\text{s})}{(9.11 \times 10^{-31} \text{ kg}) (3 \times 10^6 \text{ m/s})}$$
$$= 2.43 \times 10^{-10} \text{ m} \quad (0.243 \text{ nm})$$

X-rays

## C) Heisenberg Uncertainty Principle

The wave-particle duality of matter makes it impossible to precisely measure both the position and momentum of an object.

$\Delta x$  = uncertainty in position

$\Delta p$  = uncertainty in momentum (mv)

$$\Delta x \cdot \Delta p \geq \frac{h}{4\pi}$$

Limit on simultaneously measuring position and momentum (speed).

## V) Quantum Mechanics

Impose wave properties on  $e^-$

### A) Schrödinger's Wave Equation

Total energy of H-atom is  
sum of K.E. and P.E.

Time-Independent Sch. Eqn.:  
(in one dimension)

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi(x)}{dx^2} + V(x) \psi(x) = E \psi(x)$$

K.E.	P.E.	Total E
------	------	---------

$$\hbar = h/2\pi$$

# 1) Wave Functions

Get a series of solutions to the wave eqn.

wave functions,  $\psi$

Each  $\psi$  corresponds to a specific energy & describes a region about the nucleus, an orbital, in which an  $e^-$  w. that energy may be found

$\psi$  has no direct physical meaning

Can only determine the probability of finding  $e^-$  in a certain region of space at a given instant,

$\psi^2$  probability density

Electron density

Greater where  $e^-$  spends more of its time.

Probability of finding an  $e^-$  is high in regions of high  $e^-$  density

## B) Orbitals & Quantum Numbers

$\psi$  represents an orbital and has 3 characteristic quantum numbers associated with it,

$n$	$\ell$	$m_\ell$
energy and distance from nucleus	shape	orientation of an orbital

The first 3 arise naturally from the solution of the Sch. Eqn.

There is a 4<sup>th</sup> quantum no.

$m_s$  : spin

# 1) Principal quantum number, n

Determines:

- energy level
- average distance from nucleus
- Identifies the shell

$$n = 1, 2, 3, 4, \dots$$

Larger  $n \Rightarrow$  farther shell is from  
nucleus & higher energy

$\text{Max. no. of } e^- \text{ in shell} = 2n^2$
---

$$n = 1 \quad 2(1)^2 = 2 e^-$$

$$n = 2 \quad 2(2)^2 = 8 e^-$$

$$n = 3 \quad 2(3)^2 = 18 e^-$$



## 2) Azimuthal q. n. , $\ell$

(Angular Momentum q.n.)

identifies subshell (energy sublevels)

defines shape of orbital

# subshells in a shell = n

$$\ell = 0, 1, 2, \dots (n-1)$$

Subshell designated by letters:

$\ell$	=	0	1	2	3	4	....
		s	p	d	f	g	
#e <sup>-</sup> in							
subshell		2	6	10	14	18	
$2(2\ell + 1)$							

If  $n = 4$        $\ell = 0, 1, 2, 3$   
                    4s   4p   4d   4f

$n^{\text{th}}$ shell (n)	no. of subshells (= n)	designation	max # $e^-$ by subshell	total # $e^-$ in shell ( $2n^2$ )
---------------------------------	------------------------------	-------------	-------------------------------	---

1	1	1s	2	2
2	2	2s, 2p	2+6	8
3	3	3s, 3p, 3d	2+6+10	18
4	4	4s, 4p, 4d, 4f	2+6+10+14	32

### 3) Magnetic q.n., $m_\ell$

Describes orientation of  
orbital in space

$$m_\ell = +\ell, \dots, 0, \dots, -\ell$$

integer values from  $+\ell$  to  $-\ell$

# possible values = # orbitals in  
for  $m_\ell$  a subshell

$(2\ell + 1)$  orbitals in a subshell

Total # orbitals in shell  $n = n^2$

orbital contains a max. of  $2 e^-$

max. #  $e^-$  in subshell =  $2(2\ell + 1)$

a) Ex:

$$\ell = 0$$

$$m_\ell = 0;$$

s subshell has  
1 orbital

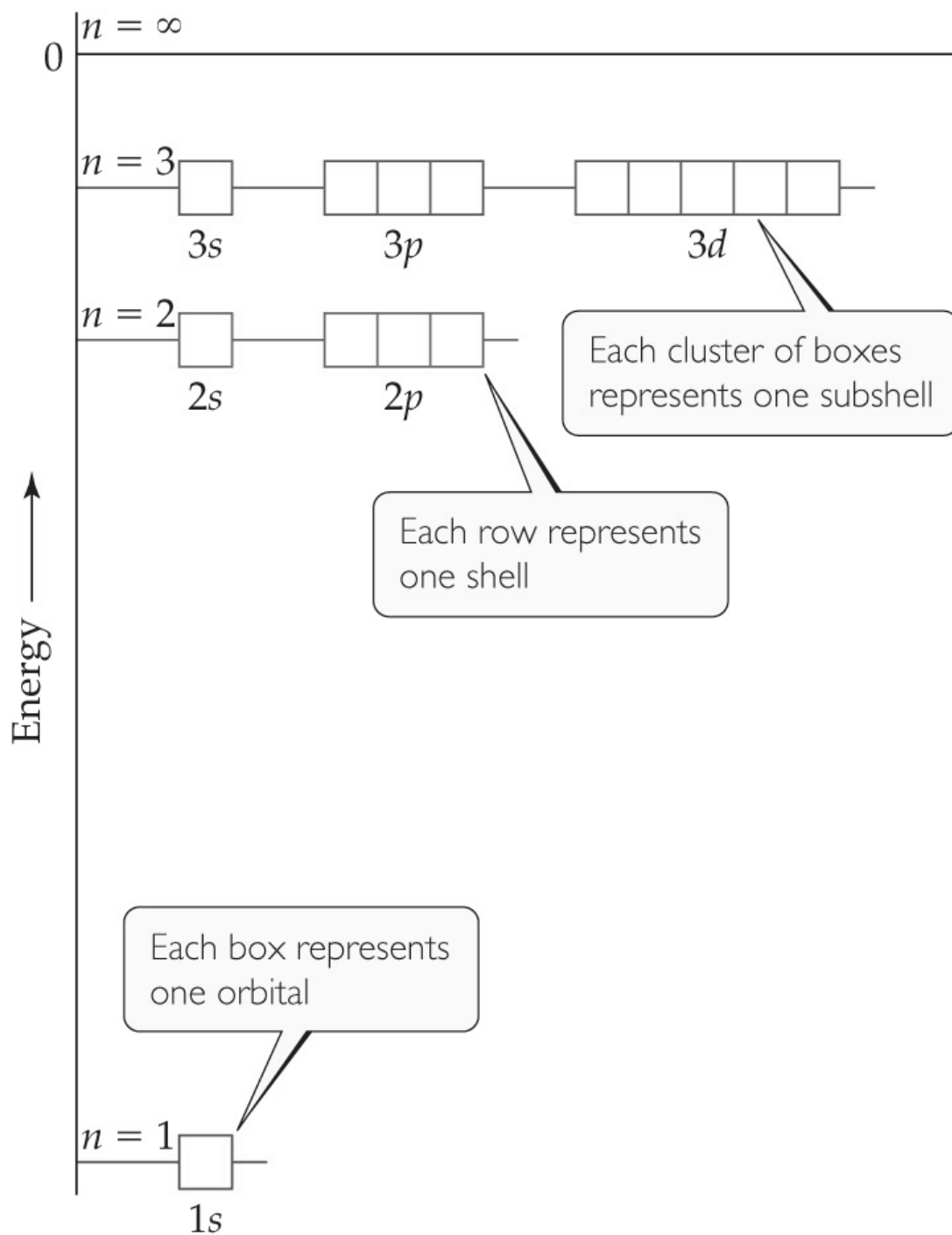
max # e<sup>-</sup>

<u><math>\ell</math></u>	<u>subshell</u>	<u># orbitals</u>	<u>max # e<sup>-</sup> in subshell</u>
0	s		
1	p		
2	d		
3	f		
4	g		
5	h		

**TABLE 6.2 • Relationship among Values of  $n$ ,  $l$ , and  $m_l$  through  $n = 4$**

$n$	Possible Values of $l$	Subshell Designation	Possible Values of $m_l$	Number of Orbitals in Subshell	Total Number of Orbitals in Shell
1	0	1s	0	1	1
2	0	2s	0	1	4
	1	2p	1, 0, -1	3	
3	0	3s	0	1	9
	1	3p	1, 0, -1	3	
	2	3d	2, 1, 0, -1, -2	5	
4	0	4s	0	1	16
	1	4p	1, 0, -1	3	
	2	4d	2, 1, 0, -1, -2	5	
	3	4f	3, 2, 1, 0, -1, -2, -3	7	

# Energy Levels in the H atom



$n = 1$  shell has one orbital

$n = 2$  shell has two subshells composed of four orbitals

$n = 3$  shell has three subshells composed of nine orbitals

© 2012 Pearson Education, Inc.

## VI) Representations of Orbitals

$\psi$  has no direct physical meaning

$\psi^2$  probability density  
(electron density)

probability of finding  $e^-$  at  
a given point in space

$(4\pi r^2) \psi^2$  radial probability density

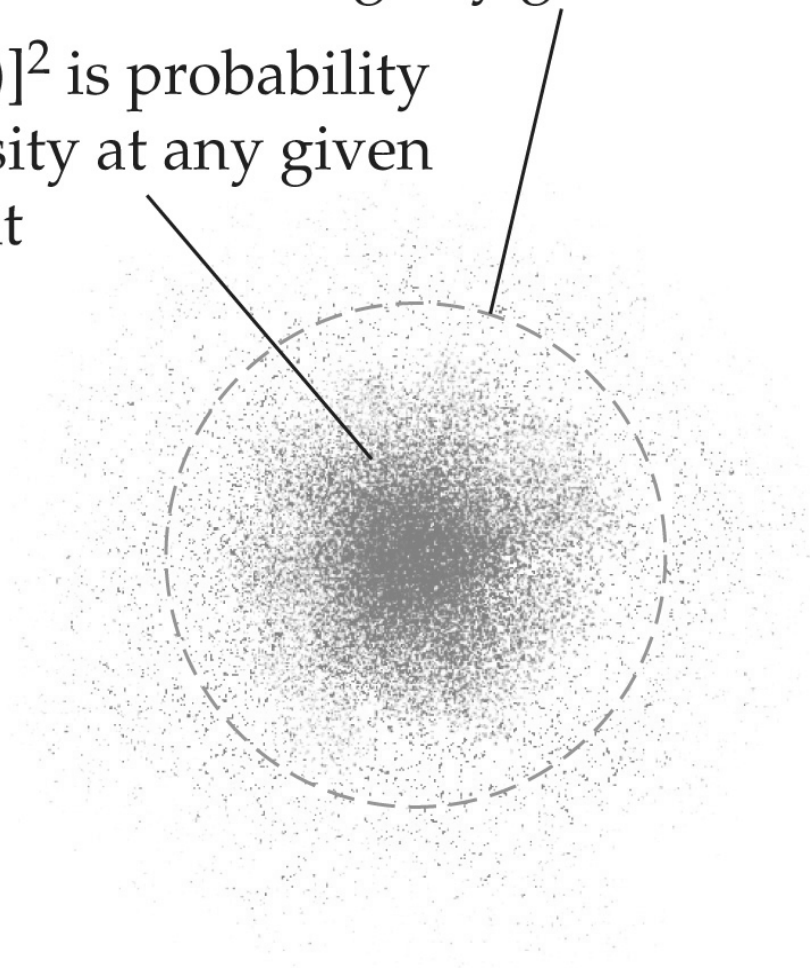
probability of finding  $e^-$  at a  
specific distance,  $r$ , from the  
nucleus

# A) s orbitals

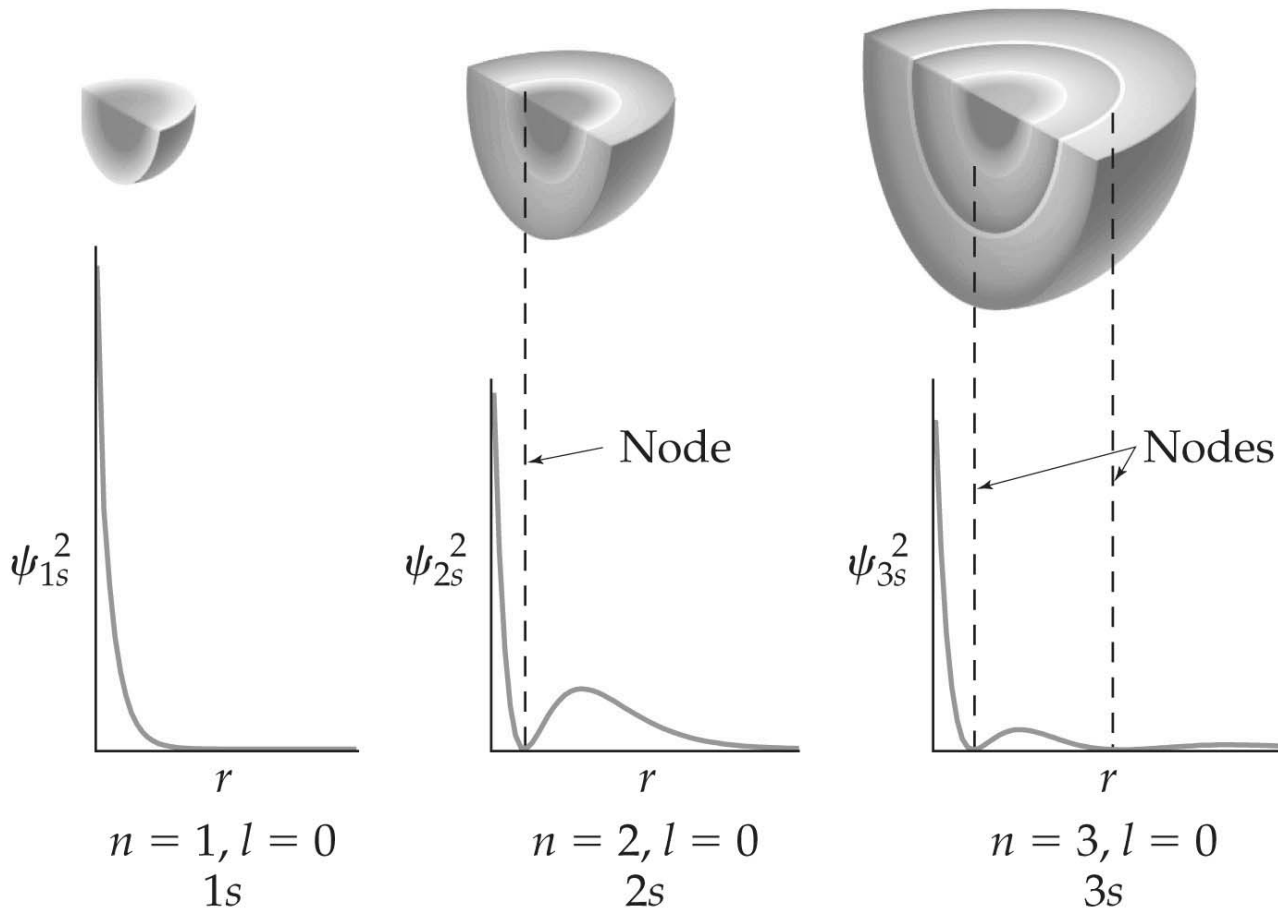
$l = 0$  All s orb. are spherical

$4\pi r^2[\psi(r)]^2$  is radial probability function = sum of all  $[\psi(r)]^2$  having any given value of  $r$

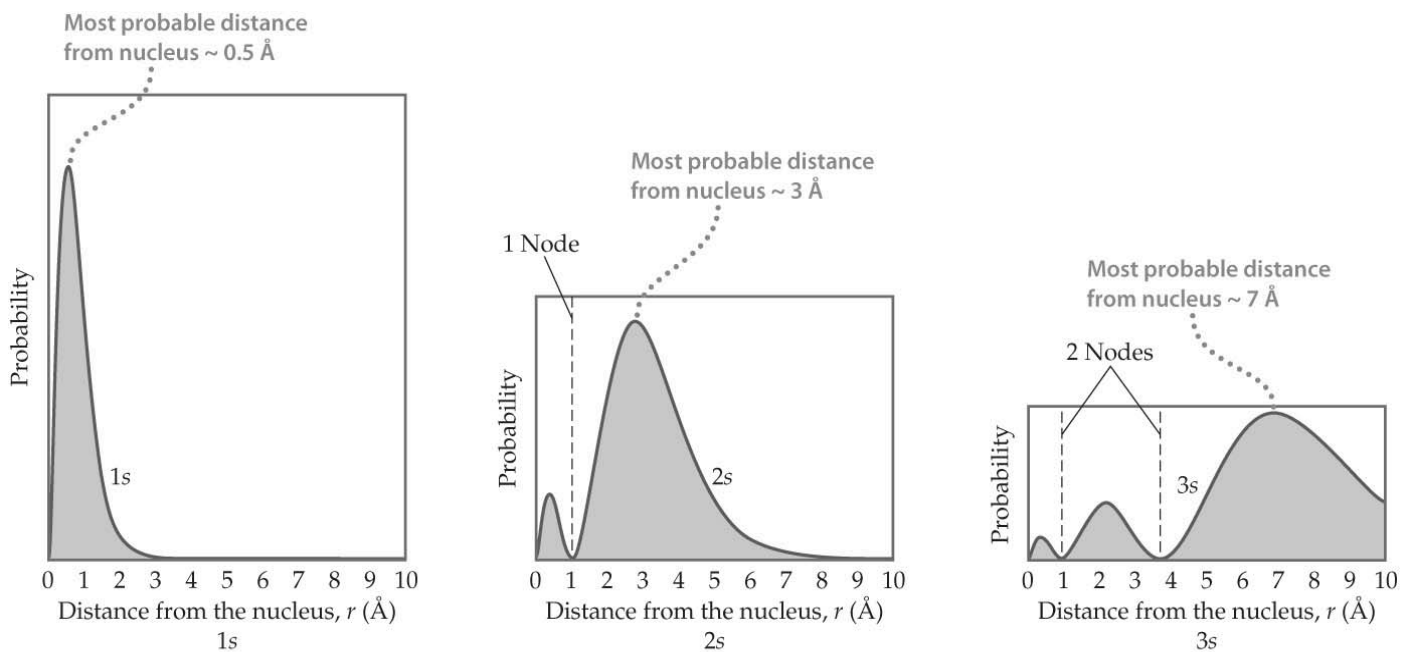
$[\psi(r)]^2$  is probability density at any given point







© 2012 Pearson Education, Inc.



© 2012 Pearson Education, Inc.

1) 3 trends from radial prop. dist.

a) Number of peaks inc. w. inc. n

$$\# \text{ peaks} = n$$

most probable distance further out & peaks get larger as move further from nucleus

b) Number of nodes inc. w. inc. n

points where the prob. is zero

$$\# \text{ nodes} = n - 1$$

$$\# \text{ spherical nodes} = n - \ell - 1$$

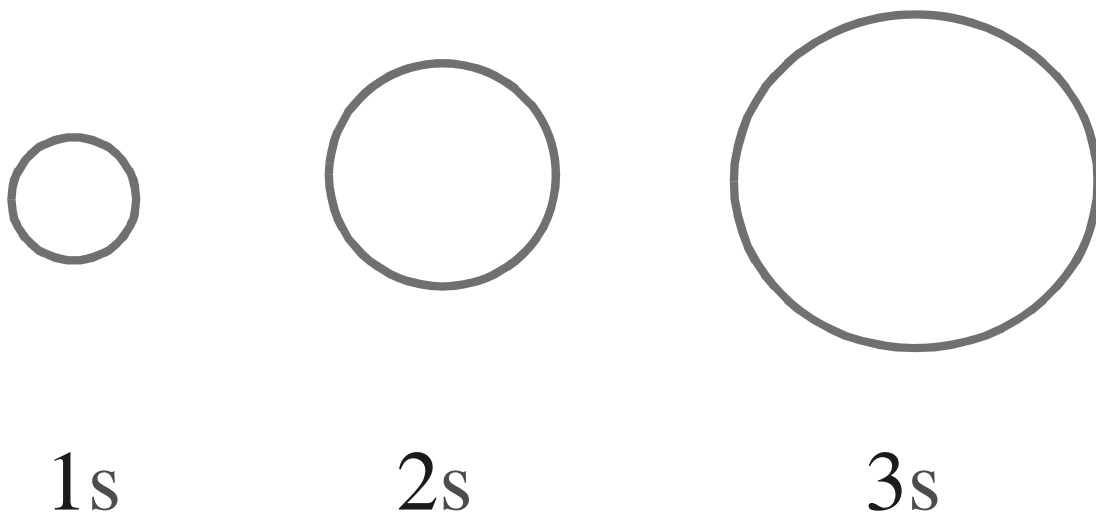
$$\# \text{ angular nodes} = \ell$$

c)  $e^-$  density spreads out w. inc. n

## 2) Contour Representation

represent a volume of space in which there is a high probability of finding the  $e^-$

usually 90%



$e^-$  in orb. of higher  $n$  will be greater avg. distance from nucleus

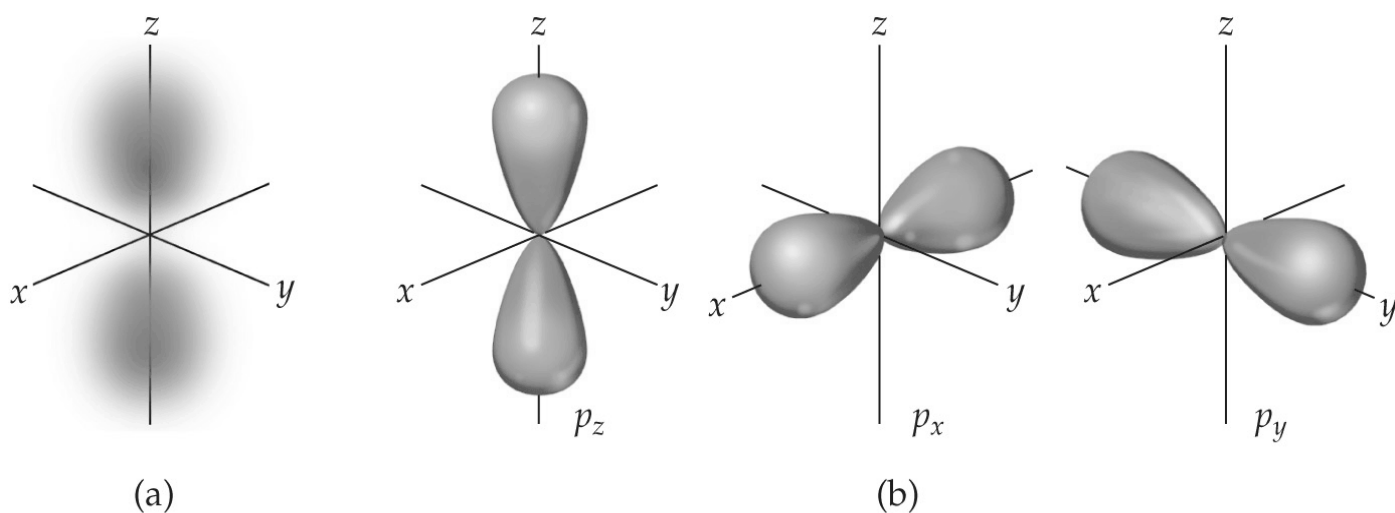
## B) p orbitals

All p orbitals have 2 lobes pointing  
in opposite directions

dumbbell or teardrop

The 3 p orbs in a subshell differ in  
their orientation in space

- at right angles to each other



© 2012 Pearson Education, Inc.

## VII) Many-Electron Atoms

H atom has only 1  $e^-$

$E_{\text{orb}}$  depends on  $n$  and is determined by attraction between positive proton and negative  $e^-$  and average distance between them

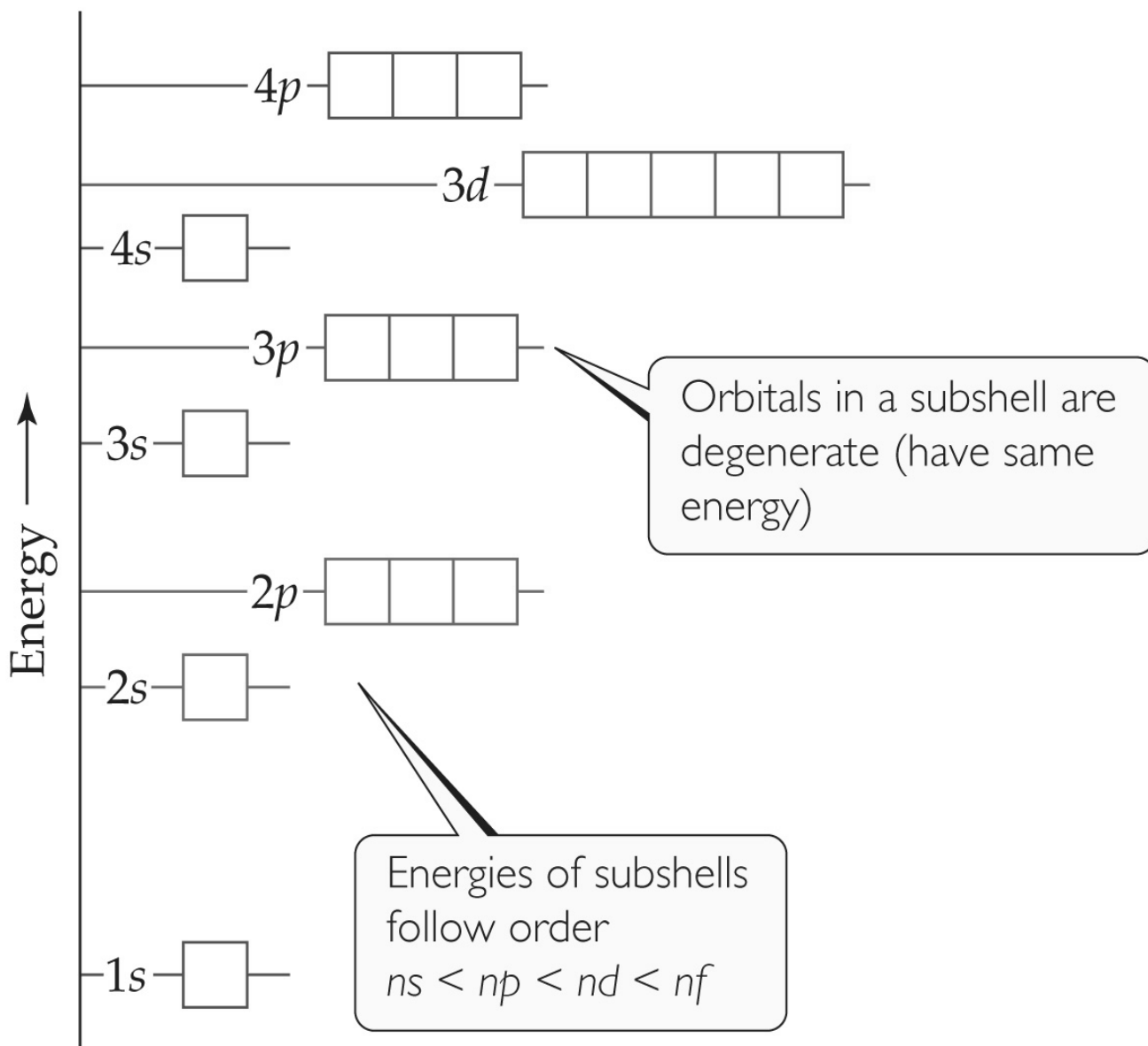
Many- $e^-$  atoms:

Add  $e^- - e^-$  repulsions to  $E$  & diff.  $e^-$ -nucleus attractions

Causes subshells to have diff.  $E$

$E_{\text{orb}}$  now depends on  $n$  and  $\ell$

$E$  of orbitals w/in subshell still degenerate



© 2012 Pearson Education, Inc.

## A) Electron Spin

$e^-$  “spins” about its own axis

- spinning charge generates a magnetic field

$e^-$  only spin in either of 2 directions

quantized

electron spin q.n.,  $m_s$

+1/2

up

↑

-1/2

down

↓

## B) Pauli Exclusion Principle

No 2  $e^-$  in an atom can have same set of 4 quantum no.'s

$$n, \ell, m_\ell, m_s$$

Look at 1s orbital

$$n = 1, \ell = 0, m_\ell = 0$$

can have only 2  $e^-$  w. diff. values of  $m_s$ ,  $+1/2$  or  $-1/2$

Limits max . #  $e^-$  in orbital to 2

- MUST have opposite spins



## C) Summary of Quantum Numbers

### 1) Shell number, n

$$n = 1, 2, 3, 4, \dots$$

energy level & avg. distance  
Period no.  $\Rightarrow$  highest n

$$\text{Max \# } e^- \text{ in shell} = 2n^2$$

### 2) Subshell, $\ell$ (shape of orbital)

$$\text{\# subshells in shell} = n$$

$$\ell = 0, 1, 2, \dots (n-1)$$

s, p, d, f, g, h....

$$\text{\# } e^- \text{ in subshell} = 2(2\ell + 1)$$

### 3) Orbitals, $m_\ell$ (orientation)

$$m_\ell = +\ell, \dots, 0, \dots, -\ell$$

$\# \text{ orb. in shell} = n^2$
----------------------------------

$(2\ell + 1)$  orbitals in a subshell

$$\text{max. } \# e^- \text{ in subshell} = 2 (2\ell + 1)$$

### 4) Spin, $m_s$

$$+1/2 (\uparrow) \quad -1/2 (\downarrow)$$

# Subshell letters, # orbitals & max # e<sup>-</sup> in subshell

	$l =$	0	1	2	3	4	5
subshell letters		s	p	d	f	g	h ...
# orbitals in subshell		1	3	5	7	9	11 ...
max # e <sup>-</sup> in subshell		2	6	10	14	18	22 ...

## VIII) Electron Configurations

Orbitals filled in order of inc. energy until all  $e^-$  have been used

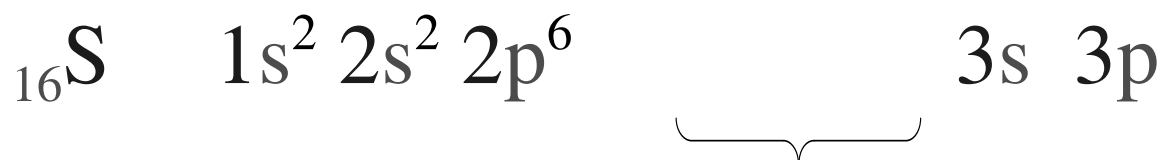
$x \leftarrow \# e^- \text{ in subshell}$

$n \ell$

shell  $\nearrow$   $\nwarrow$  subshell



A) Ex: Consider sulfur,  ${}_{16}\text{S} : 16 e^{-}$



valence shell  
(outer shell)

${}_{16}\text{S}$  is in 3<sup>rd</sup> period ;  $n_{\text{max}} = 3$

${}_{16}\text{S}$  is in group VI A,  $6 e^{-}$  in outer  
or valence shell

valence  $e^{-} \Rightarrow e^{-}$  in outer or  
valence shell

core  $e^{-} \Rightarrow e^{-}$  in inner shells

Note: For representative elements

Period no.  $\Rightarrow$  n value of valence shell

Group no.  $\Rightarrow$  # of valence  $e^-$

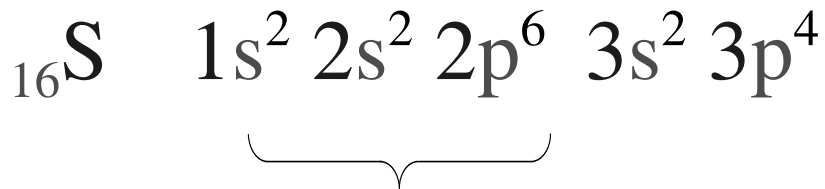
Elements in a group have similar chemical and physical properties

- same valence shell  $e^-$  configuration

$e^-$  in outer shell are ones involved in chemical reactions

## B) Shorthand Electron Configuration

Focus attention on valence shell  $e^-$



completed subshells  $\Rightarrow$  [Ne]

noble gas from  
previous period



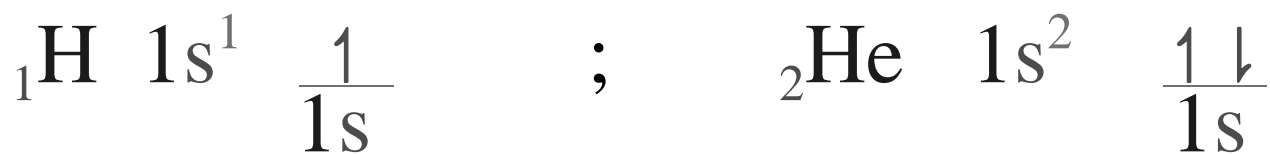
1) Ex:  ${}_6\text{C}$



## C) Orbital Diagrams

A dash  $\text{---}$  indicates an orbital

Use arrows,  $\uparrow$  or  $\downarrow$  to indicate  $e^-$   
with up or down spin



1s 2s 2p





single  $e^-$  in an orbital, 1, unpaired

paramagnetic substance

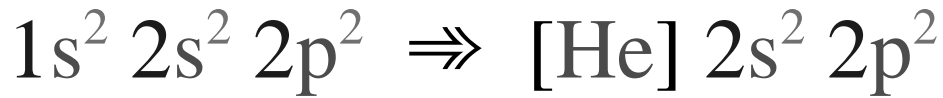
- unpaired  $e^-$ 's
- attracted by magnetic field

2  $e^-$  in same orbital, 1↓, paired

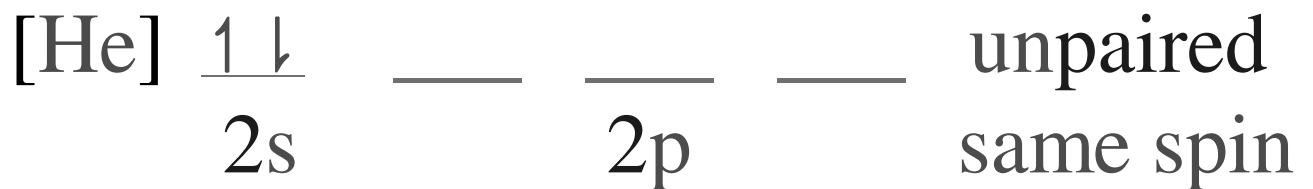
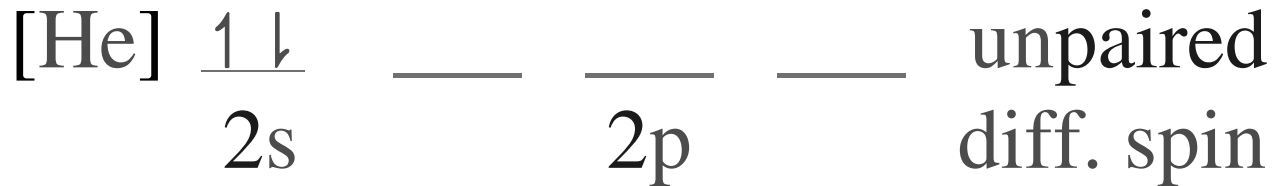
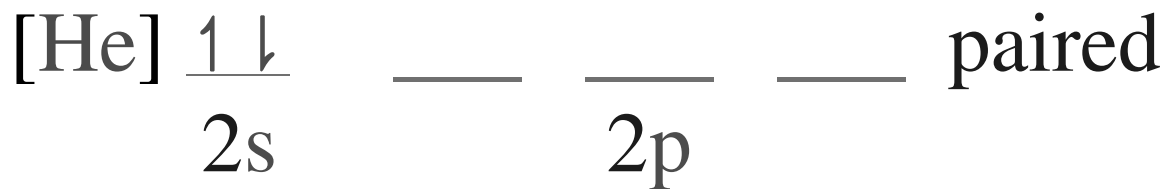
Diamagnetic substance

- all  $e^-$  paired
- not attracted by magnetic field

## D) Hund's Rule



3 possible orbital diagrams:



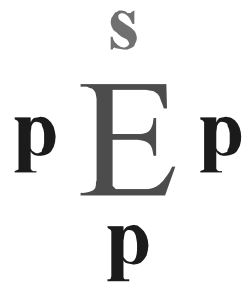
Hunds Rule:  $e^-$  occupy diff. orbitals of a subshell until all are singly occupied before  $e^-$  pairing occurs.

## E) Electron-Dot Symbols

Represent  $e^-$  in the s & p orb. of the valence shell as dots arranged around the symbol of the element.

There are 4 s & p orb. & 4 positions about the symbol

- treat like orb. diagrams



Note: only real useful for representative elements

A) Ex's: Draw e<sup>-</sup> dot symbols

1)  ${}_6\text{C}$



2)  ${}_{12}\text{Mg}$



3)  ${}_{16}\text{S}$



# IX) Electron Conf & Periodic Table

Look at  ${}_{32}\text{Ge}$



n = 4	5s	___			
	4p	___	___	___	
	3d	___	___	___	___
	4s	___			
n = 3	3p	___	___	___	
	3s	___			
n = 2	2p	___	___	___	
	2s	___			
n = 1	1s	___			

# What was happening?

left, filling s orb

$2e^-$ , 2 columns

right, filling p orb.

$6e^-$ , 6 columns

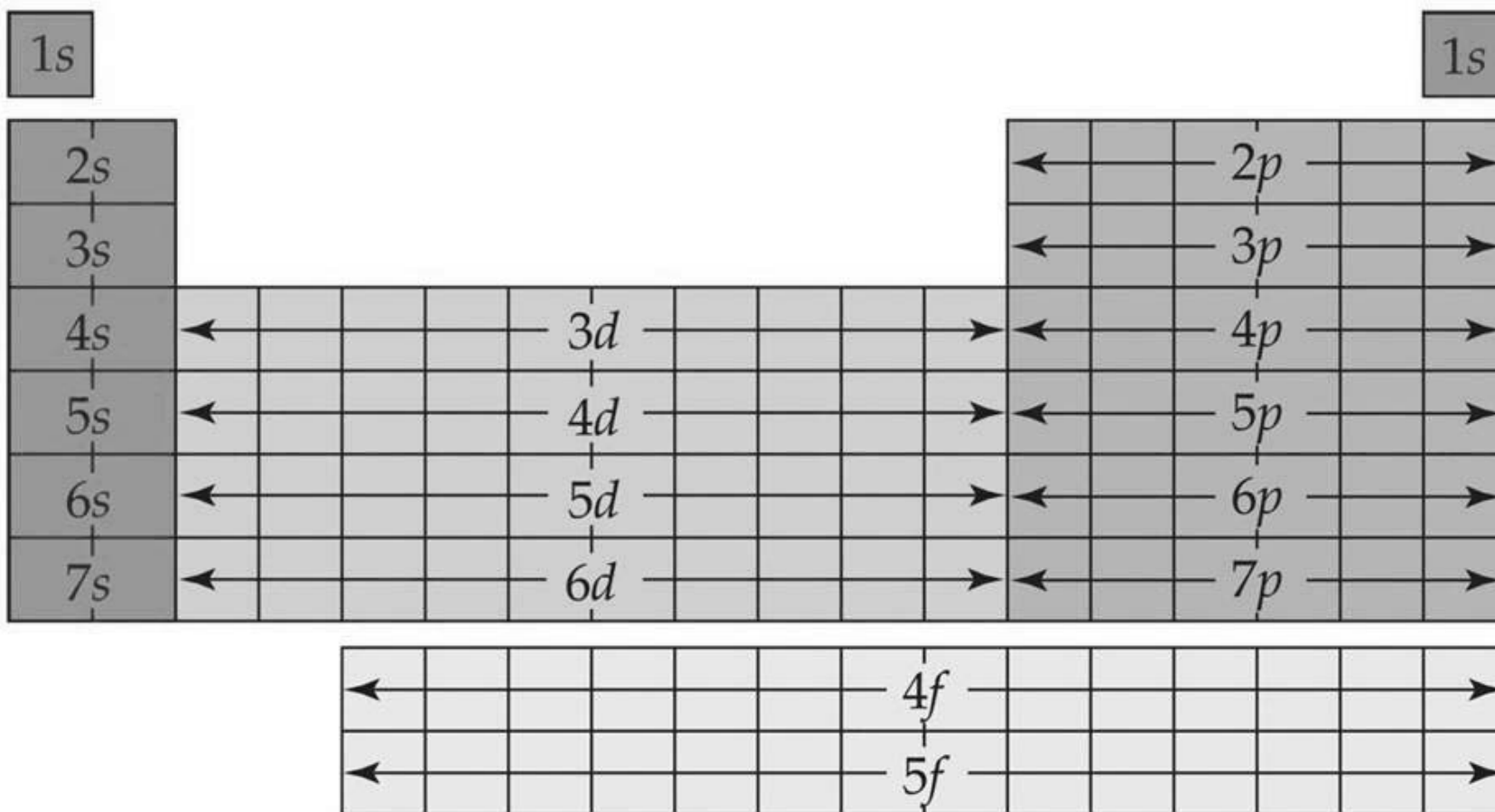
center, filling d orb

$10e^-$ , 10 columns

Period no.  $\Rightarrow$  n value of s & p

subshells of valence shell

Group no.  $\Rightarrow$  # of valence  $e^-$



Representative s-block elements

Representative p-block elements

Transition metals

f-Block metals

Copyright © 2006 Pearson Prentice Hall, Inc.

A) Ex's:



Period no.

3

Group no.

VI A





	IA	IIA	IIIB	IVB	VB	VIB	VII B	VIII B					IB	IIB	IIIA	IVA	VA	VIA	VIIA	VIIIA
1	1.008 <b>H</b> 1																		4.003 <b>He</b> 2	
2	6.941 <b>Li</b> 3	9.012 <b>Be</b> 4												10.81 <b>B</b> 5	12.011 <b>C</b> 6	14.007 <b>N</b> 7	15.999 <b>O</b> 8	18.998 <b>F</b> 9	20.179 <b>Ne</b> 10	
3	22.990 <b>Na</b> 11	24.305 <b>Mg</b> 12												26.98 <b>Al</b> 13	28.09 <b>Si</b> 14	30.974 <b>P</b> 15	32.06 <b>S</b> 16	35.453 <b>Cl</b> 17	39.948 <b>Ar</b> 18	
4	39.098 <b>K</b> 19	40.08 <b>Ca</b> 20	44.96 <b>Sc</b> 21	47.88 <b>Ti</b> 22	50.94 <b>V</b> 23	52.00 <b>Cr</b> 24	54.94 <b>Mn</b> 25	55.85 <b>Fe</b> 26	58.93 <b>Co</b> 27	58.69 <b>Ni</b> 28	63.546 <b>Cu</b> 29	65.38 <b>Zn</b> 30	69.72 <b>Ga</b> 31	72.59 <b>Ge</b> 32	74.92 <b>As</b> 33	78.96 <b>Se</b> 34	79.904 <b>Br</b> 35	83.80 <b>Kr</b> 36		
5	85.47 <b>Rb</b> 37	87.62 <b>Sr</b> 38	88.91 <b>Y</b> 39	81.22 <b>Zr</b> 40	92.91 <b>Nb</b> 41	95.94 <b>Mo</b> 42	98 <b>Tc</b> 43	101.07 <b>Ru</b> 44	102.91 <b>Rh</b> 45	106.42 <b>Pd</b> 46	107.87 <b>Ag</b> 47	112.41 <b>Cd</b> 48	114.82 <b>In</b> 49	118.69 <b>Sn</b> 50	121.75 <b>Sb</b> 51	127.60 <b>Te</b> 52	126.90 <b>I</b> 53	131.39 <b>Xe</b> 54		
6	132.91 <b>Cs</b> 55	137.33 <b>Ba</b> 56	138.91 <b>La</b> 57	178.39 <b>Hf</b> 72	180.95 <b>Ta</b> 73	183.85 <b>W</b> 74	186.21 <b>Re</b> 75	190.23 <b>Os</b> 76	192.22 <b>Ir</b> 77	195.08 <b>Pt</b> 78	196.97 <b>Au</b> 79	200.59 <b>Hg</b> 80	204.38 <b>Tl</b> 81	207.2 <b>Pb</b> 82	208.98 <b>Bi</b> 83	209 <b>Po</b> 84	210 <b>At</b> 85	222 <b>Rn</b> 86		
7	223 <b>Fr</b> 87	226.03 <b>Ra</b> 88	227.03 <b>Ac</b> 89	261 <b>Rf</b> 104	262 <b>Ha</b> 105	263 <b>Sg</b> 106	262 <b>Ns</b> 107	265 <b>Hs</b> 108	266 <b>Mt</b> 109	269 <b>110</b>	272 <b>111</b>	277 <b>112</b>								

6	Lanthanide Series	140.12 <b>Ce</b> 58	140.91 <b>Pr</b> 59	144.24 <b>Nd</b> 60	145 <b>Pm</b> 61	150.36 <b>Sm</b> 62	151.96 <b>Eu</b> 63	157.25 <b>Gd</b> 64	158.93 <b>Tb</b> 65	162.50 <b>Dy</b> 66	164.93 <b>Ho</b> 67	167.26 <b>Er</b> 68	168.93 <b>Tm</b> 69	173.04 <b>Yb</b> 70	173.04 <b>Lu</b> 71
7	Actinide Series	232.04 <b>Th</b> 90	231.04 <b>Pa</b> 91	238.03 <b>U</b> 92	237.05 <b>Np</b> 93	<b>Pu</b> 94	<b>Am</b> 95	<b>Cm</b> 96	<b>Bk</b> 97	<b>Cf</b> 98	<b>Es</b> 99	<b>Fm</b> 100	<b>Md</b> 101	<b>No</b> 102	<b>Lr</b> 103

A PERIODIC CHART OF THE ELEMENTS  
(Based on <sup>12</sup>C)

3)  ${}_{43}\text{Tc}$

4)  ${}_{82}\text{Pb}$

## B) Exceptions



Reason: 4s and 3d are very close in energy. (Can act like degenerate orb)

$\frac{1}{2}$  filled & filled subshells  
are more stable.